



UC SANTA CRUZ



PARTY CALL PHYSICS

Likelihoods and Classifiers

Dr. Giordon Stark 

Lund Seminar

September 15th, 2022

 giordonstark.com



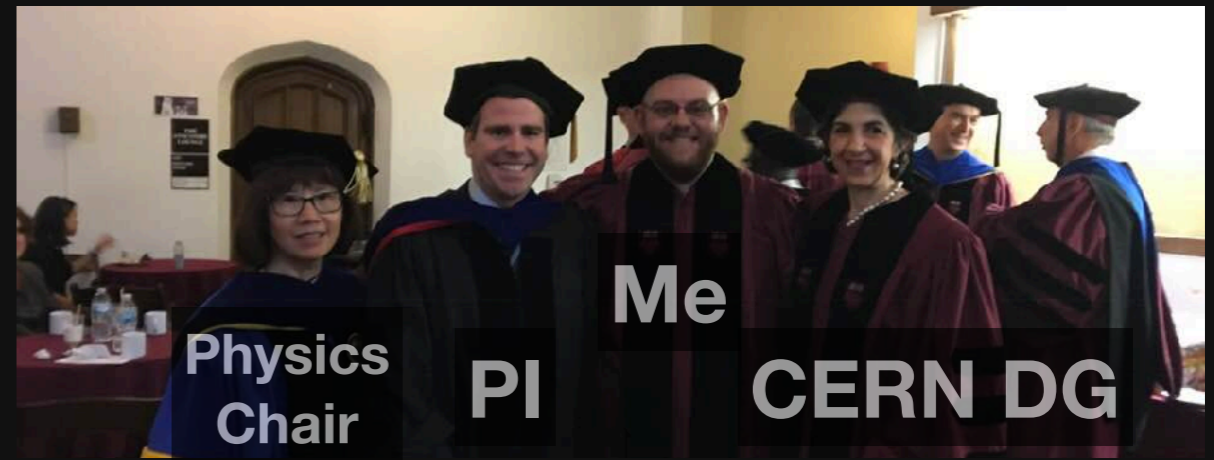
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2016-06-04 03:47:03 **1**

if you can read this, you're too close

About Me



- ✦ B.S Caltech 2012 (LIGO)
 - ✦ Brownian Thermal Noise
- ✦ Working on ATLAS since 2014
- ✦ PhD UChicago 2018 (ATLAS)
 - ✦ Search for new (hadronic) physics and instrumentation upgrades
- ✦ Currently postdoc at SCIPP, UC Santa Cruz since 2018
 - ✦ Large-scale physics analysis combinations, software development, and instrumentation upgrades
- ✦ Lots of outreach/teaching/DEI experience (bootcamps, workshops, committees)

Overview of Today

The long road to making physics accessible
(and how rocky it has been at times)

- ✦ The Standard Model... and **beyond!**
- ✦ The Large Hadron Collider, **ATLAS**, and you
- ✦ Searching for signs of **new physics**
- ✦ Experimentalist introduction to **statistics and hypothesis testing**
- ✦ **Accessibility** in physics education
- ✦ What does the **future** have in store for us?

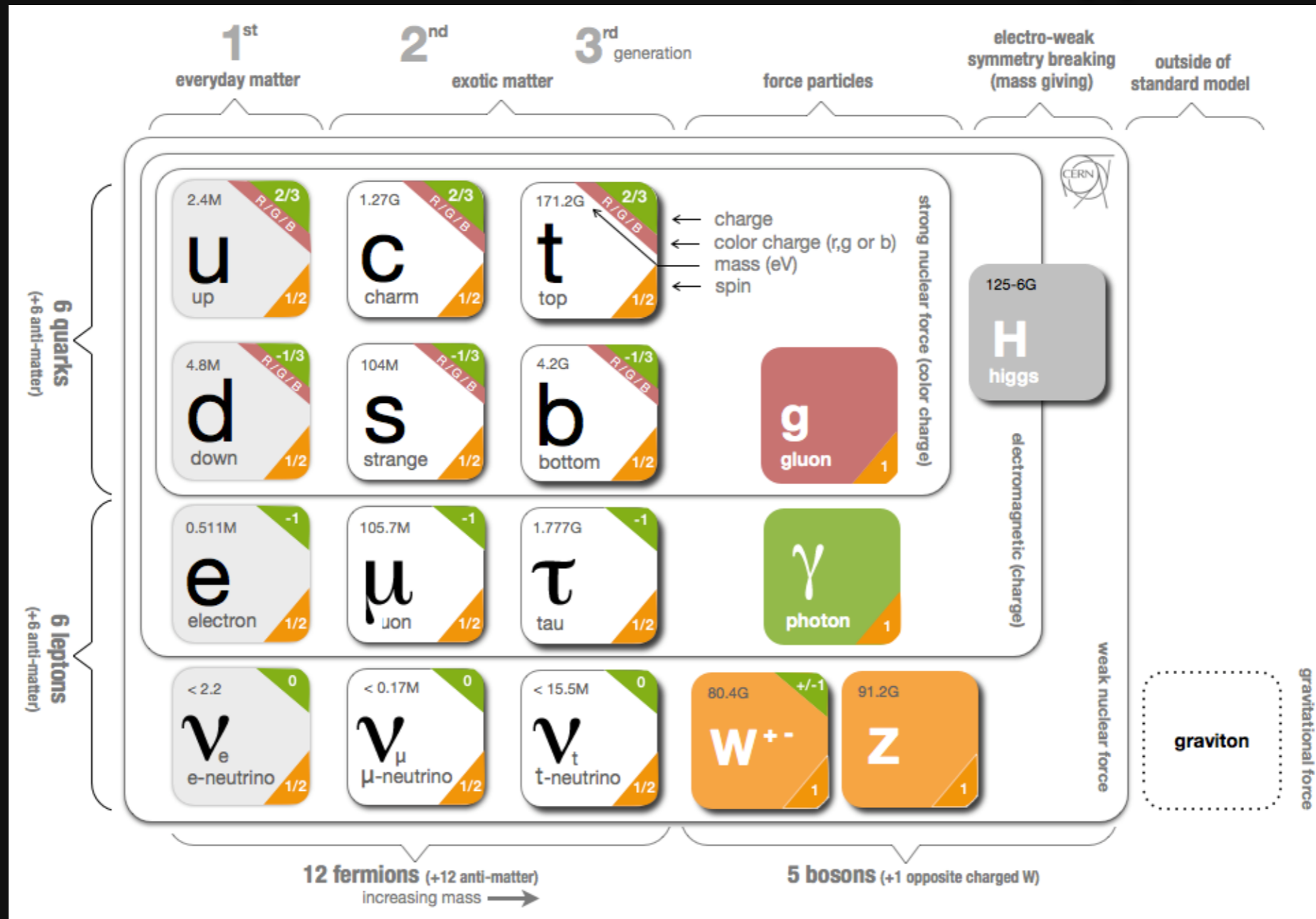
 **Let's get started**

The Standard Model

A study of particles and their interactions

*“The story so far: In the beginning the Universe was created.
This has made a lot of people very angry and been widely
regarded as a bad move.”* — Douglas Adams

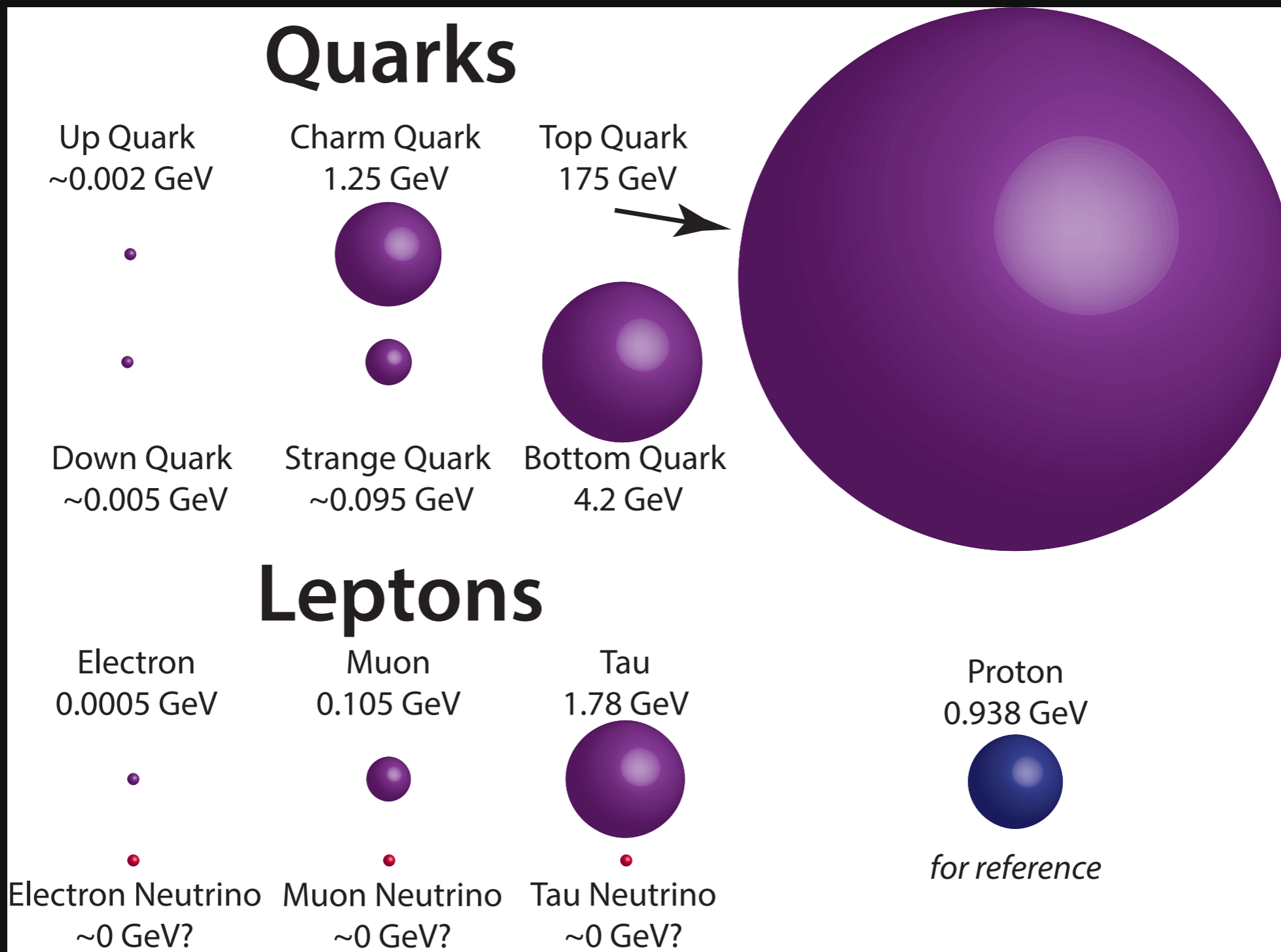
The Standard Model (I)



12 fermions
(6 leptons, 6 quarks)
1/2-spins

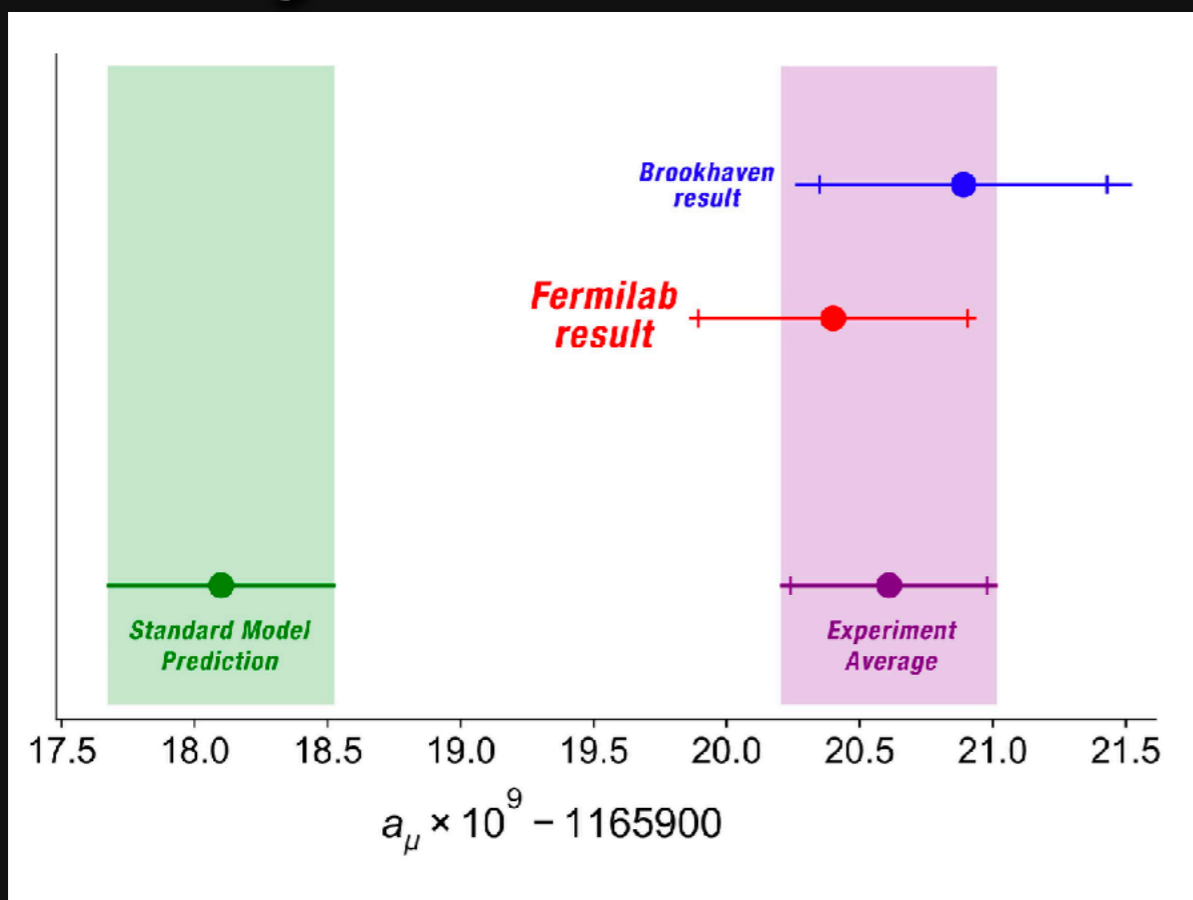
5 bosons
(4 force-carriers)
integer spins

