



MuCol: a new Muon Collider initiative for Europe

Lund Science Coffee seminar

PRESENTED BY NATALIA MILAS

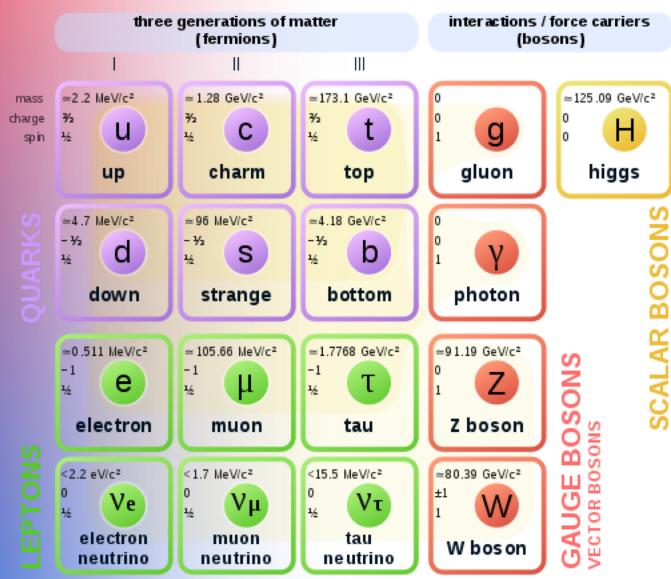
2022-09-05



1 Why Muons?

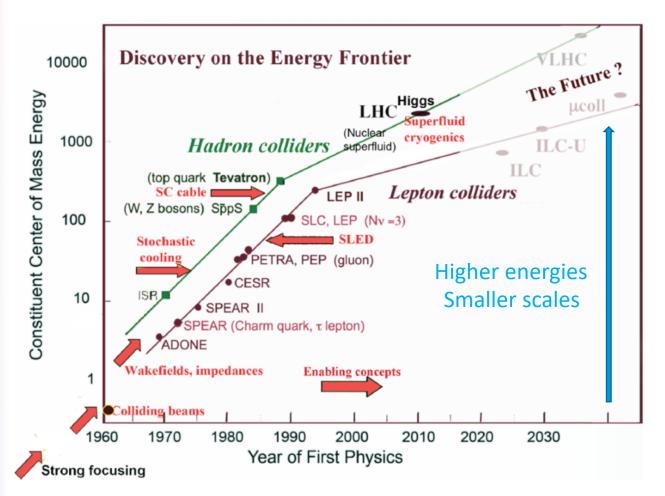
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Why Muons?



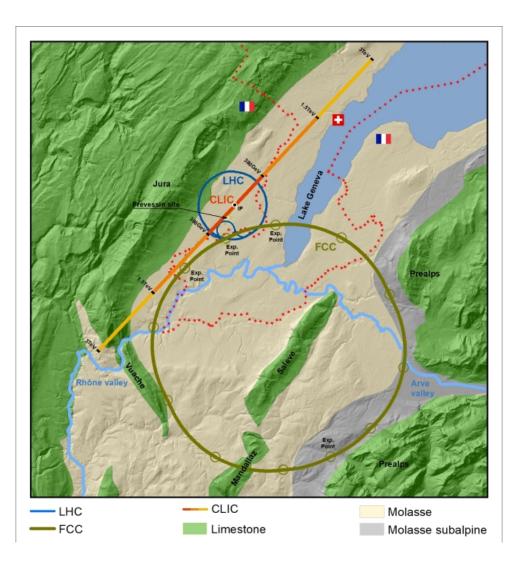
Standard Model of Elementary Particles

What limits a collider energy?



