

Interferometric Signatures of Hydrodynamics in Small Systems

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We've seen collectivity in small systems.

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Possible explanations include:

- ▶ Initial-state correlations (e.g., CGC)
- ▶ String hadronization models (e.g., Lund string)
- ▶ Escape mechanisms
- ▶ *Hydrodynamics*

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- ▶ *Hydrodynamics*

How can we discriminate between models?

Various observables:

- ▶ Anisotropic flow, strangeness enhancement, etc.
- ▶ **This talk:** Hanbury Brown–Twiss (HBT) interferometry
 - Collective signatures include K_T -scaling of radii, etc.
 - **Today:** *discuss scaling of radii with $dN_{\text{ch}}/d\eta$*

HBT from hydrodynamics

$$C(\vec{p}_1, \vec{p}_2) \equiv E_{p_1} E_{p_2} \frac{d^6 N}{d^3 p_1 d^3 p_2} / \left(E_{p_1} \frac{d^3 N}{d^3 p_1} E_{p_2} \frac{d^3 N}{d^3 p_2} \right)$$

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$$\rightarrow C_{\text{fit}}(\vec{q}, \vec{K}) \equiv 1 + \lambda \exp \left(- \sum_{i,j=o,s,l} R_{ij}^2(\vec{K}) q_i q_j \right)$$
$$\vec{q} \equiv \vec{p}_1 - \vec{p}_2, \vec{K} \equiv \frac{1}{2} (\vec{p}_1 + \vec{p}_2)$$

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Hydrodynamic $S(x, p)$ defined by Cooper-Frye formula:

$$S(x, p) = \frac{1}{(2\pi)^3} \int_{\Sigma(x_f)} \frac{p \cdot d^3 \sigma(x_f) \delta^4(x - x_f)}{e^{(p \cdot u(x_f) - \mu)/T} \pm 1}$$

HBT from hydrodynamics

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For Gaussian sources:

$$\implies R_{ij}^2(\vec{K}) \equiv \langle (\tilde{x}_i - \beta_i \tilde{t})(\tilde{x}_j - \beta_j \tilde{t}) \rangle,$$

$$\langle f(x) \rangle \equiv \frac{\int d^4 x f(x) S(x, K)}{\int d^4 x S(x, K)}$$

$$\tilde{x}_i \equiv x_i - \langle x_i \rangle, \tilde{t} \equiv t - \langle t \rangle, \vec{\beta} \equiv \vec{K} / K^0$$

In particular, we have these “pocket relations”:

$$R_s^2 = \langle \tilde{x}_s^2 \rangle$$

$$R_o^2 = \langle \tilde{x}_o^2 \rangle - \beta_T \langle \tilde{x}_o \tilde{t} \rangle + \beta_T^2 \langle \tilde{t}^2 \rangle$$

$$R_l^2 = \langle \tilde{x}_l^2 \rangle - \beta_L \langle \tilde{x}_l \tilde{t} \rangle + \beta_L^2 \langle \tilde{t}^2 \rangle$$