The Standard Model (II)

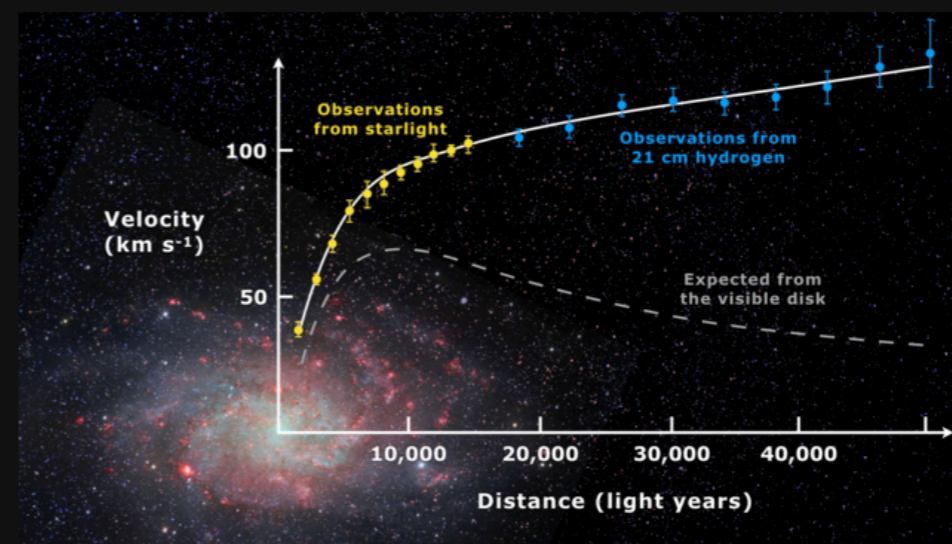


Volume proportional to the mass

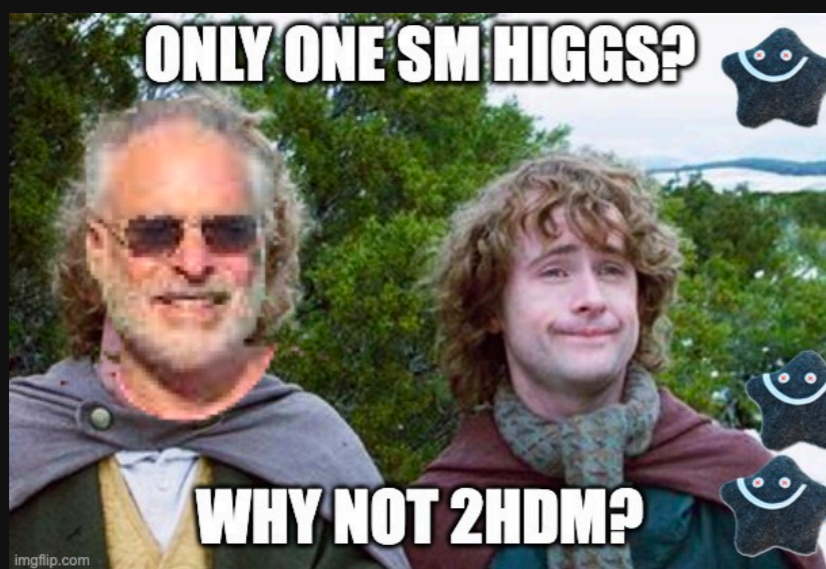
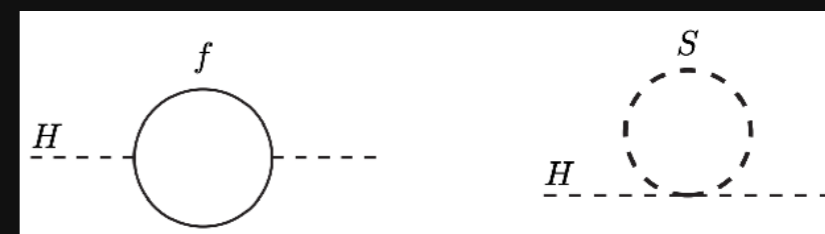
Beyond the Standard Model



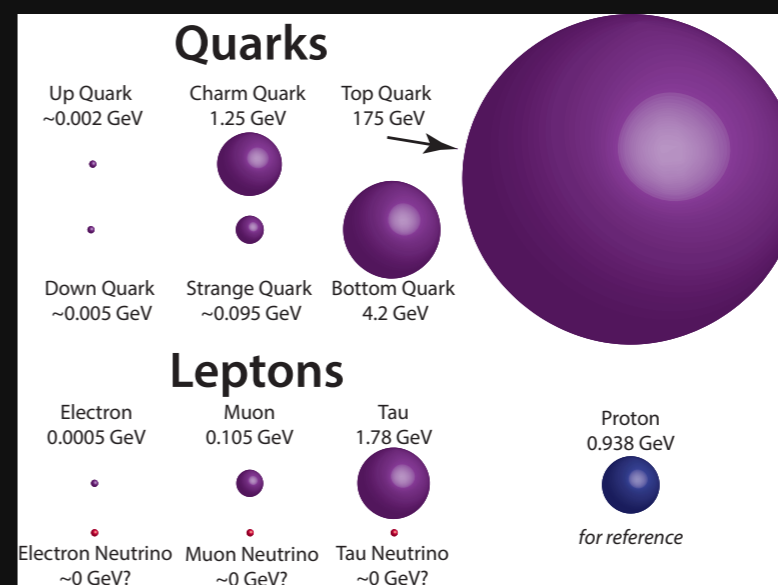
anomalous magnetic dipole moment of muon



Cold Dark Matter candidates?



Flavour anomalies?



Hierarchy Problem

Beyond the Standard Model

What is dark matter?

Where did all the antimatter go?

Why does the standard model look the way it does?

Why is the weak force so much stronger than gravity? (Hierarchy problem)

Supersymmetry (SUSY) is a framework with good theoretical motivations in which theorists can study BSM physics



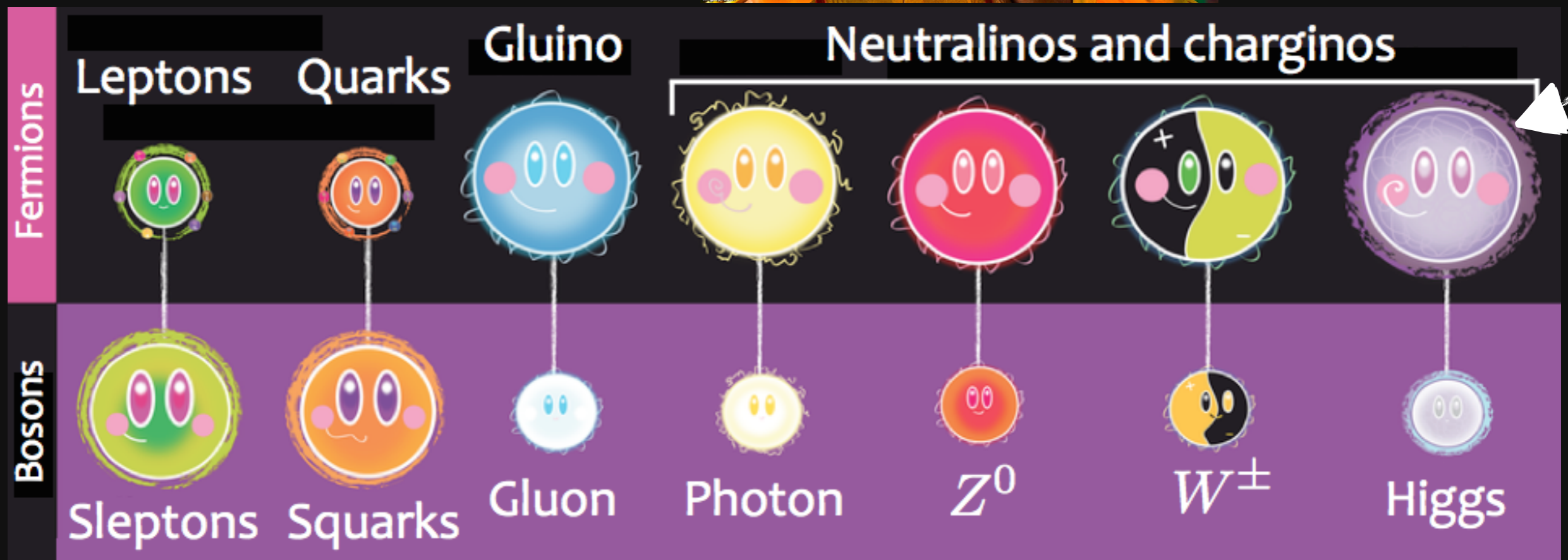
Supersymmetry (SUSY) is a set of benchmark models to help experimentalists answer these questions!

What is supersymmetry?

small = Standard Model
large = Beyond the SM



Higgsinos



A particle physics tango between fermions and bosons

The ATLAS Detector

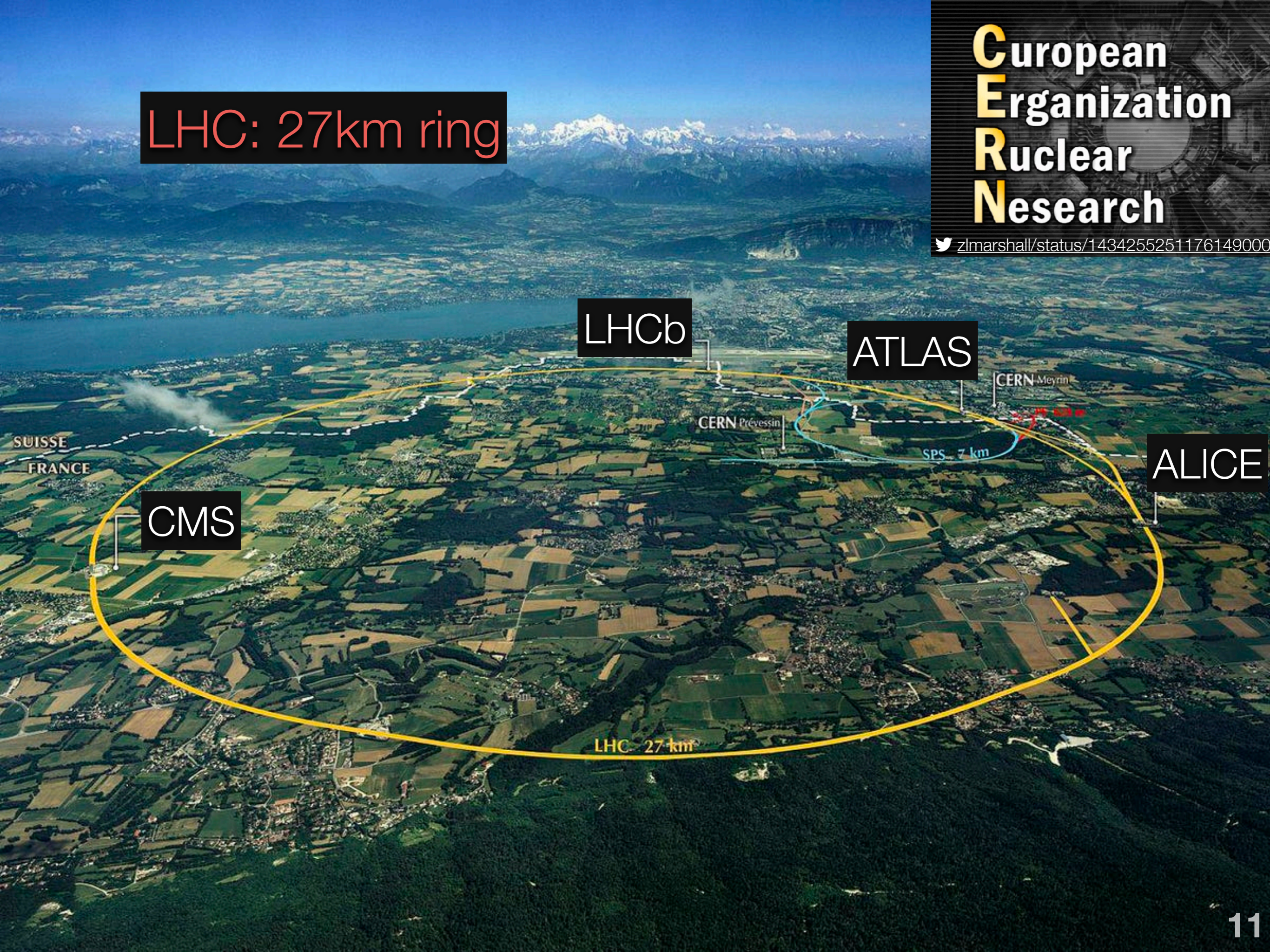
Taking pictures of proton-proton collisions