- Hadron (proton) colliders are limited by bending magnetic field
 - This limits the bending radius -> tunnel length
- Electron rings are limited by Synchrotron radiation
 - Synchrotron radiation is much stronger for low mass particles
- Electron linacs are limited by available acceleration
 - Cannot accelerate electrons without extremelly long tunnels

Example of Colliders

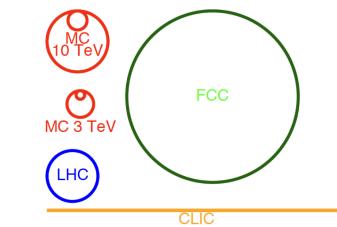


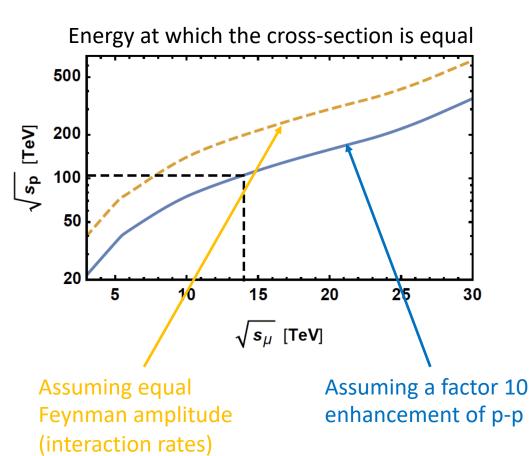
- LEP/LHC:
 - 14 TeV proton-proton (LHC)
 - 209 GeV for electron-position (LEP)
 - 27 km circumference
- FCC (proposed):
 - 90-350 GeV for electron-position
 - 100 TeV for proton-proton
 - 100 km circumference
- CLIC (proposed)
 - 500 GeV (3 TeV) for electron-position
 - 11 (50.1) km



Advantages of using Muons

- Proton collisions have a lot of "wasted energy"
- More energy available for the collision since muons are elementary particles
- m_µ = 200 m_e this mitigates the radiation issues since power dissipated in synchrotron radiation ∝ 1/m⁴





https://muoncollider.web.cern.ch/introduction

Physics Case Behind the Muon Collider

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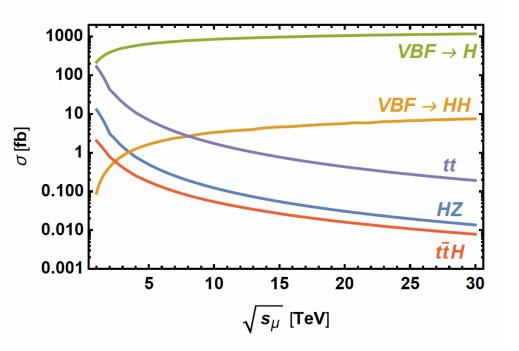




What can be done with the Muons ?

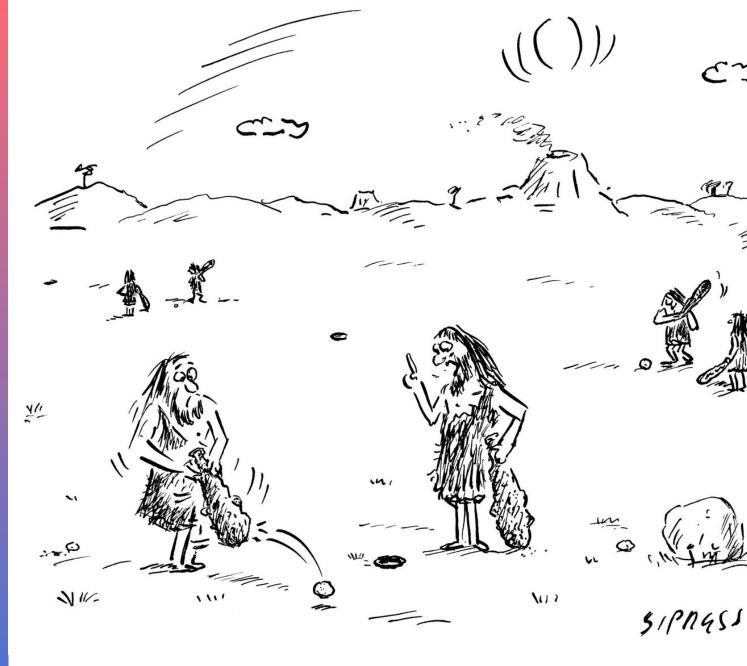
An accelerator physicist view

- Because of lack of breamstrahlung a muon collider can have a small energy spread thus allowing precision measurements of mass width.
- Direct exploration on several beyondthe-SM questions:
- search extensively for new particles coupled with the Higgs boson,
- Dark matter
- As well as probe **New Physics** indirectly:
- some of the key processes for Higgs physics, have very large cross-sections and a muon collider might allow a satisfactory program of Higgs couplings determination.