- ▶ For $\beta_L = 0$, R_s^2 and R_l^2 dominated by *spatial* geometry
 - ▶ For $\beta_T \neq 0$, R_o^2 contains both spatial *and* temporal information
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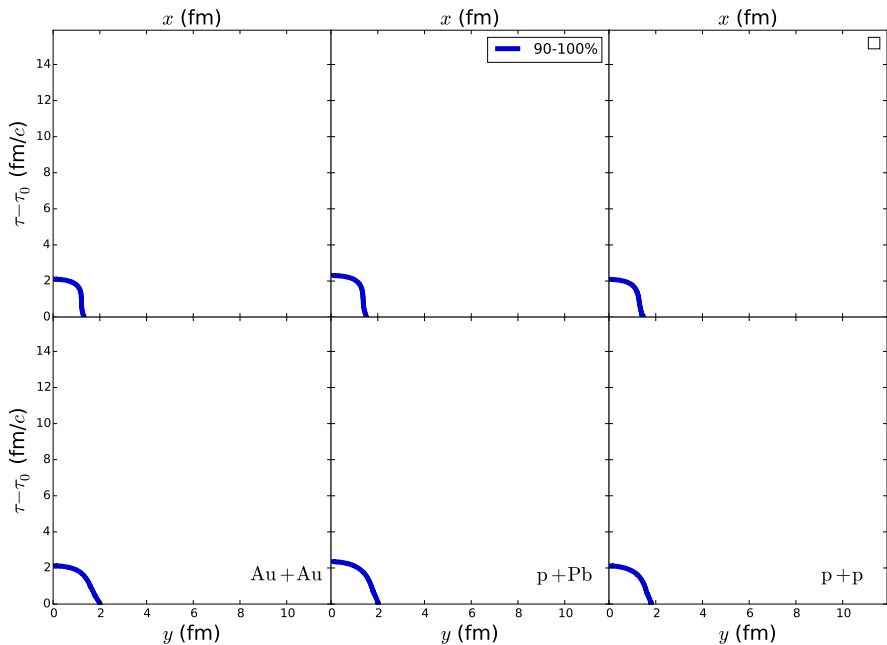
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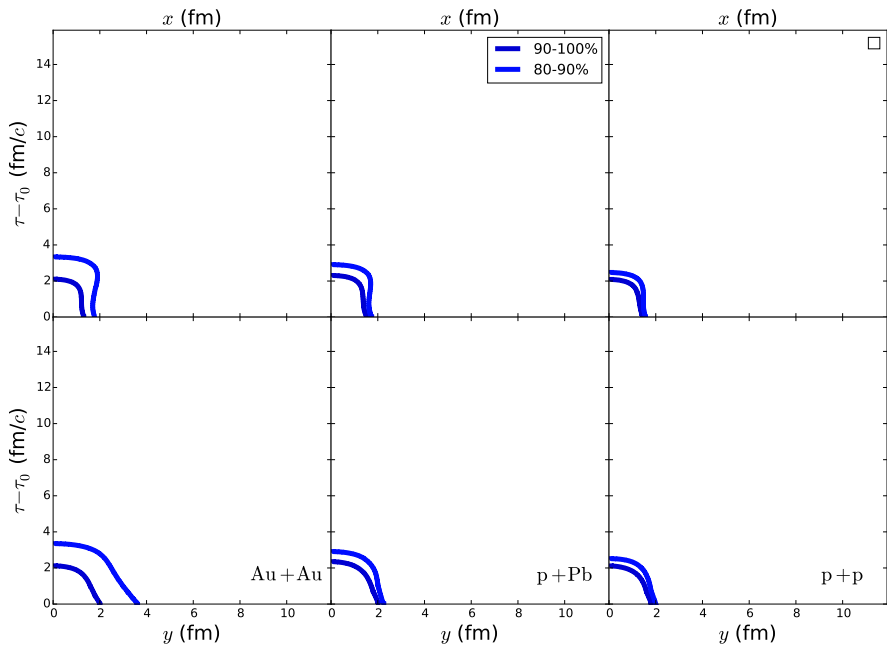
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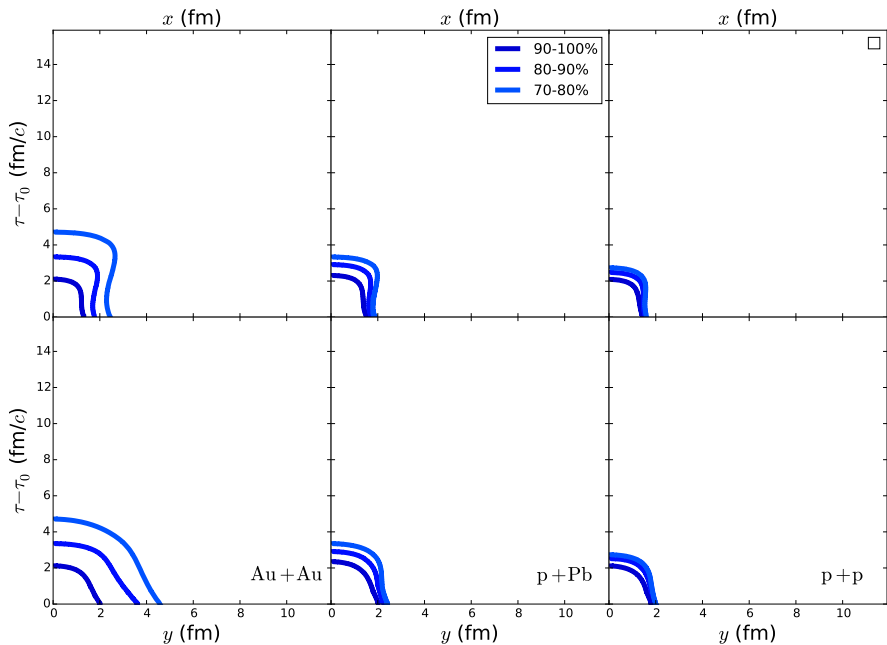
In this talk:

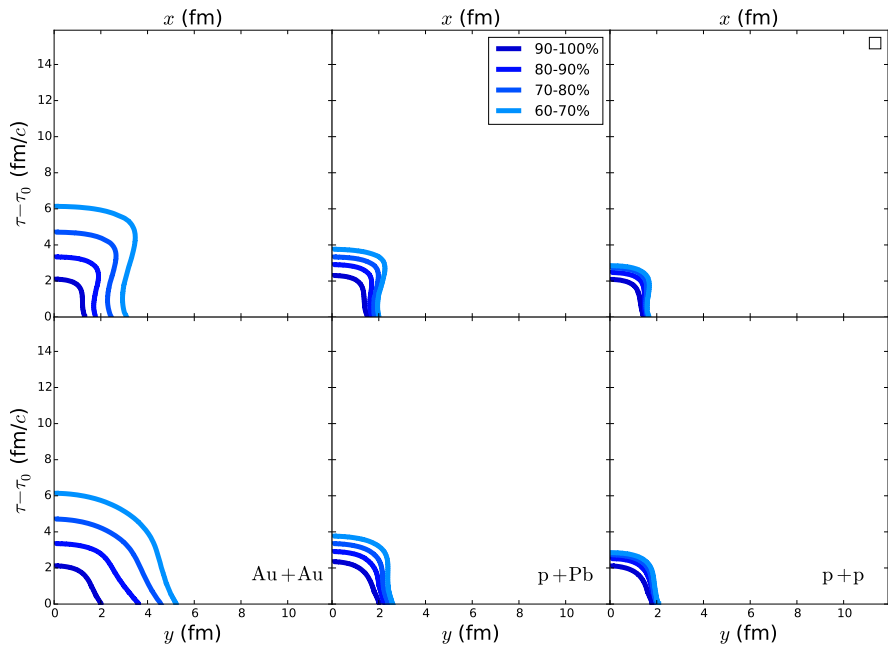
- ▶ **Run** hydrodynamics for p+p (7 TeV), p+Pb (5.02 TeV), and Au+Au (200 GeV)
- ▶ **Compute** the HBT radii from the correlation function $C(\vec{q}, \vec{K})$
- ▶ **Compare** with data
- ▶ **Interpret** the results using pocket relations