*“The single most important component of a camera is the
twelve inches behind it.”* — Ansel Adams

LHC: 27km ring

European Organization Nuclear Research

[zlmarshall/status/1434255251176149000](https://twitter.com/zlmarshall/status/1434255251176149000)



LHCb

ATLAS

CERN Meyrin

CERN Prévessin

SPS 7 km

SUISSE
FRANCE

CMS

ALICE

LHC 27 km

LHC

ATLAS

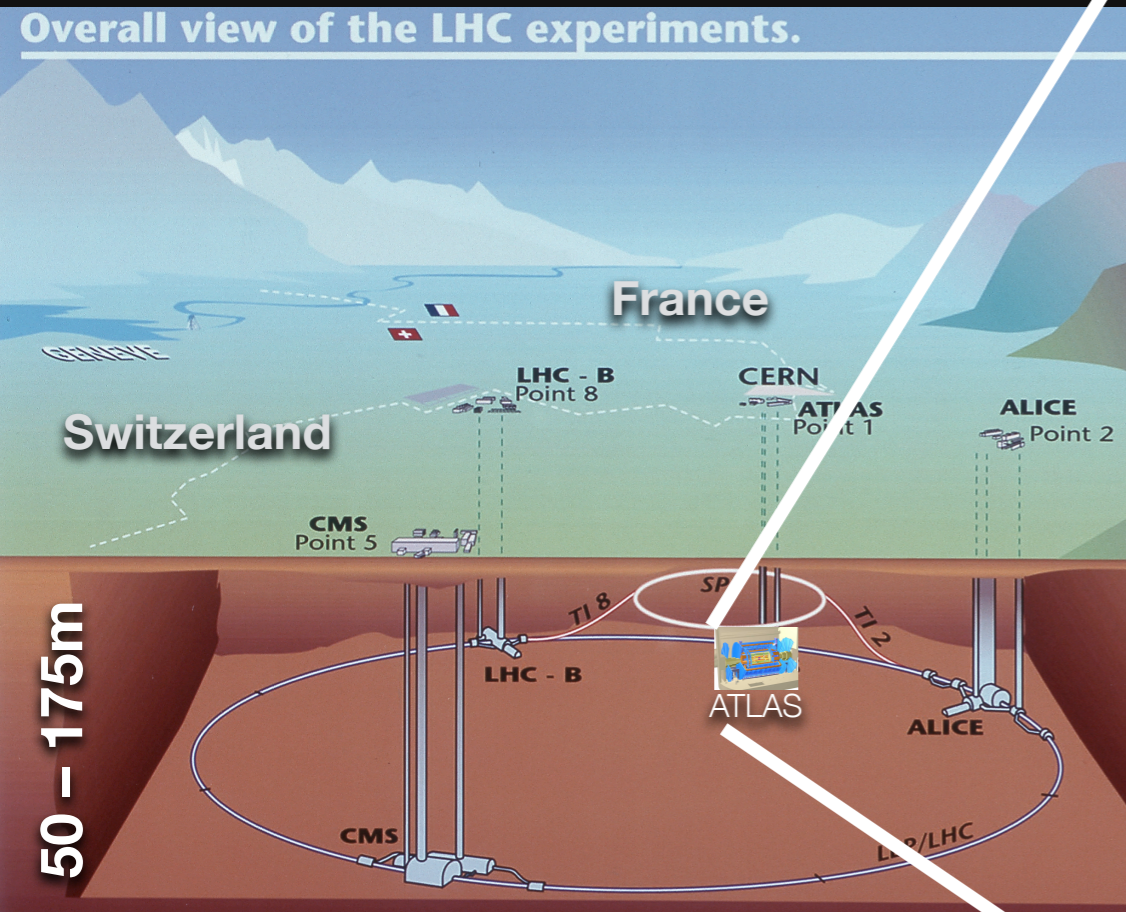
A collider and a detector

The **L**arge **H**adron **C**ollider is a massive, 27 km collider, operational since Sept. 2008

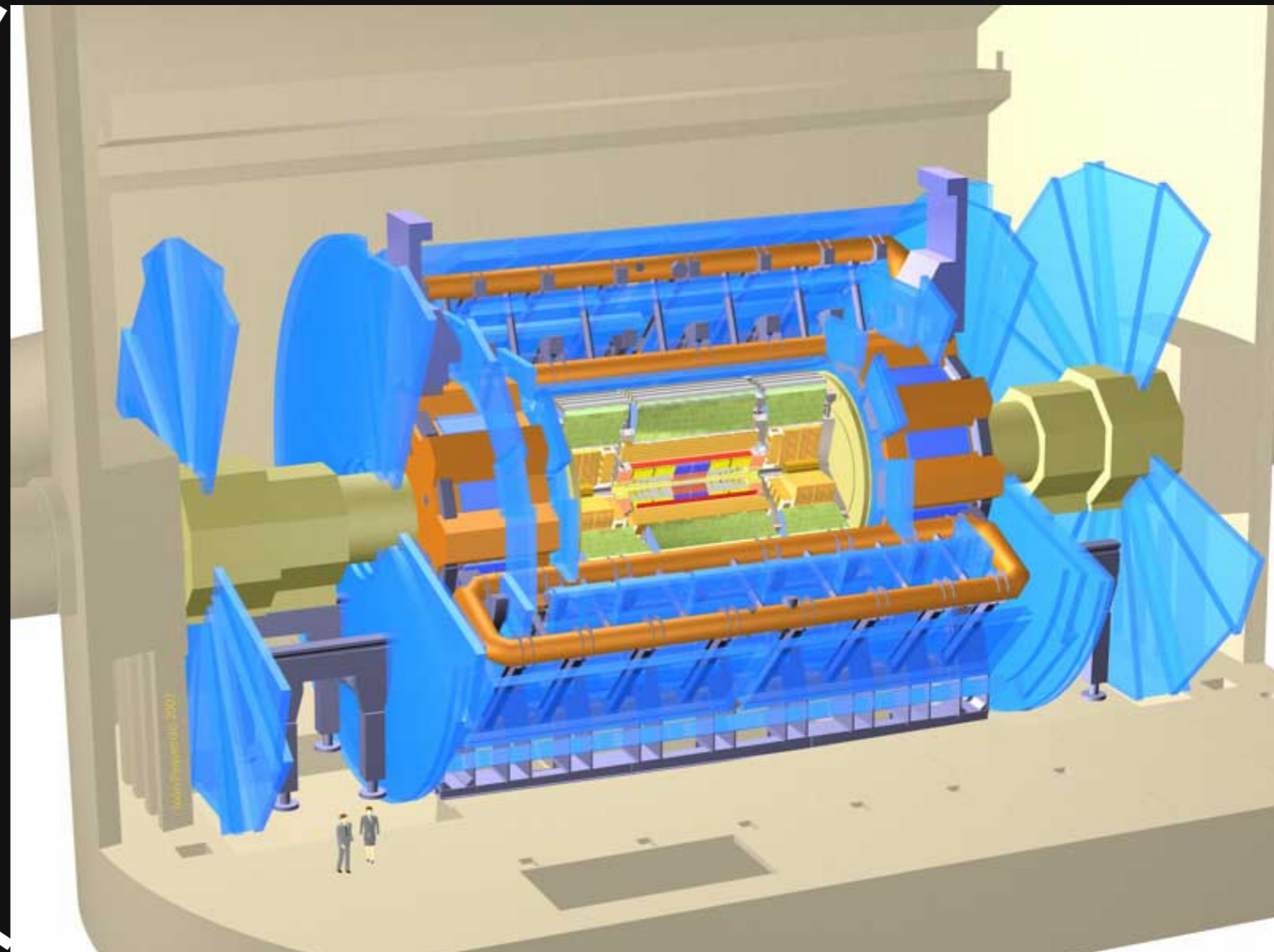
Four points along the ring at which the proton-proton beams cross

ATLAS is a large 7000 ton general purpose detector (46m x 25m)

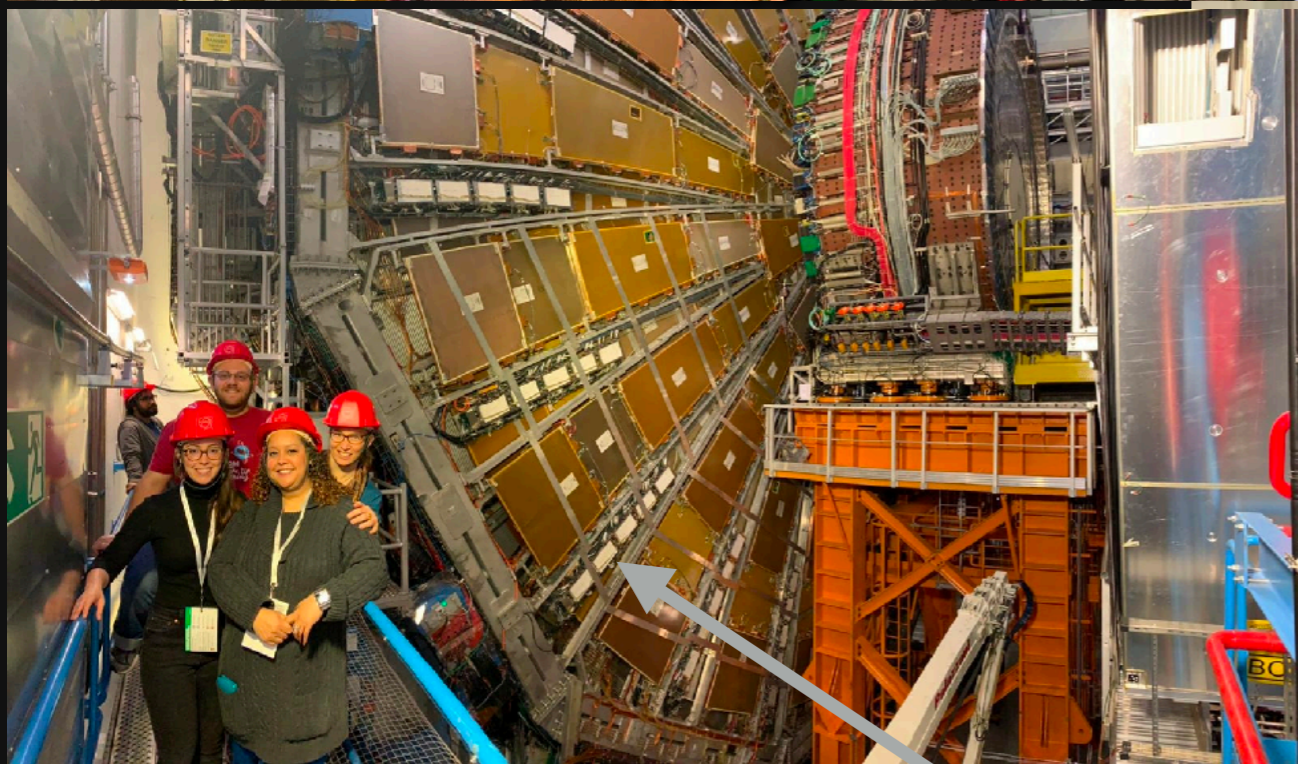
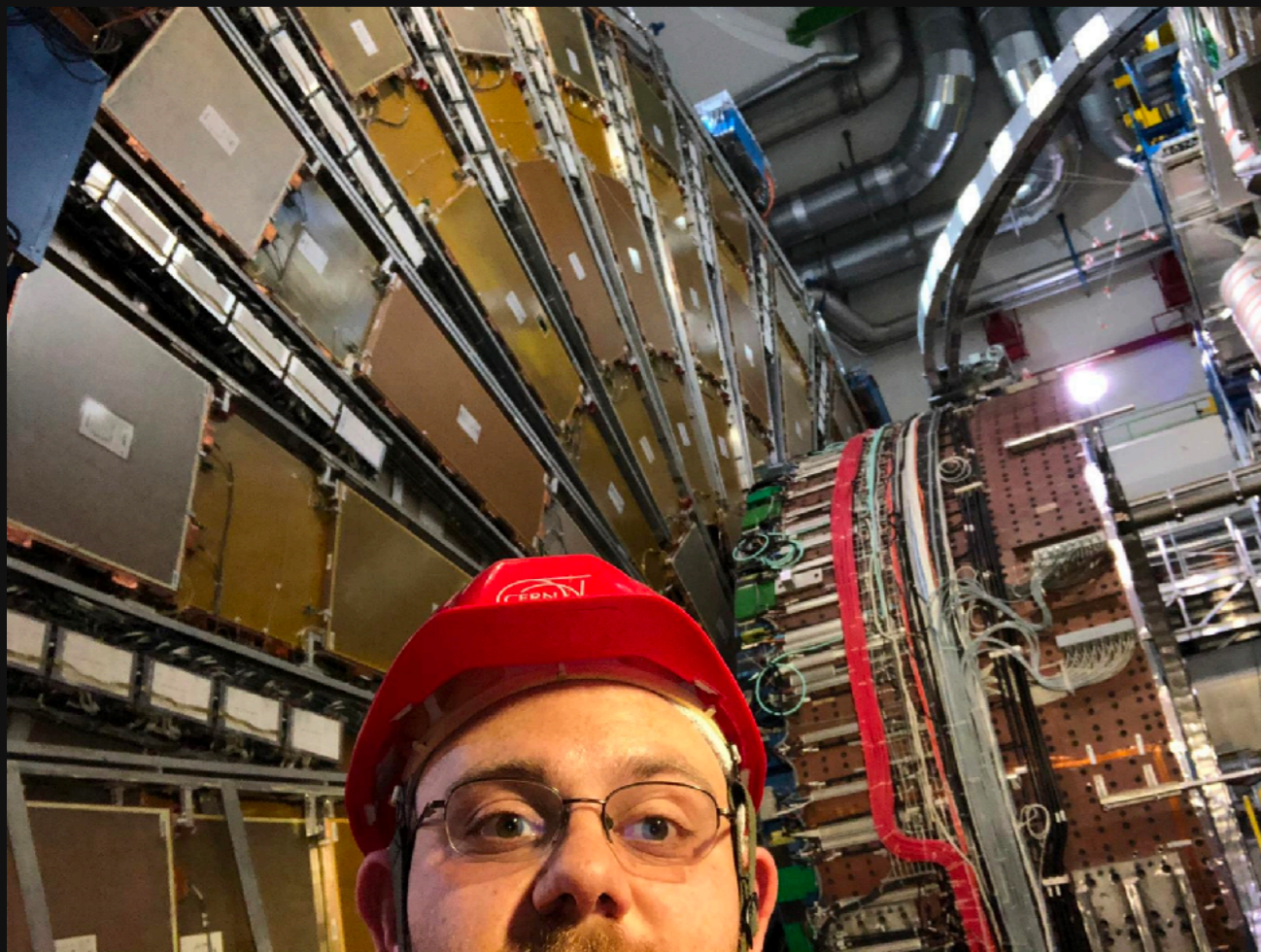
Located at collision Point 1