- The decay of the stored muons lead to a neutrino flux about 1000 times the ones presently available.
 - Neutrino oscillations and other interesting neutrino physics.

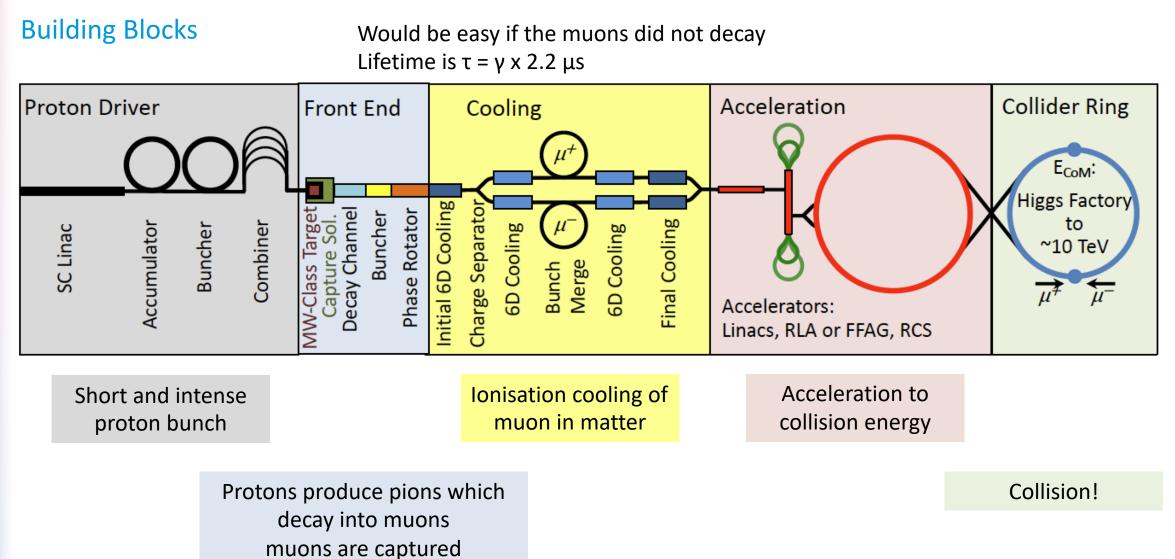
Main components of a Muon Collider



"Maybe we need different types of clubs for different types of situations."

A Muon Collider in a nutshell



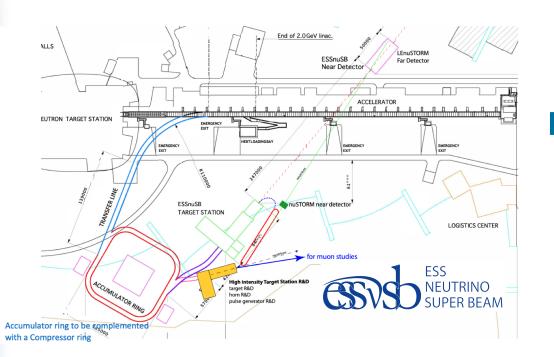




Creation of short and intense proton bunch

O(kA) peak current

- Parameters are driven by the main collider.
- Latest parameters (MuCol): 2MW, 5Hz with a possible upgrade to 5 MW.
- 10¹⁴ 10¹⁵ protons per bunch
- 1 ns bunch length at the Target

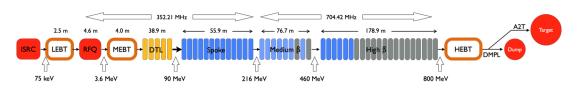


In order to meet the Muon requirements:

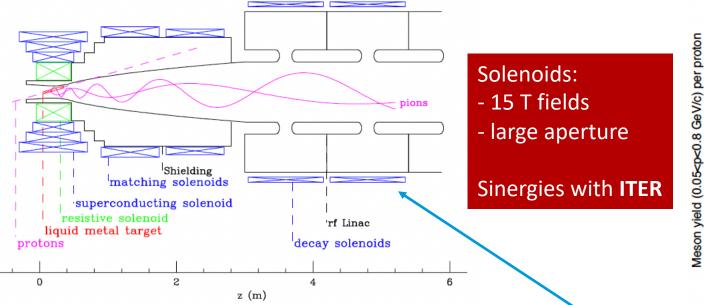
- Increase the current and/or
- Increase the Pulse length **and/or**
- Increase the energy

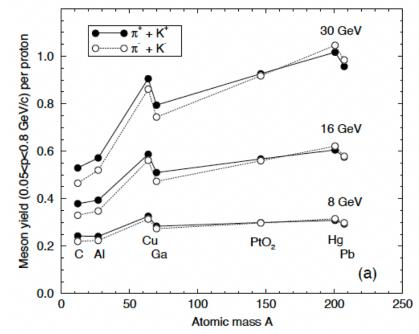
Protons to the target:

- 5 MW
- 1 us pulse
- About 10¹³ protons per pulse
- 14 Hz rep rate
- 2.5 GeV



Pion production and capture



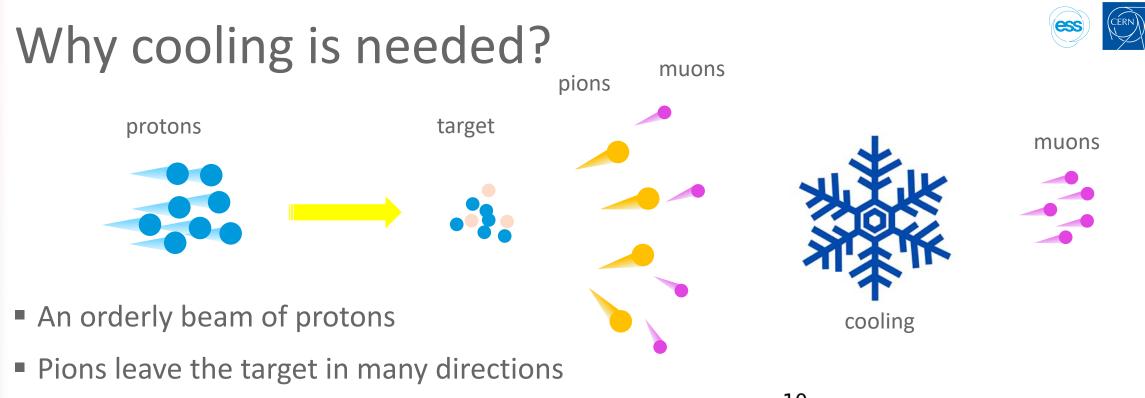


- The high muon beam intensity for the collisions requires a very high efficiency pion/muon capture
- We aim at colliding ¾ of the muons that are cooled.
- However estimates show that only ¼ of the muons survive the capture and cooling.

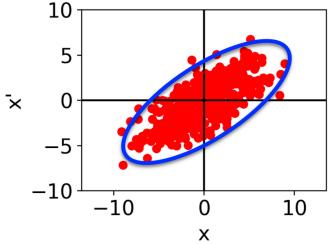
Phase rotation linac

The pions (and muons) created in the collision of protons with the target have an rms energy spread of 100% and mean energy of 100 MeV.

Phys.Rev STAB, vol 2, 081001 (1999)



- Pions decay also in many different directions
- Low Brightness beam
- High Emittance beam

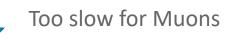




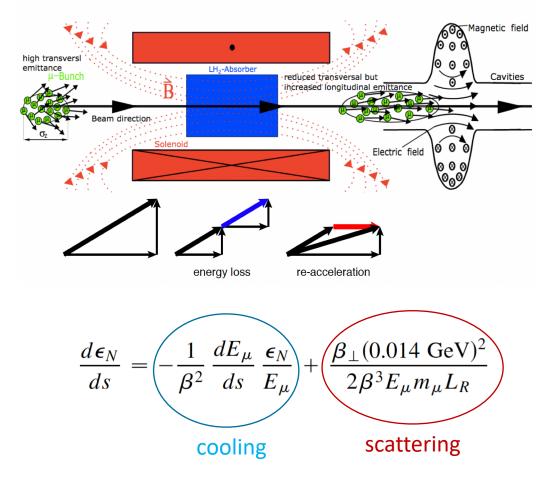
Muon cooling

Cooling must happen in a time scale compatible with the muon lifetime.