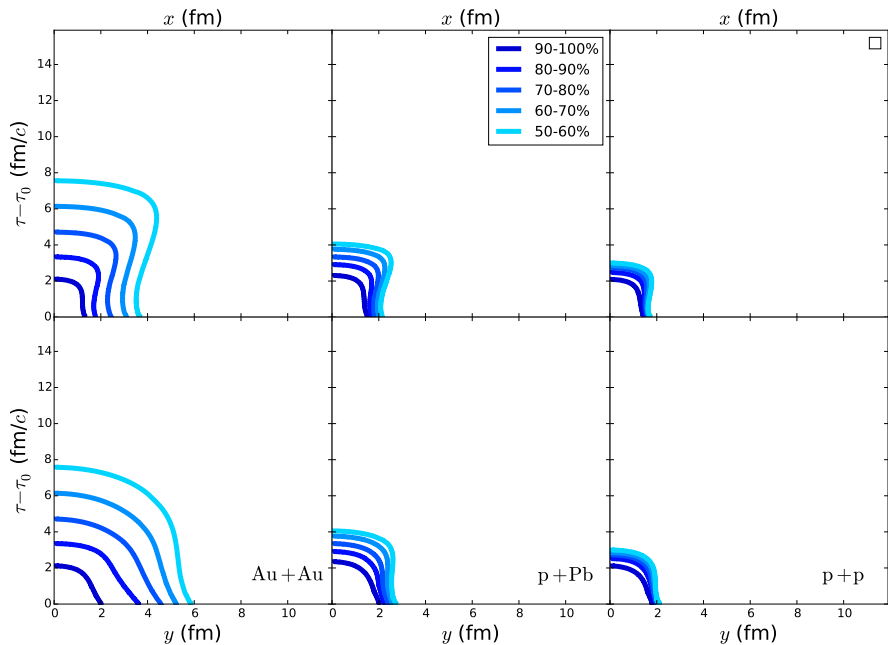
First: what kinds of geometries does hydrodynamics predict?