Stable rock at that depth



LHC

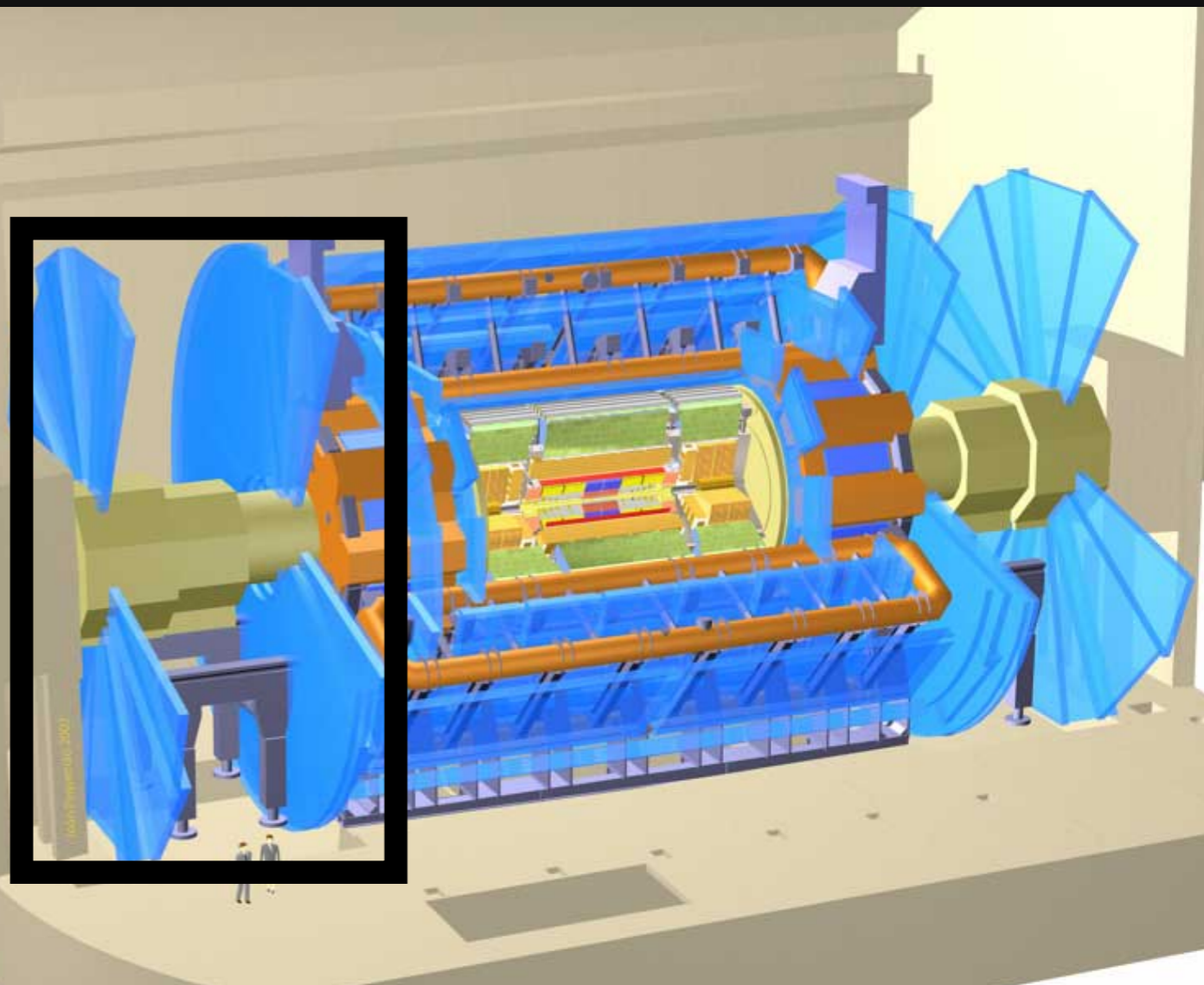


ATLAS

a detector

ATLAS is a large 7000 ton general purpose detector (46m x 25m)

Located at collision Point 1

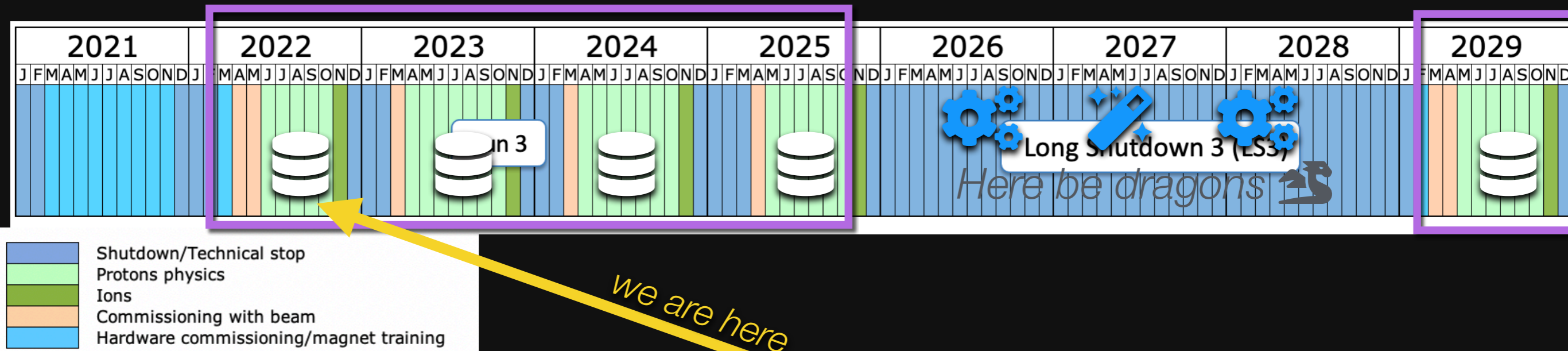


LHC Schedule

No data collection

Run 3

Run 4+



Experiments recently started up, **Run 3 data** 🗄 July 5th 2022!

- ✦ Focus on doing more with what we have
 - ✦ clever techniques to find new physics (SUSY?) in existing data
 - 🚨 **global fits and large scale combinations to determine future directions**
- ✦ **Finalized calibrations** on physics objects (electrons, muons, jets, photons) and pushed **object definitions to lower energies**

No data collection

LHC Schedule

<https://www.youtube.com/watch?v=06kFq1QF5-s>

Run 4+

2021											
J	F	M	A	M	J	J	A	S	O	N	D

2028											
J	F	M	A	M	J	J	A	S	O	N	D

2029											
J	F	M	A	M	J	J	A	S	O	N	D

- Shutdown/Tech
- Protons physics
- Ions
- Commissioning
- Hardware com



Ex

2022!

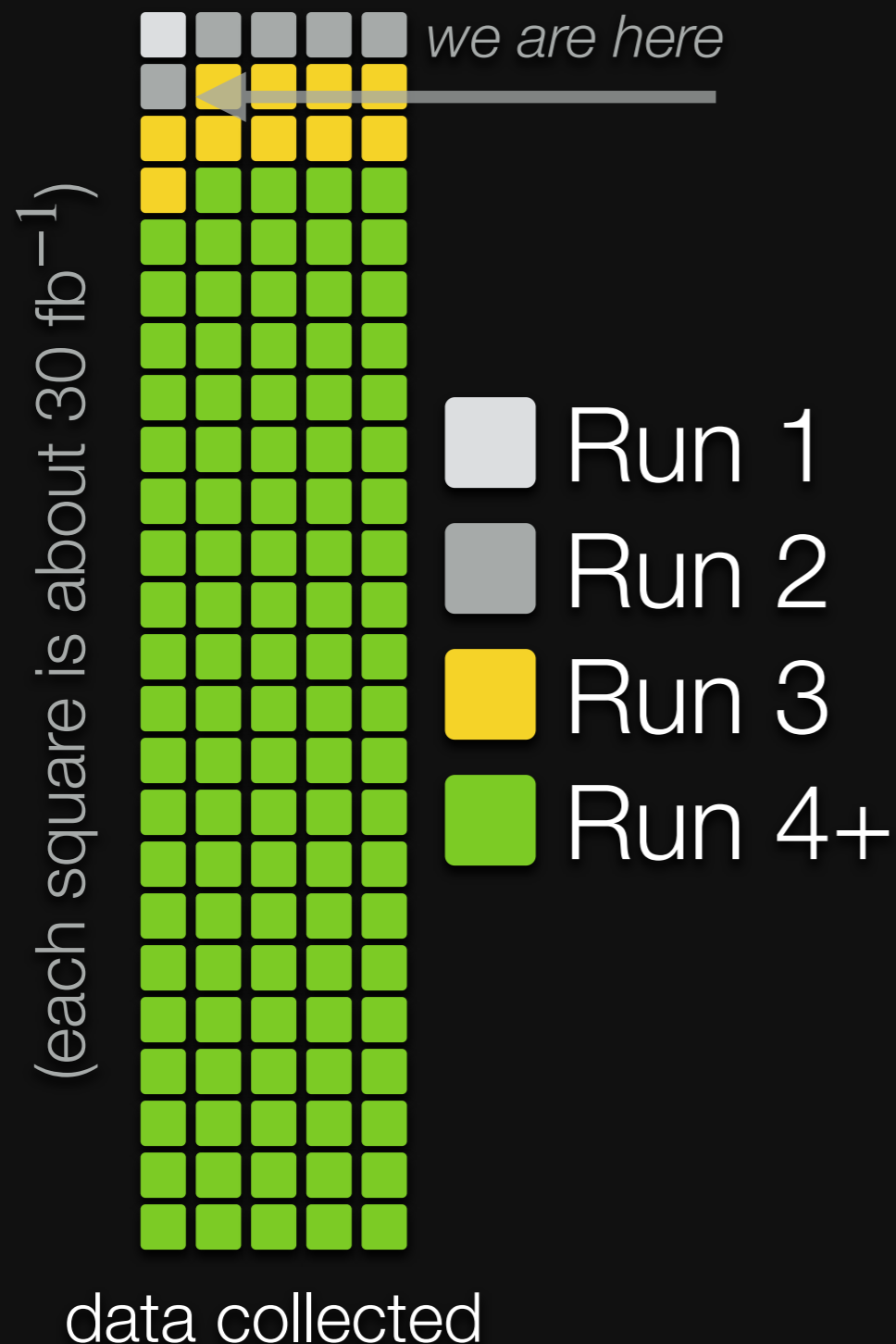
- Focus on
- cleve
- ! glob

ections

- Finalized calibrations on physics objects (electrons, muons, jets, photons) and pushed object definitions to lower energies