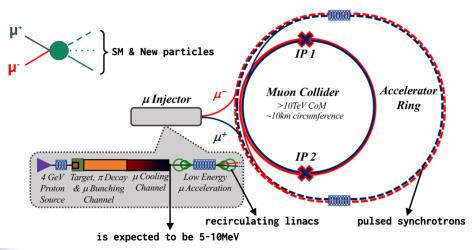
- Electron cooling : overlap the beam with another with matching emittance and transfer the "heat" from one to another
- Stochastics cooling: Measured the beam profiles and have a set of correctors to shrink it



- Ionisation cooling: Uses the fact that muon weak interaction with matter. Pass the beam through a material which causes energy losses through ionization (transverse and longitudinal). Replenishes the longitudinal part using RF -> Effective reduction of transverse momenta.
- For the longitudinal plane cooling -> have a trajectory that is energy dependent and shape the ionisation material as wedges -> higher energy particles loose more energy



Acceleration and collision



Muons will decays into electrons and neutrinos inside the collider. Where all those particles end and what is the effect over machine components and beyond?

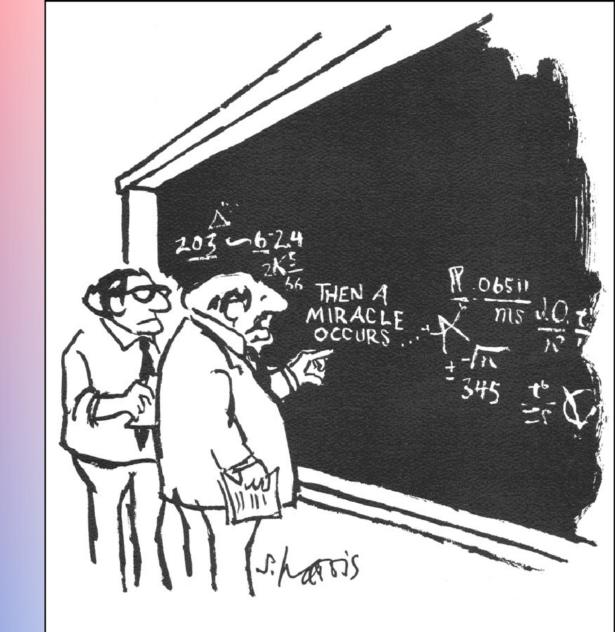
Acceleration:

- Crucial to be as fast as possible up to ~100 GeV because of the muon decay time
- Muon energy of cooling channels is ~ 10 MeV -> linacs are more efficient (recirculating linacs are options)
- Latter FFAs or RCS are more efficient.
- Likely the most expensive part of the complex.
- Collider:
 - Receives 2 bunches of particles, a μ^+ and a μ^- one.
 - Mainy constraints to be able to deliver the required luminosity.
 - IR with extremelly low β^* of a few cm (for the 3 TeV case)
 - Acceptable dynamic aperture
 - Strong instability effects (high charge and small bunch length).
 But in this case the Muon decay my be a blessing...



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What are the challenges?

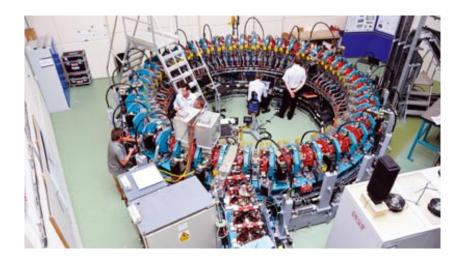


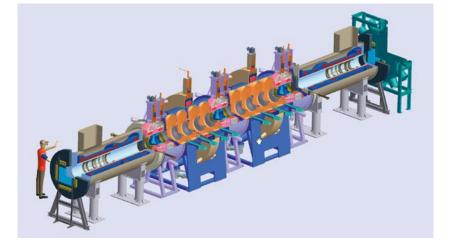
"I THINK YOU SHOULD BE MORE EXPLICIT HERE IN STEP TWO,"