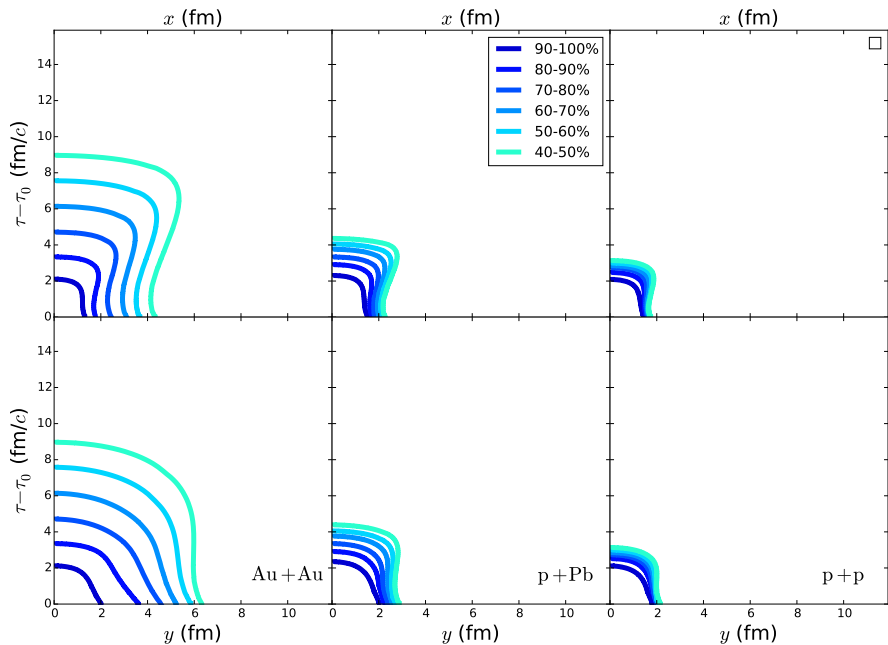


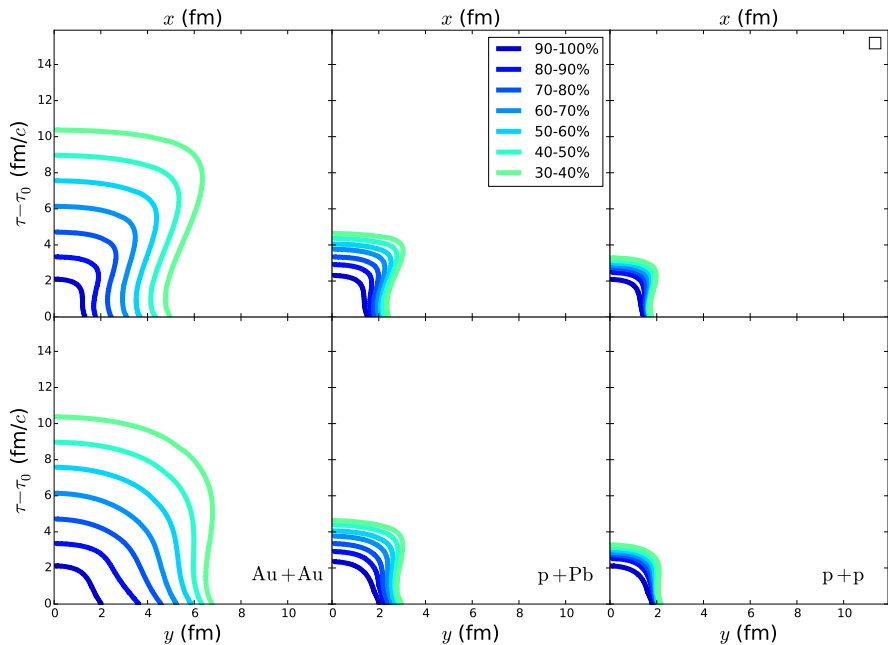


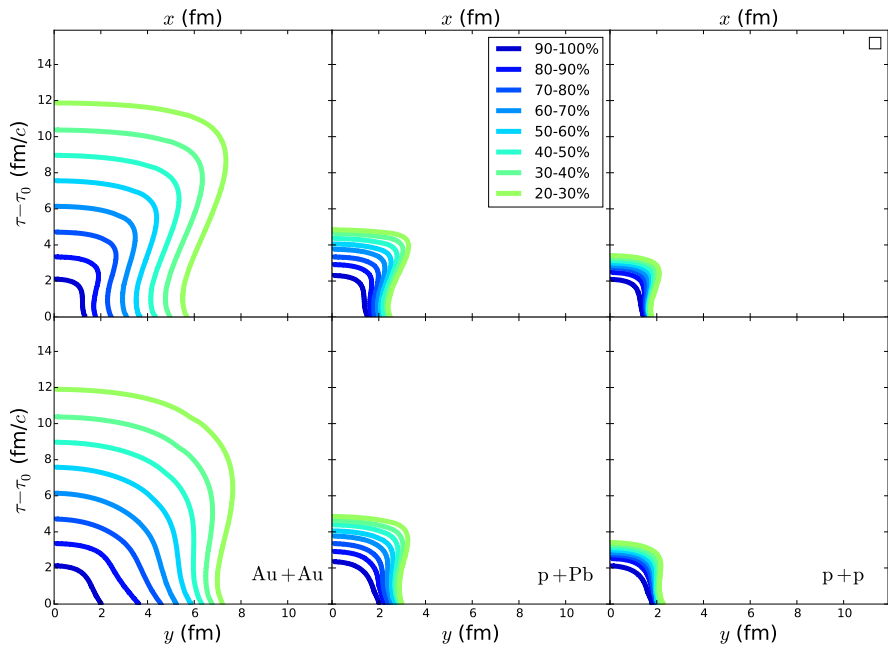


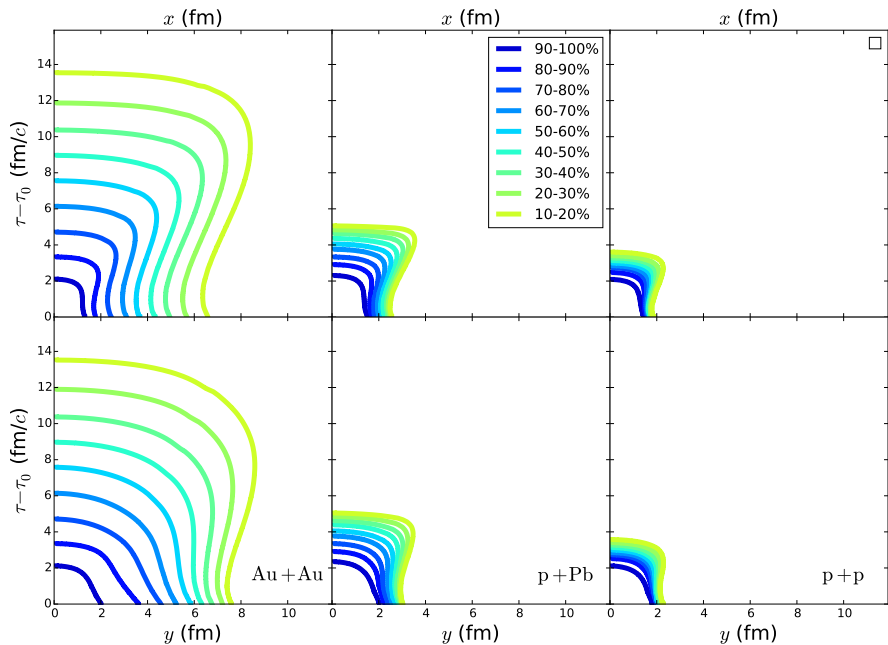


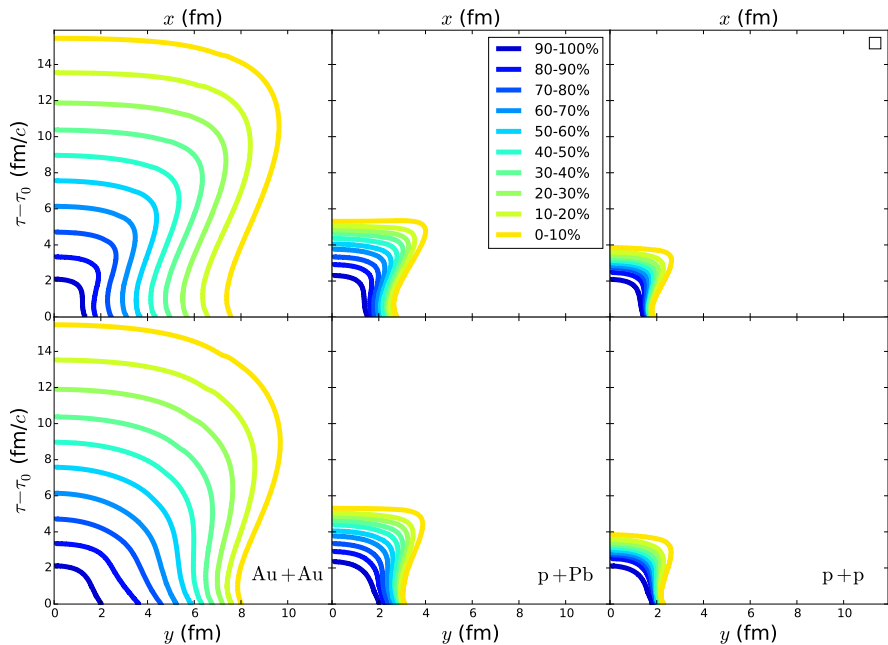


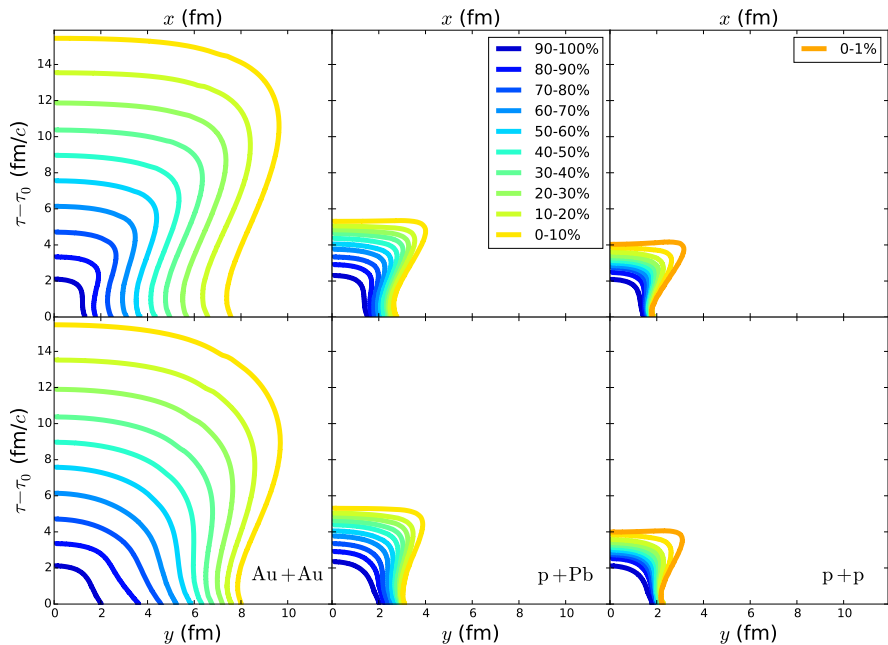


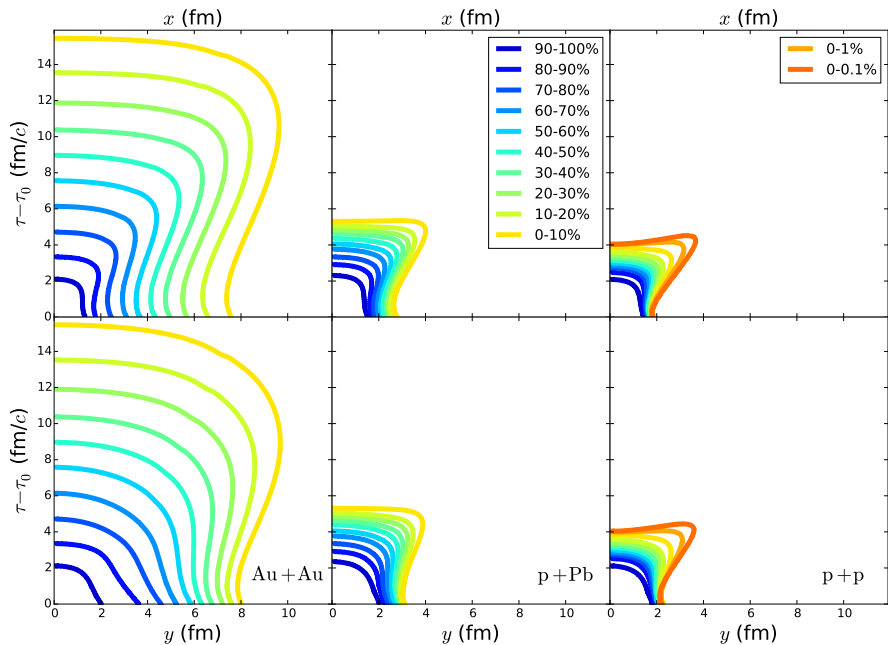