The Large Hadron Collider

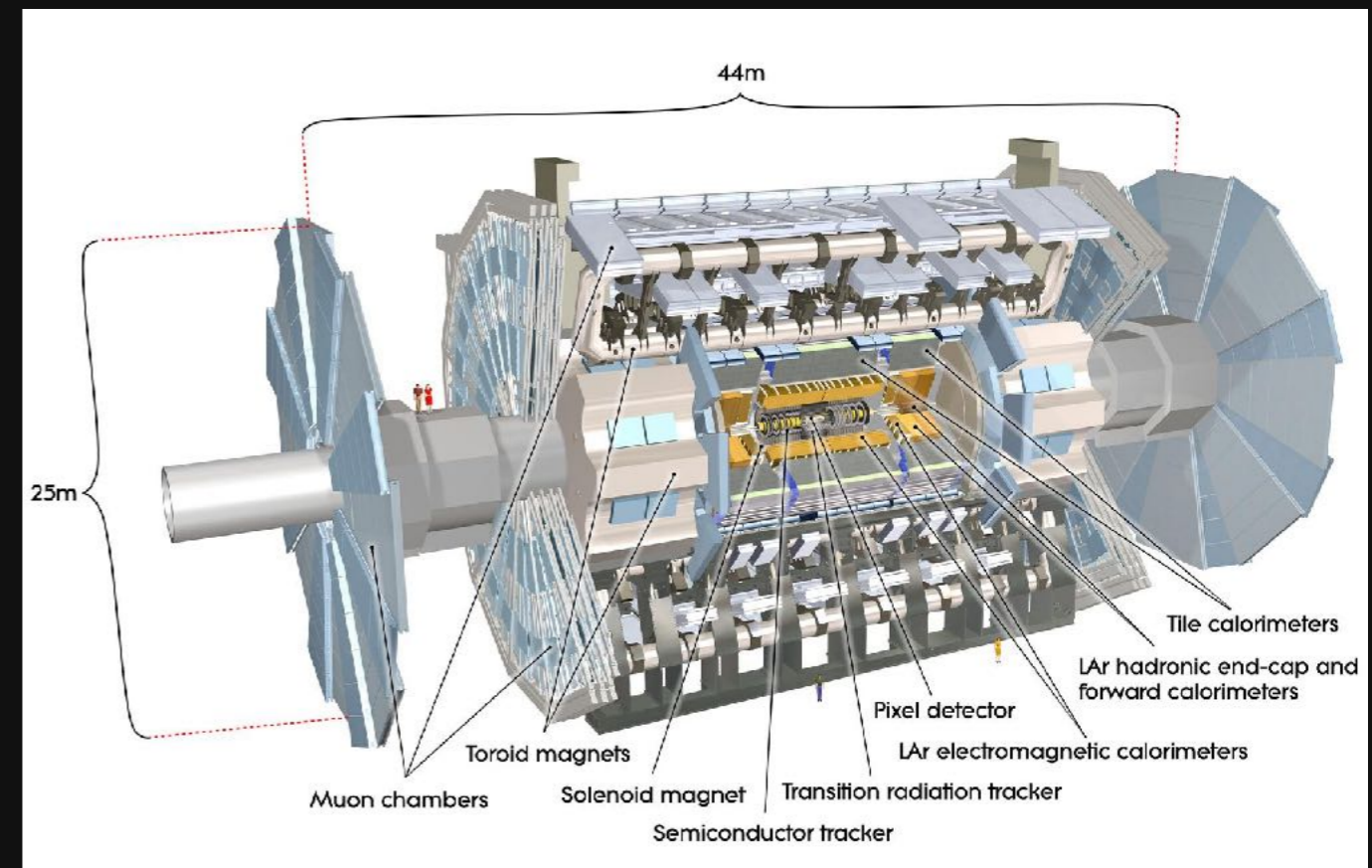
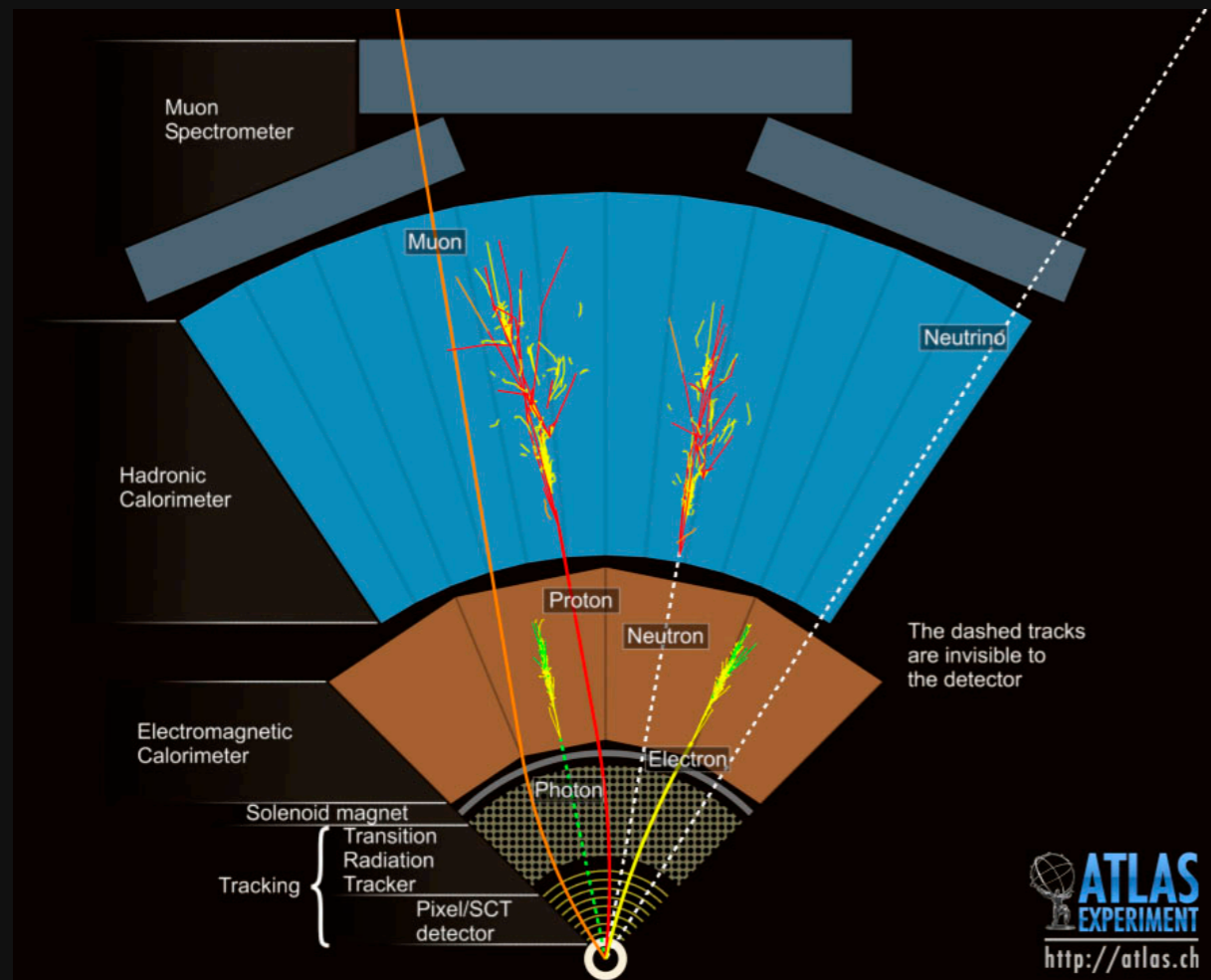
LUMIRDLE



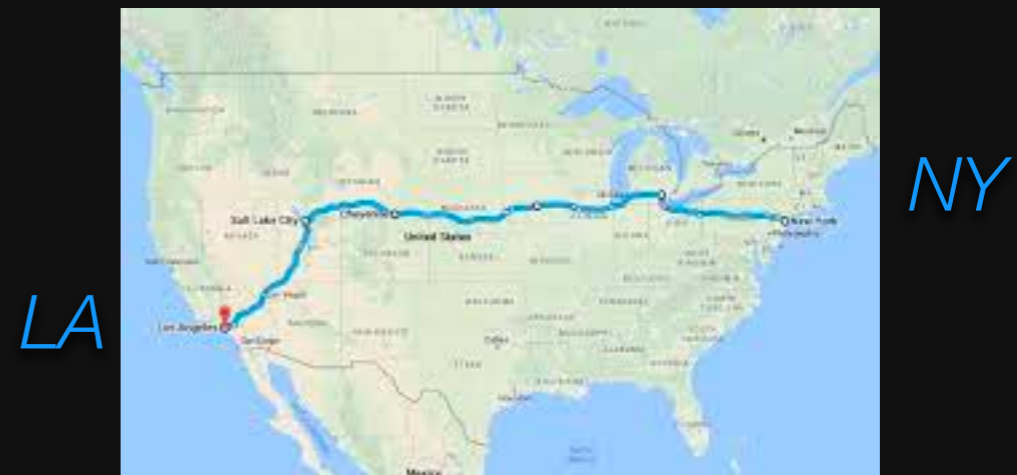
- ✦ **20 more years** of (HL-)LHC physics
- ✦ In 2015-2018: collected 5x more data than in the dataset where we discovered the Higgs
 - ✦ 40 million collisions per second (one every **25 ns**)
 - ✦ 90 petabytes/year of data

The ATLAS Detector

4 main subsystems: inner detector, muon spectrometer, calorimetry, and trigger



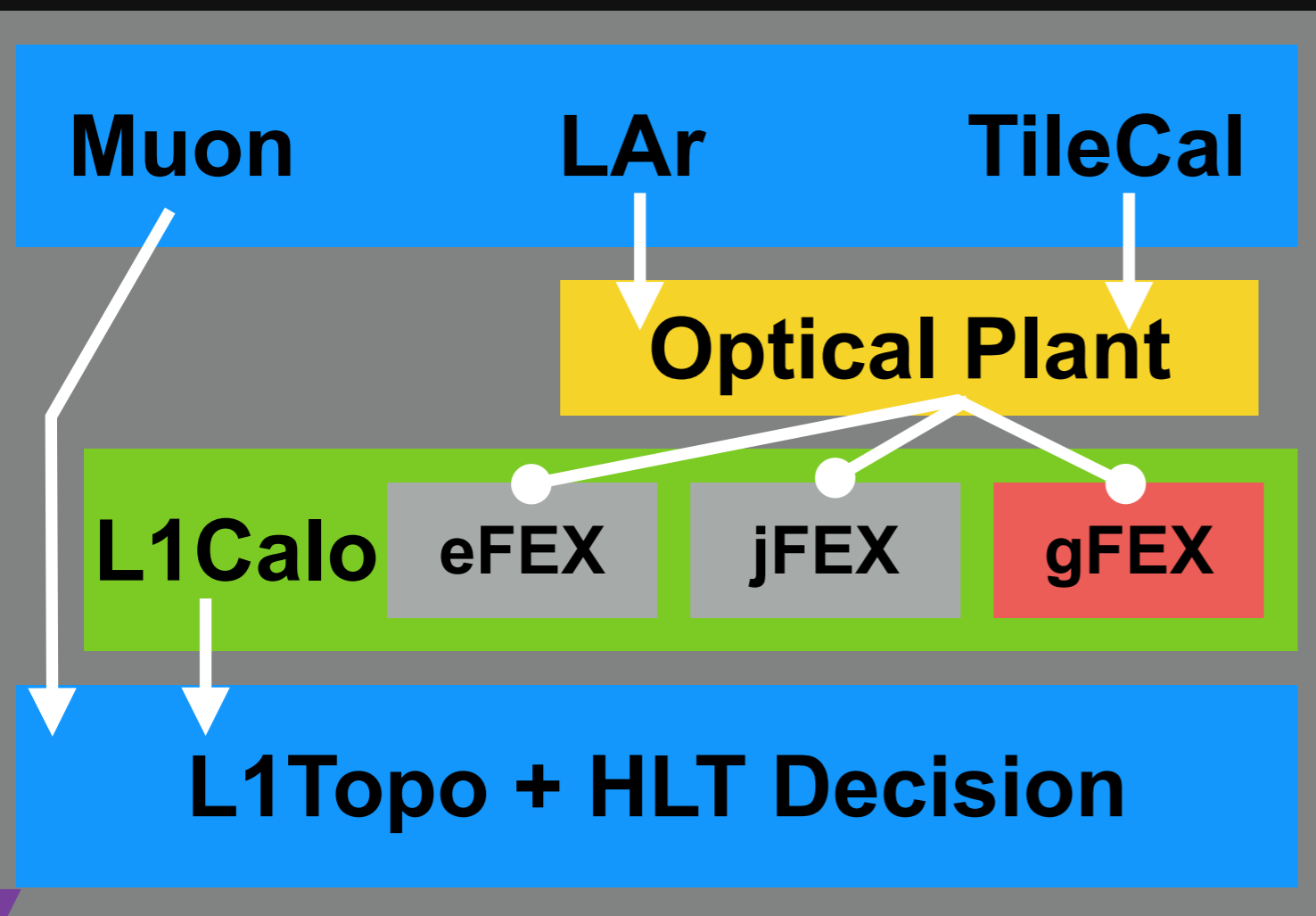
A single complex detector comprising many subsystems. Over 100 million electronic channels and **3000km of cables!**