Muon R&D in the last decade

- MERIT:
 - Demonstrated principles of pion production in solenoid fields
- EMMA:
 - Demonstrated acceleration in FFAs
- MUCOOL:
 - Radio-frequency accelerating cavities R&D
 - Demonstrated operation of cavities at high gradient inside strong magnetic fields
 - Breakdown suppression using high pressure gas
 - Careful RF coupler design and cleaning
- Muon Ionization Cooling Experiment (MICE)
 - Need to increase the beam brightness
 - Proof-of-principle of Ionization cooling

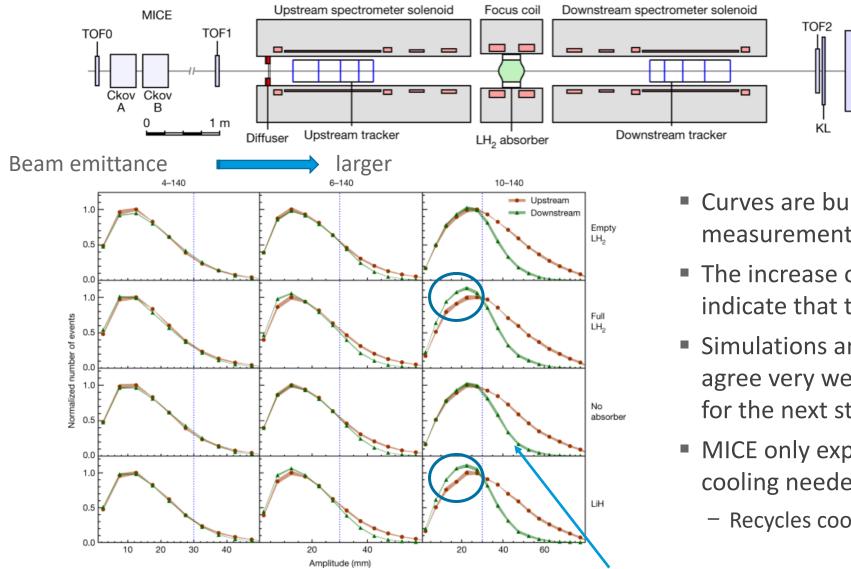






The MICE experiment





Curves are build ups of single muons measurements

EMR

- The increase on muons in the core indicate that there is cooling
- Simulations and measurements agree very well which is important for the next stages modeling
- MICE only explored a fraction of the cooling needed for a Muon Collider
 - Recycles cooling -> RINGS

interactions with any material bathed by the neutrino radiation disk.

Table 1: Radial distance, R, from the ring center with center-of-mass energy, \sqrt{s} , and depth, d, needed to reduce neutrino-induced dose at surface to DOE (100 mrem) and Fermilab (10 mrem) annual off-site limits at N_D decays/yr.

The Neutrino problem

The radiation hazard arises from the showers of

Neutrino radiation and detector background

penetrating charged particles produced through neutrino

	\sqrt{s} (TeV)	0.5	1	2	3	4
	$N_D imes 10^{21}$	0.2	0.2	2	2	2
100 mrem	R (km)	0.4	1.1	6.5	12	18
	<i>d</i> (m)	≤ 1	≤ 1	3.3	11	25
10 mrem	R (km)	1.2	3.2	21	37	57
	<i>d</i> (m)	≤ 1	≤ 1	34	107	254
						7