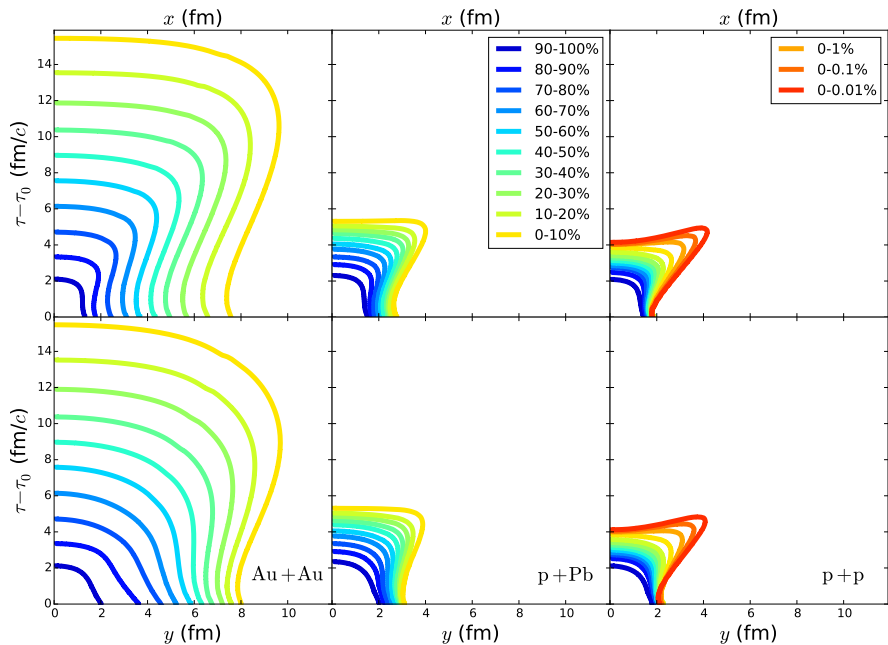


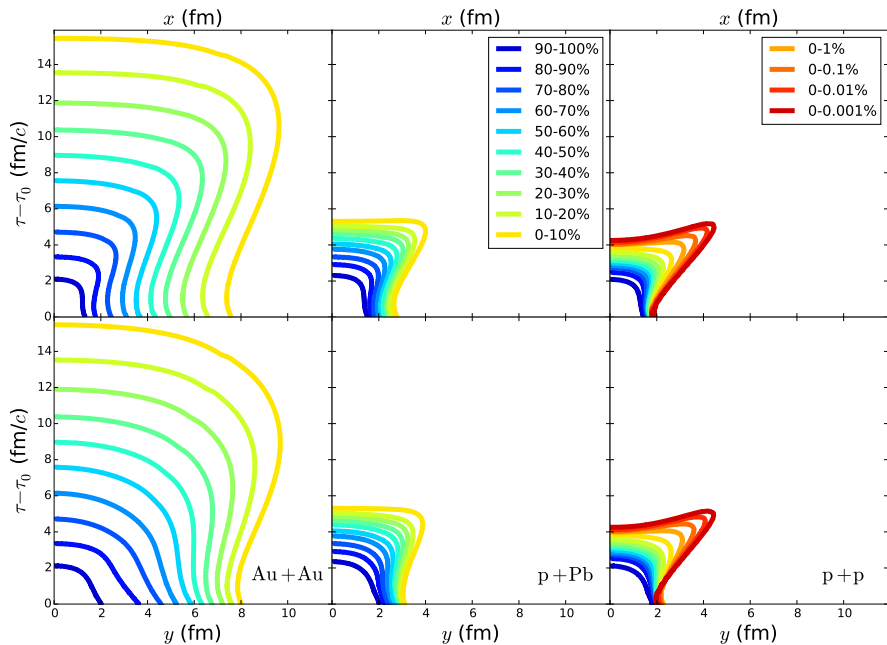




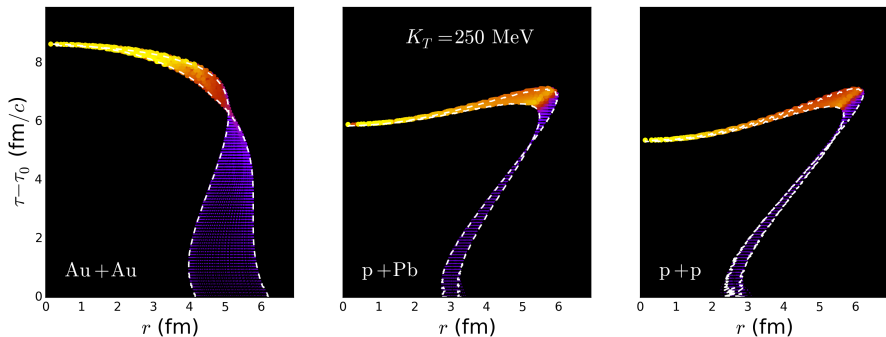








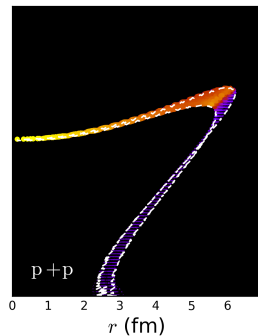
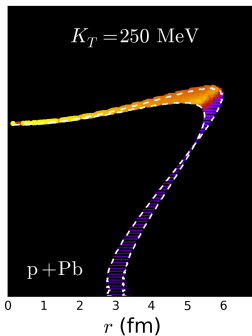
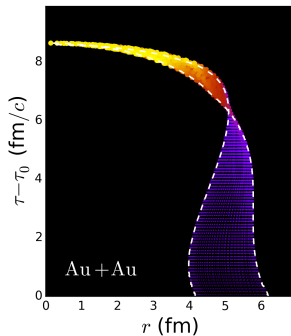
$$dN_{\text{ch}}/d\eta = 100$$



Observations:

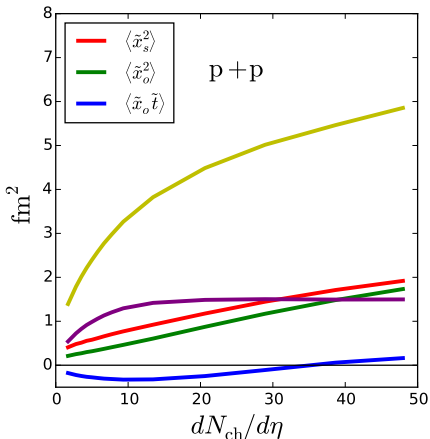
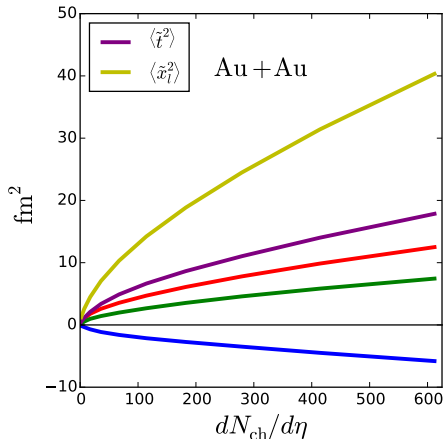
- ▶ Small systems highly elongated due to strong collective flow
- ▶ Geometric size of emission region $\langle \tilde{x}_i^2 \rangle$ similar across collision systems
- ▶ Emission duration $\langle \tilde{t}^2 \rangle$ decreases from Au+Au to p+Pb, p+p