⚠ One event/bunch-crossing has many simultaneous proton-proton collisions

Calorimetry and Trigger

40MHz Largely “junk” events

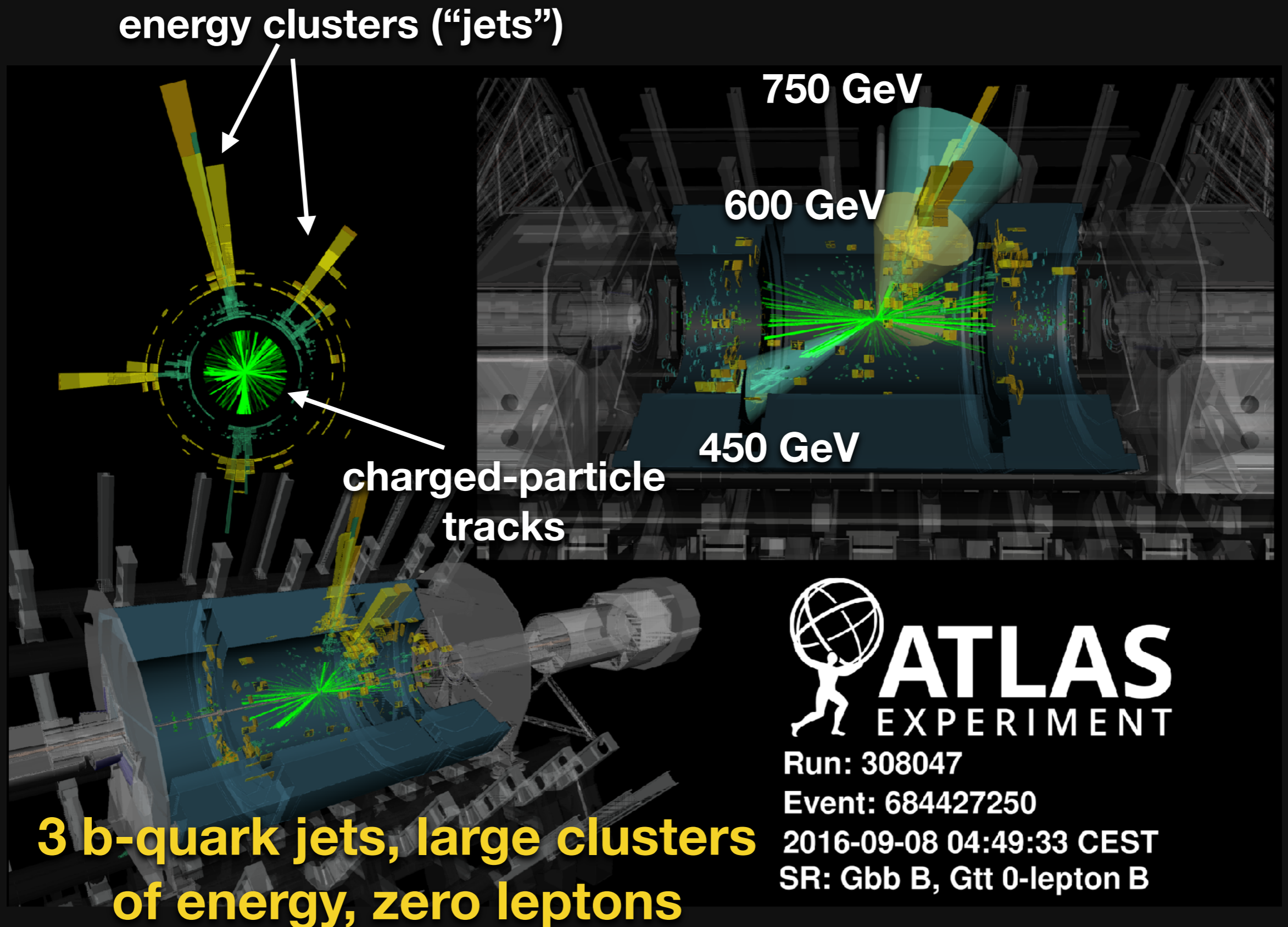


- The trigger system uses data from the calorimeters
- Bunches of protons collide 40 million times per second
 - Can only save ~1000 events per second to disk
- **Goal:** retain efficiency of processes sought for in ATLAS
 - Need a lot of smart rejection
 - Need it fast and performant
 - Keep rates under control

1kHz Largely “interesting” events

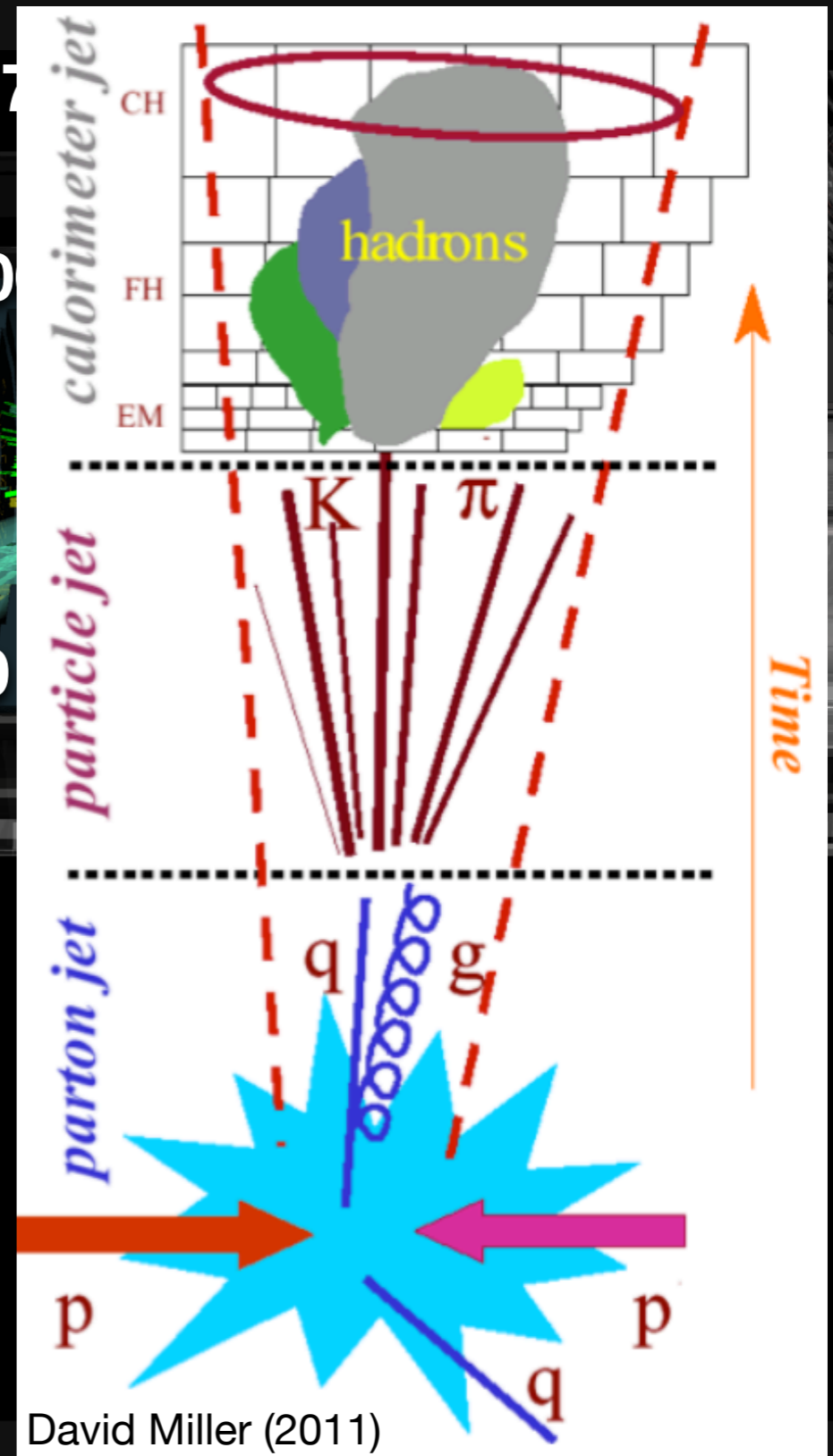
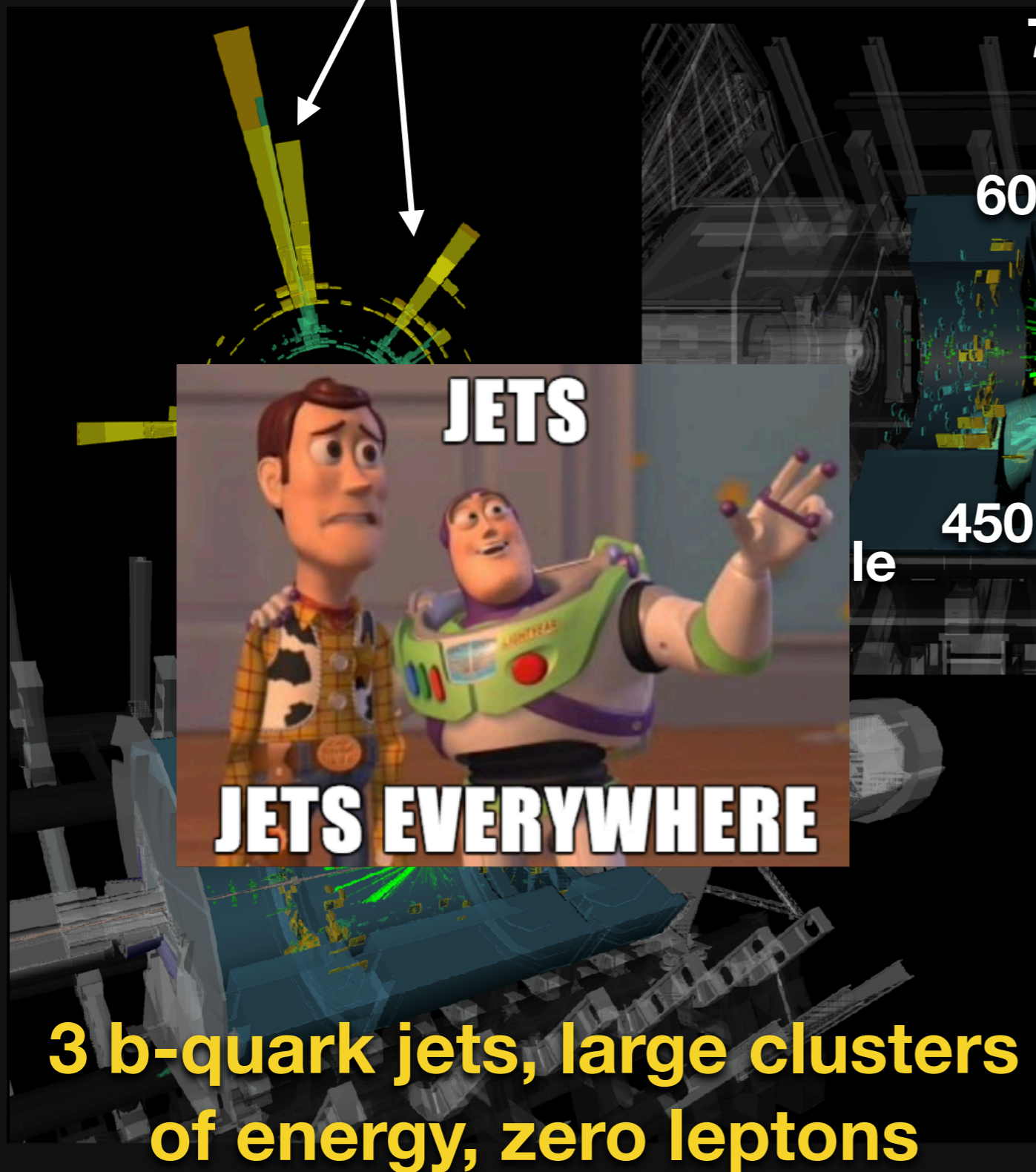
❓ Make a selection to **keep** 📁 or **throw away** 🗑 data forever

A fully reconstructed event



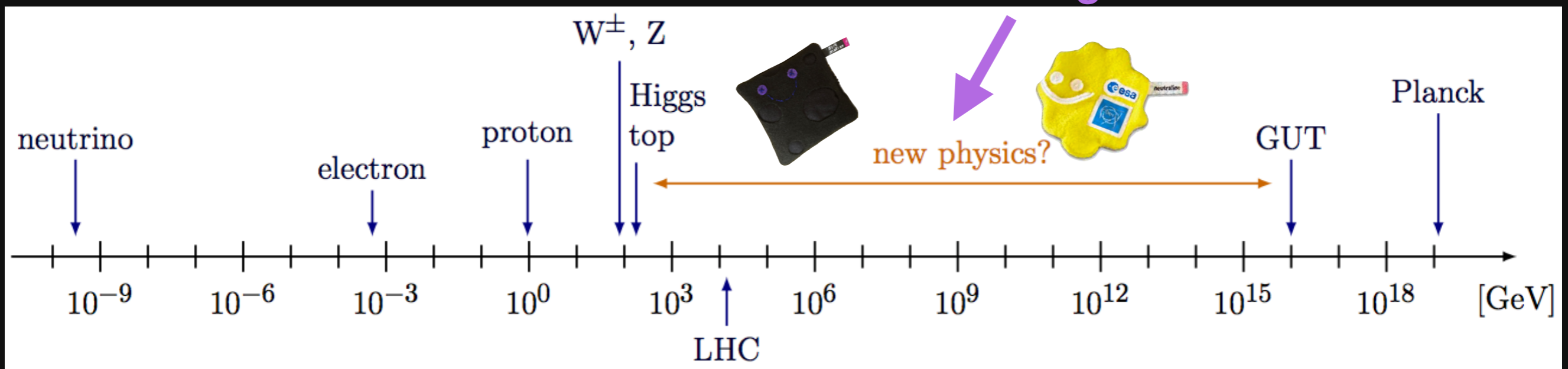
A fully reconstructed event

energy clusters ("jets")



Searching for SUSY

Where is SUSY hiding?