Depth of LCH tunnel: between 45 and 170 m ESS public average dose: 10 mrem/year

arXiv:hep-ex/0005006v1 3 May 2000

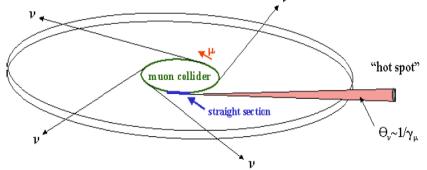


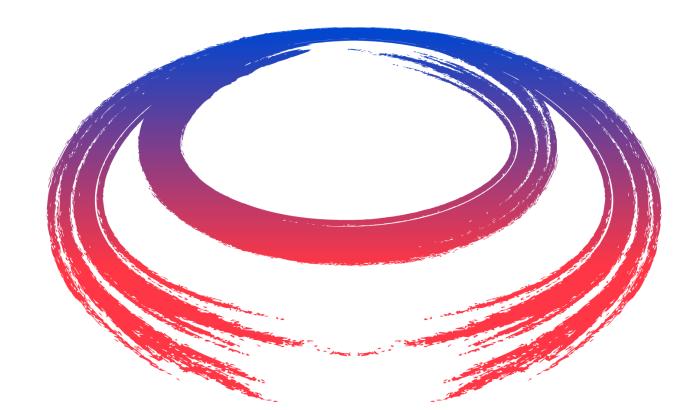
FIGURE 1. The decays of muons in a muon collider will produce a neutrino radiation disk emanating out tangentially from the collider ring. Radiation hot spots in the disk will occur directly downstream from straight sections in the collider ring.

center of mass energy, E _{CoM}	0.1	4 TeV	10 TeV	100 TeV
additional description	H ⁰ factory	"lite"	2 nd gen.	3 rd gen.
collider luminosity, \mathcal{L} [cm ⁻² .s ⁻¹]	$1 imes 10^{31}$	$6 imes 10^{33}$	$1 imes 10^{36}$	$1 imes 10^{36}$
collider int. lum., $\int \mathcal{L} [fb^{-1}/yr]$	0.1	60	$10\ 000$	10000
muon beam energy, E_{μ} [TeV]	0.05	2	5	50
muon decays/yr, N^+_{μ} [10 ²⁰]	6	0.08	8	0.4
collider reference depth, d [m]	10	300	100	100
ν beam distance to surface, L [km]	11	62	36	36
ν beam radius at surface [m]	24	3.3	0.8	0.08
ave. rad. dose in plane [mSv/yr]	2×10^{-5}	$5 imes 10^{-4}$	2.3	10
str. sec. len. for 10x ave. rad. [m]	1.9	1.1	1.0	4.2



MuCol initiative

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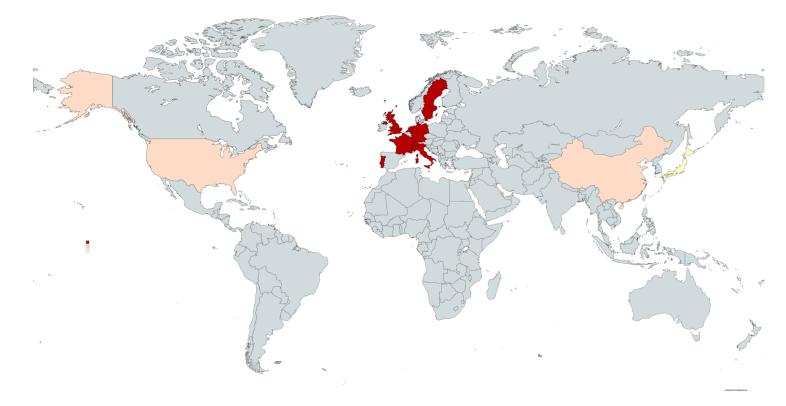


UON Collider Collaboration

Collaboration in EU and beyond

- Collaboration meetings started in the beginning of 2021.
- Presentation of an new Muon Collider Study to the CERN Council

Preparation and Submission of a EU grant **HORIZON-INFRA-2022-DEV-01** -> Approved in August 2022 and starting January 2023 (running for 4 years)





Next Steps

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Studies in the pipeline

- Review proton complex
 - Average power of 2 MW is no problem
 - Merging into 5 pulses of 400 kJ per second needs to be verified
 - Target design to sustain a 5Mw beam
- A lot more simulation and R&D work on the cooling part (crucial for the project)
- Collective effects across the whole complex to identify botlenecks
 - Review apertures, feedback and other specifications
 - First results for aperture requirements
 - Potential instability of interaction of muon beam with mater
- Preliminary deisgn of the acceleration complex chain
- Power and cost optimisation
- Vacuum and absorber, instrumentalon, cryogenics, magnets, RF...
- Reuse of existing infrastructure, e.g. LHC tunnel to house the accelerator