$$dN_{\text{ch}}/d\eta = 100$$



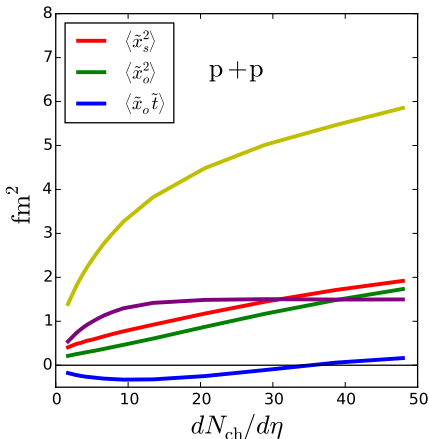
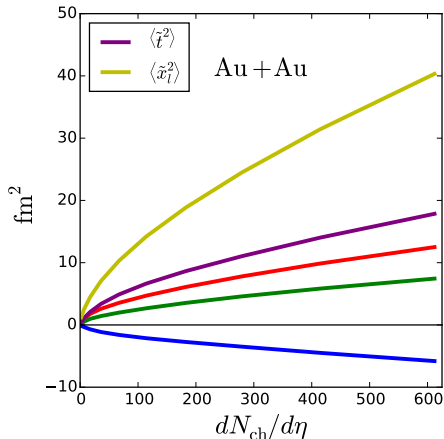
Expectations:

- ▶ Geometric contributions to R_i should grow with $dN_{\text{ch}}/d\eta$
- ▶ Emission duration and spatial-temporal correlation behave differently in small vs. large systems



$$R_s^2 = \langle \tilde{x}_s^2 \rangle \quad R_o^2 = \langle \tilde{x}_o^2 \rangle - \beta_T \langle \tilde{x}_o \tilde{t} \rangle + \beta_T^2 \langle \tilde{t}^2 \rangle \quad R_l^2 = \langle \tilde{x}_l^2 \rangle$$

- ▶ Geometry ($\langle \tilde{x}_s^2 \rangle$, $\langle \tilde{x}_o^2 \rangle$, $\langle \tilde{x}_l^2 \rangle$) scales *monotonically* with $dN_{\text{ch}}/d\eta$
- ▶ $\langle \tilde{t}^2 \rangle$ monotonic in Au+Au, levels off in p+p
- ▶ $\langle \tilde{x}_o \tilde{t} \rangle < 0$ monotonic in Au+Au, changes sign in p+p



$$R_s^2 = \langle \tilde{x}_s^2 \rangle$$

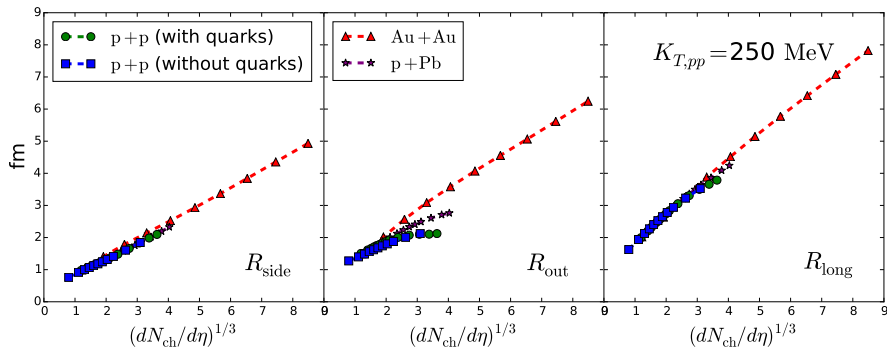
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⇒ All R_i scale strongly in Au+Au

⇒ R_o scales more weakly than R_s, R_l in p+p

Full hydro results



Hydrodynamics implies:

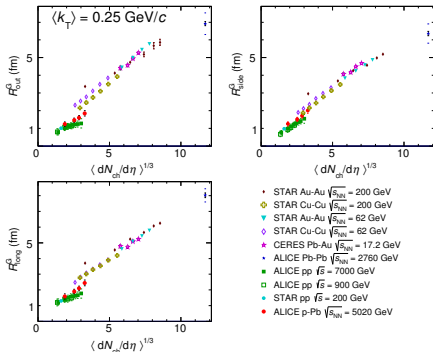
$$V \sim R_{out} R_{side} R_{long} \propto dN_{ch}/d\eta$$

Suggests:

$$R_i \propto (dN_{ch}/d\eta)^{1/3}$$

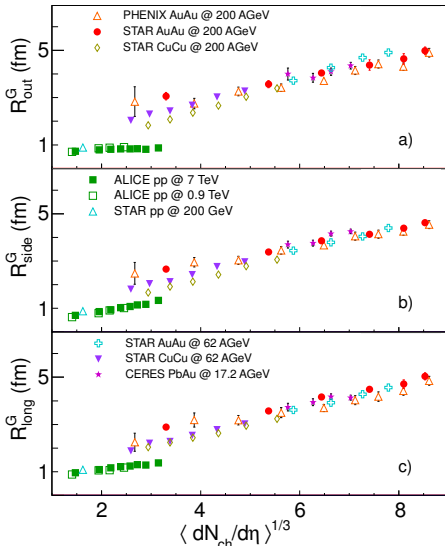
What does the data show?