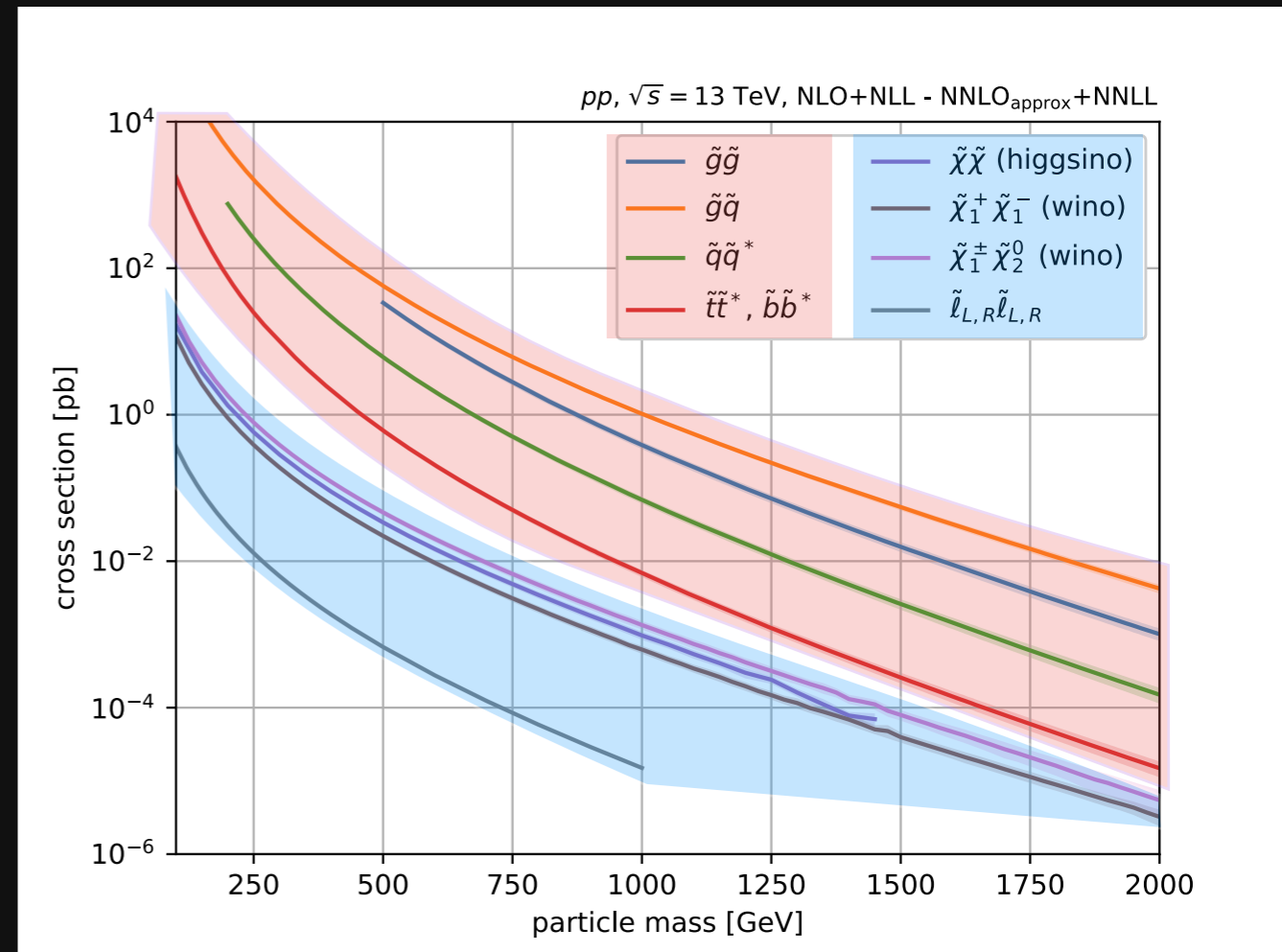
“SUSY is just around the corner.”
— Carlos Wagner



! “Naturalness” motivates light gluinos and stops

Maybe gluinos/squarks?

- ✦ LHC at 13 TeV well-motivated to search for SUSY (*some searches are possible for the first time!*)
- ✦ **Gluinos and squarks**, with large color coupling, have highest cross-section in SUSY
- ✦ **Electroweak-produced** sparticles are subdominant



Search for strongly-produced sparticles!

(electroweak states may be first detected if high mass limits on strong production)

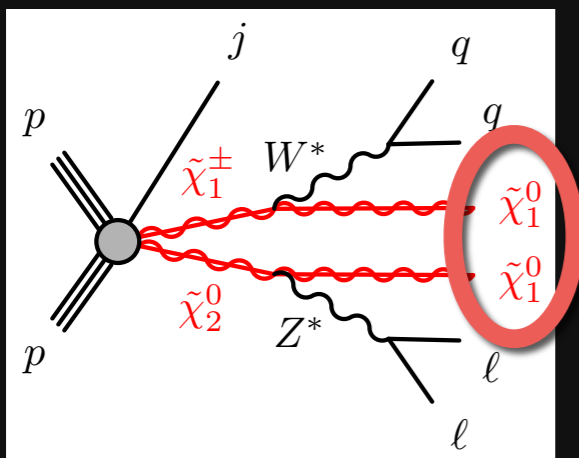
What does SUSY look like?

Lots of kinematics! Grouped into three categories:

What does SUSY look like?

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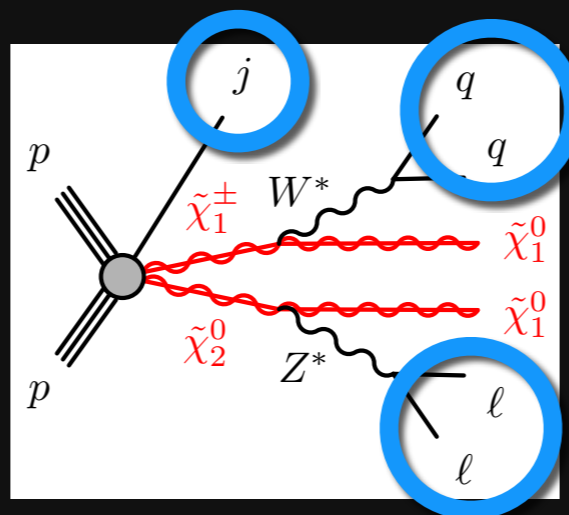
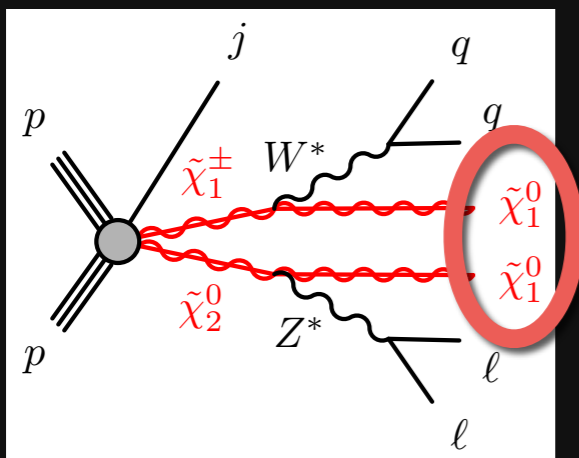
- **Missing momentum-type**: sensitive to the properties of the invisible states
 - ❓ **how many neutralinos in the event?** (*RPC: LSP escapes detection*)



What does SUSY look like?

Lots of kinematics! Grouped into three categories:

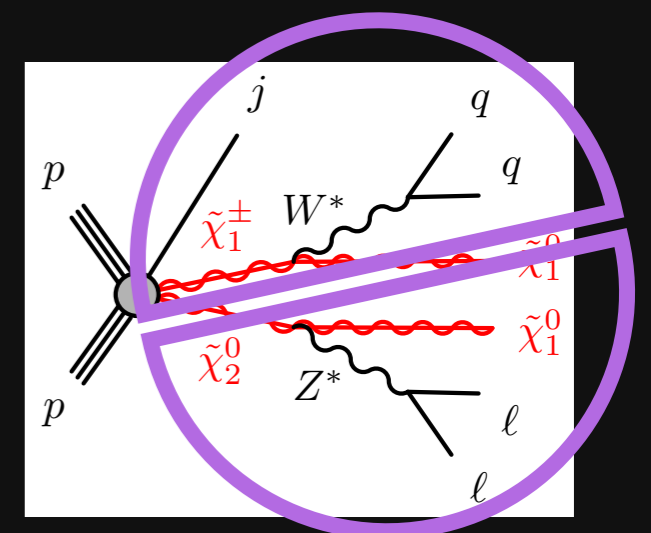
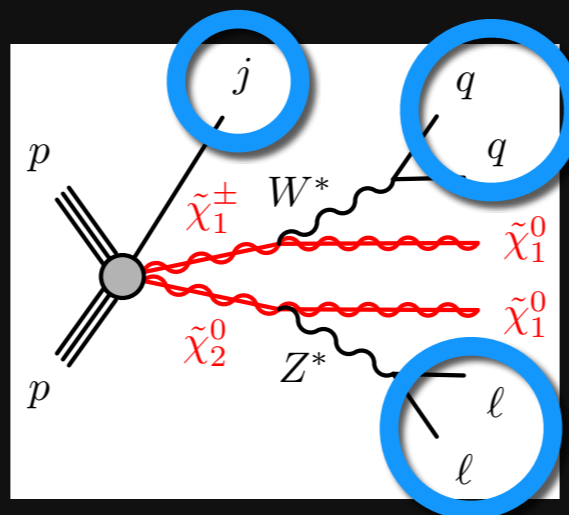
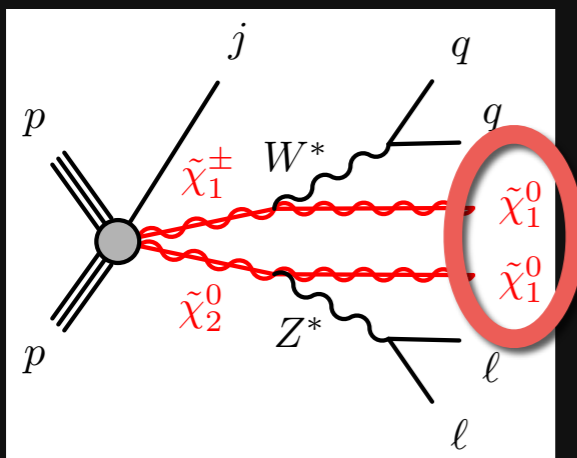
- ✦ **Missing momentum-type**: sensitive to the properties of the invisible states
 - ❓ **how many neutralinos in the event?** (*RPC: LSP escapes detection*)
- ✦ **Energy scale-type**: sensitive to the overall energy scale of the event
 - ❓ **how much energy in the event?** (*SUSY can reach high mass scales*)