ESS, CERN and Uppsala



In the final Muon collider a cooling facto of 10⁶ is needed

Planning **demonstrator** production target and c

Suitable site exists on C proton beam

could combine with

Other sites should be e



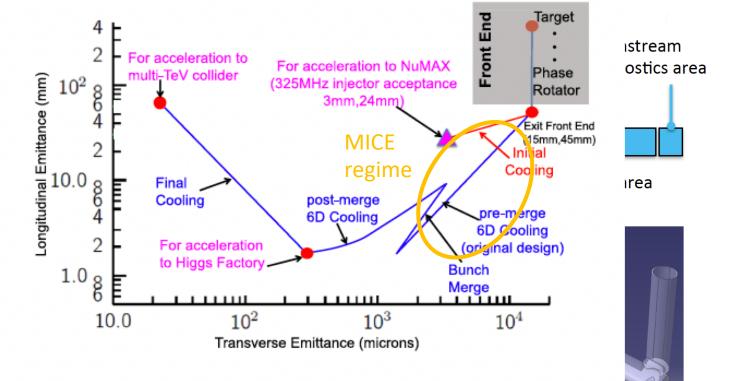


Fig. 3. Ionization cooling path in the 6D phase space.

R. Losito et al.



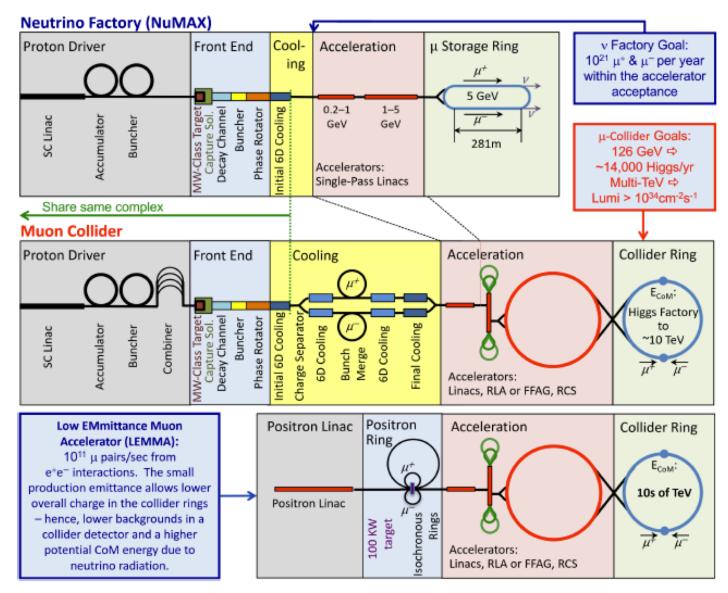


Thank you! Questions?

2022-09-05



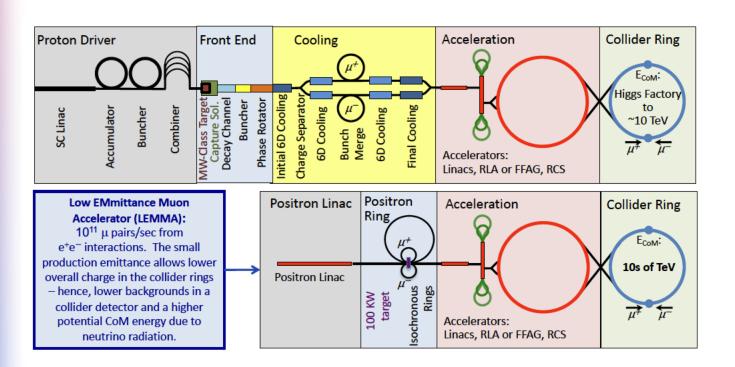
Muon Collider Staging



An alternative to the proton driver

An alternative driver

A novel approach of the Low Emittance Muon Accelerator (LEMMA) based on muon pair production with a positron beam impinging on electrons at rest in a target, was recently proposed and is now under conceptual study.



Pros:

- Muon beam with smaller emittance – no cooling needed
- Muon produced already at 22 GeV (500 μs lifetime)

■ **R&D**:

- Intensity requirement of the positron source (order of magnitude higher than ILC levels)
- Problem with beam density on the target