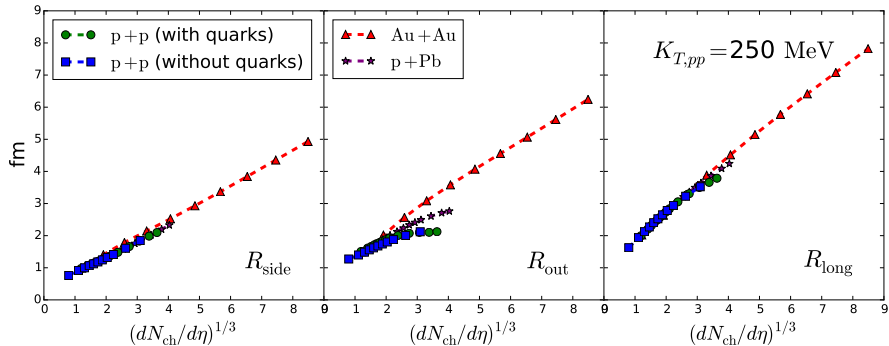
ALICE (2015)



- ▶ Radii different in pp vs. AA
- ▶ Slopes different in pp vs. AA
- ▶ Slope for pp R_{out} less than for R_{side} , R_{long}

ALICE (2011) - $K_T = 400 \text{ MeV}$





Overall, hydrodynamics does fairly well:

	System radii	System slopes	Slope hierarchy
R_{out}	✓	✓	✓
R_{side}	✗(?)	✓	✓(?)
R_{long}	✗(?)	✗(?)	✓(?)

Conclusions

- ▶ Data show non-trivial hierarchy of $dN_{\text{ch}}/d\eta$ -scaling in p+p R_o^2 as compared with R_s^2, R_l^2
- ▶ Same feature is *not* present in Au+Au

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 - Small systems have much stronger collective flow than large systems
 - Flow generates *very* different geometries for different systems at same multiplicity

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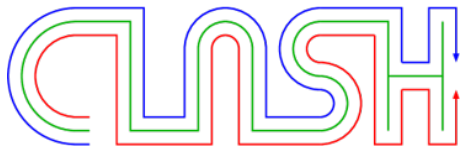
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Summary assessment: hydrodynamics reproduces qualitative features, but simple approach is quantitatively lacking

Conclusions

So what's missing?

- ▶ Initial-state fluctuations:
 - may shift multiplicity-dependence of average radii
 - enables new observables which probe fluctuations of shape and size
- ▶ Hadronic phase:
 - generates additional flow and smaller source sizes without (significantly) changing multiplicity
 - naturally relaxes of $\beta_L = 0$ assumption



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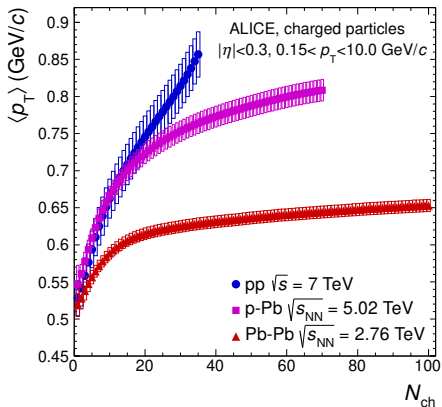
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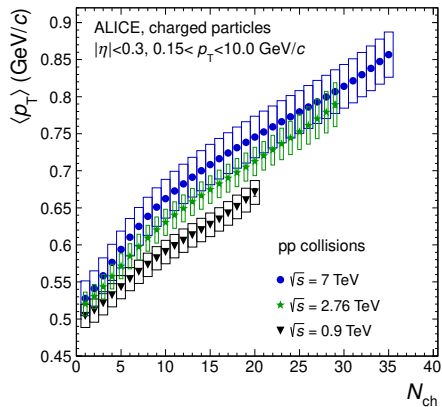
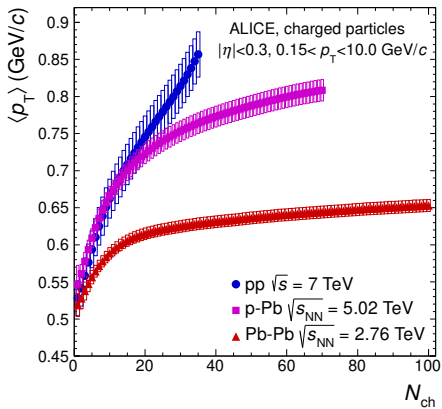
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Děkuju!

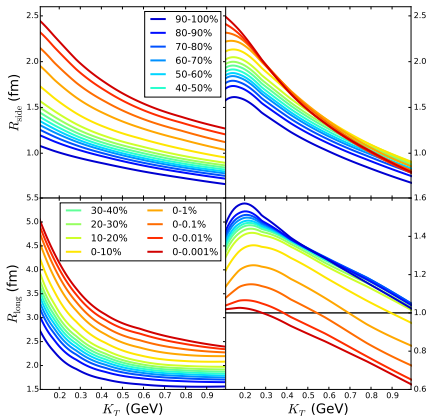


Backup slides

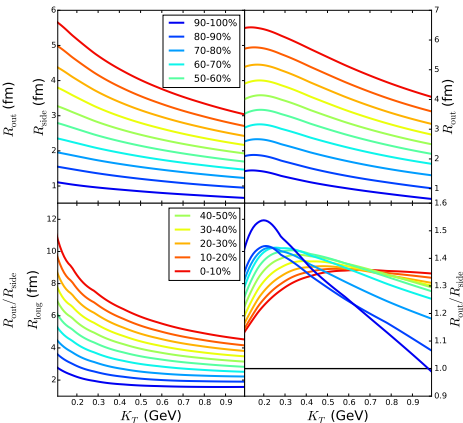




p+p



Au+Au



Generic consequences of strong collective flow:

- ▶ K_T -scaling
- ▶ Possibility of $R_o^2/R_s^2 < 1$

p + p (with quarks)