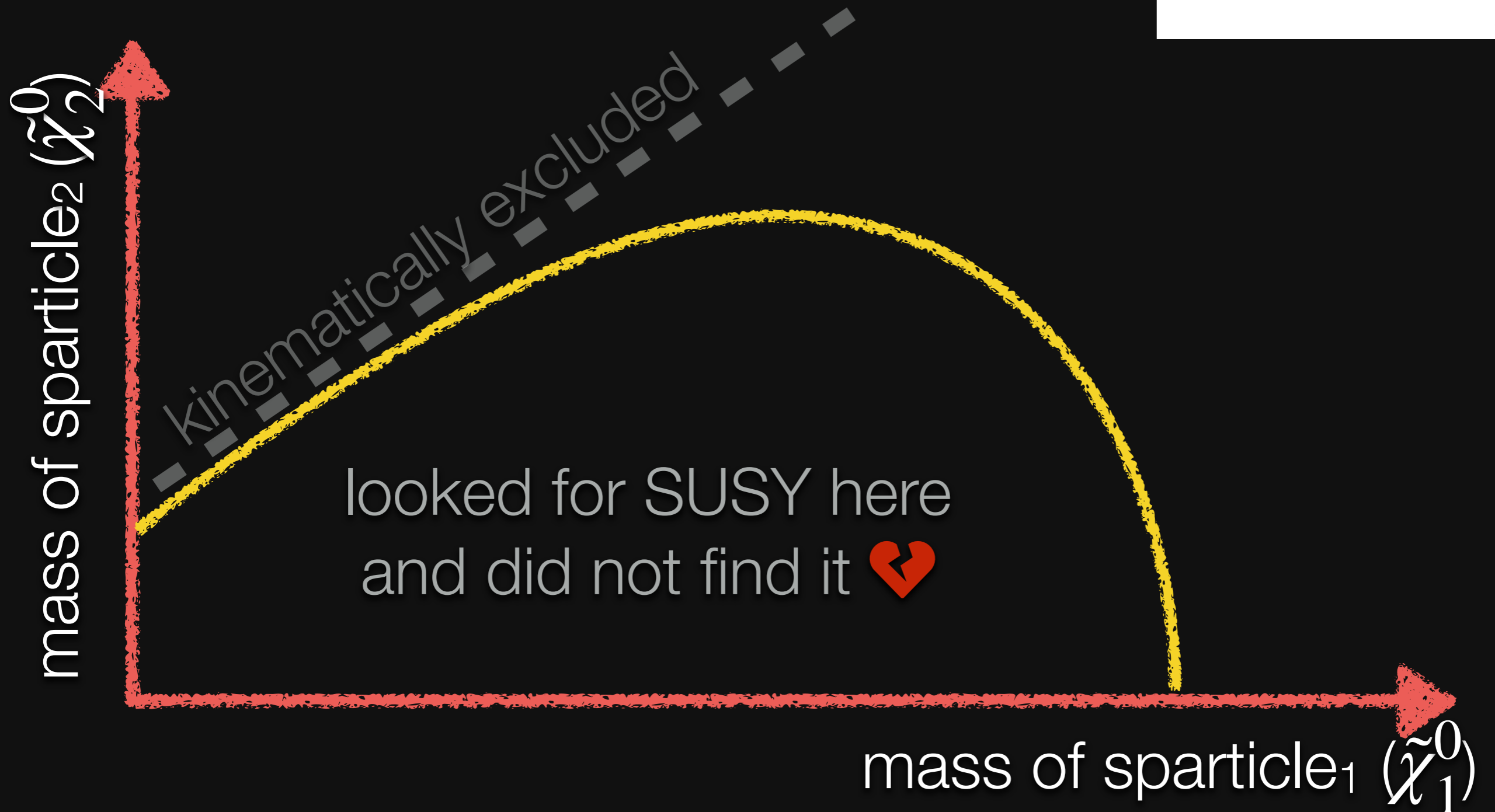
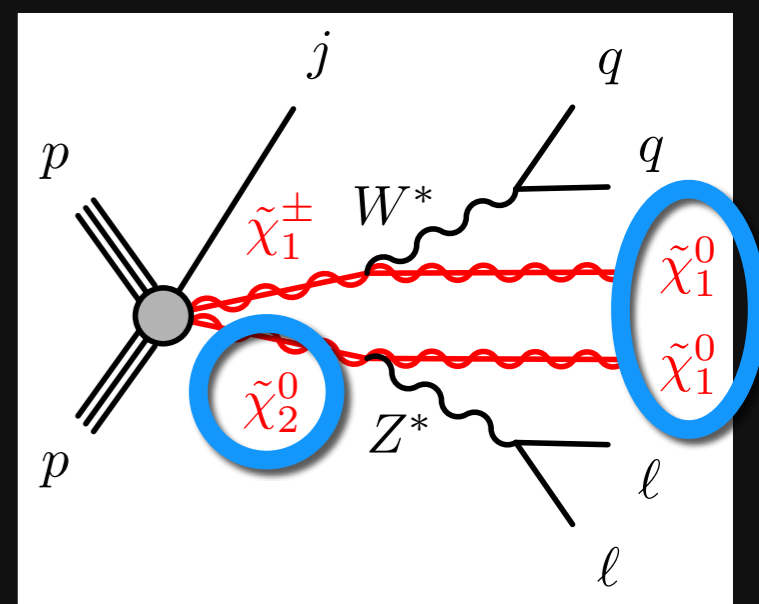
What does SUSY look like?

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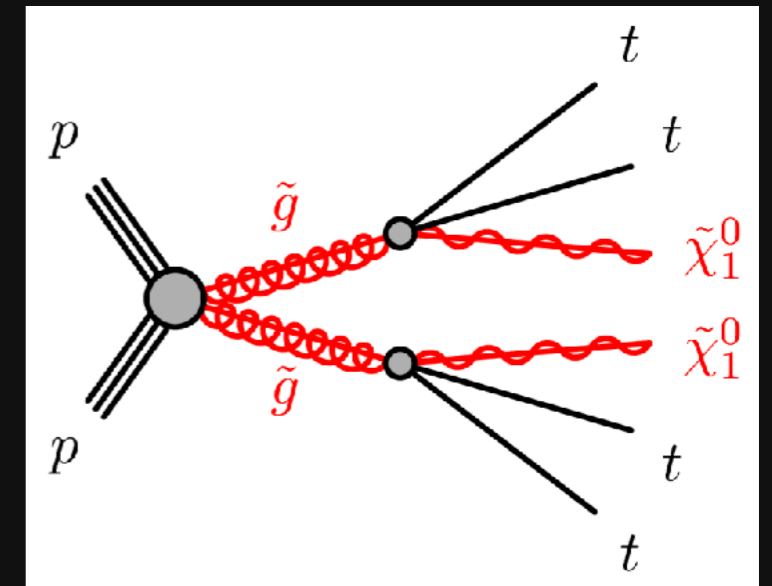
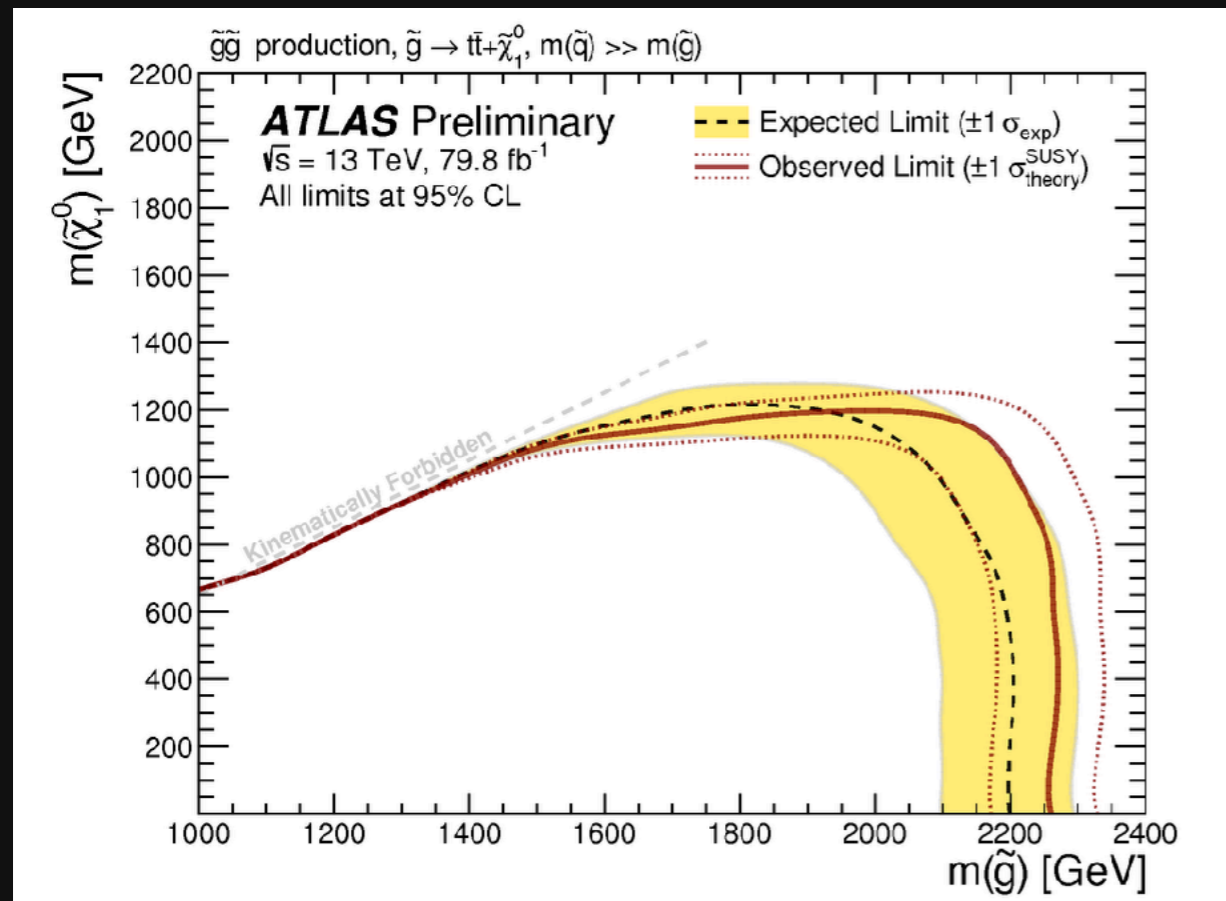
- **Missing momentum-type**: sensitive to the properties of the invisible states
 - ❓ how many neutralinos in the event? (*RPC: LSP escapes detection*)
- **Energy scale-type**: sensitive to the overall energy scale of the event
 - ❓ how much energy in the event? (*SUSY can reach high mass scales*)
- **Energy structure-type**: sensitive to the structure of the visible energy
 - ❓ how is the energy of the decay partitioned across the final state visible/invisible objects? (*e.g. decay angle between LSP and jets*)



Interpreting Results



(My thesis analysis)



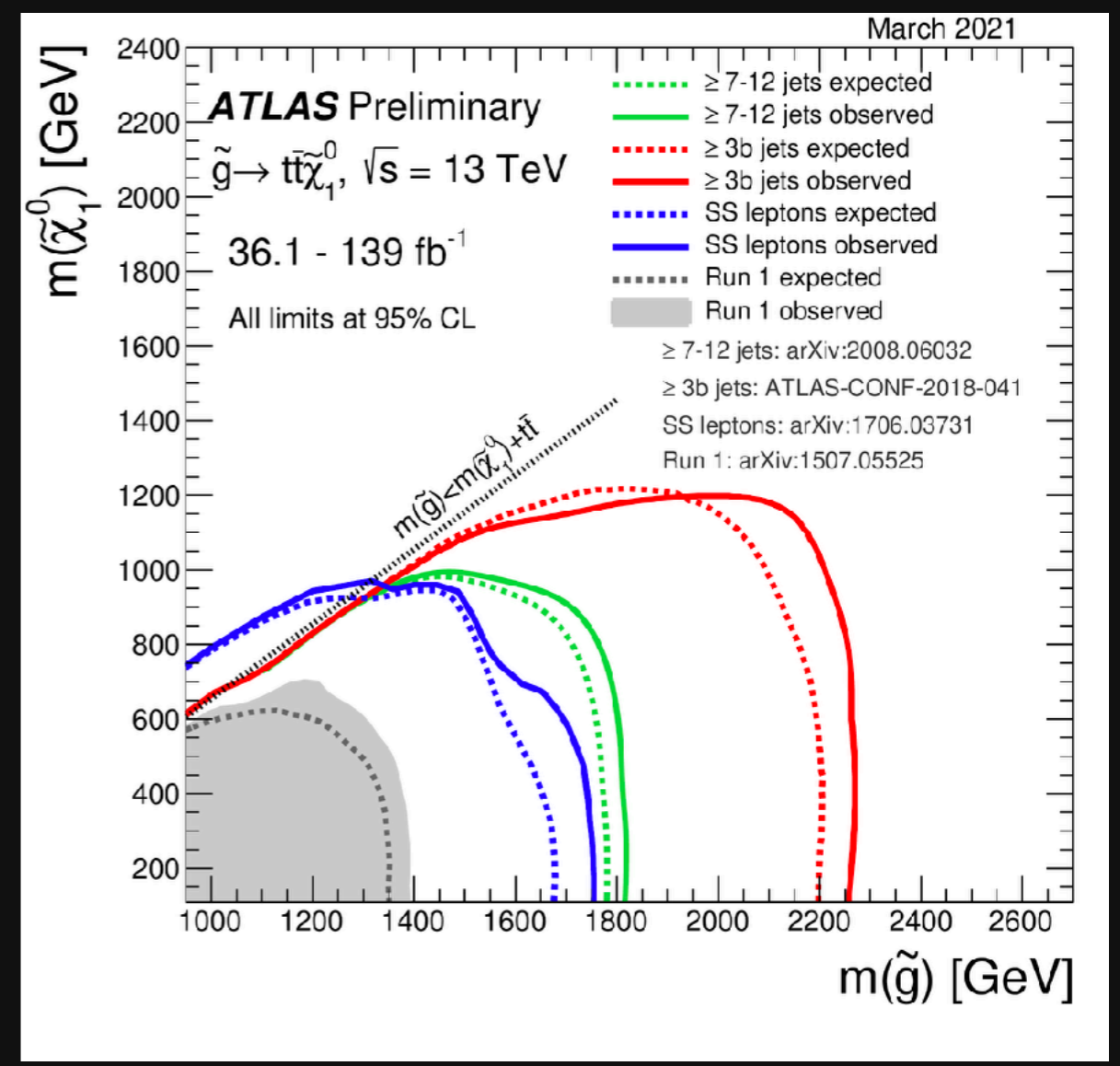
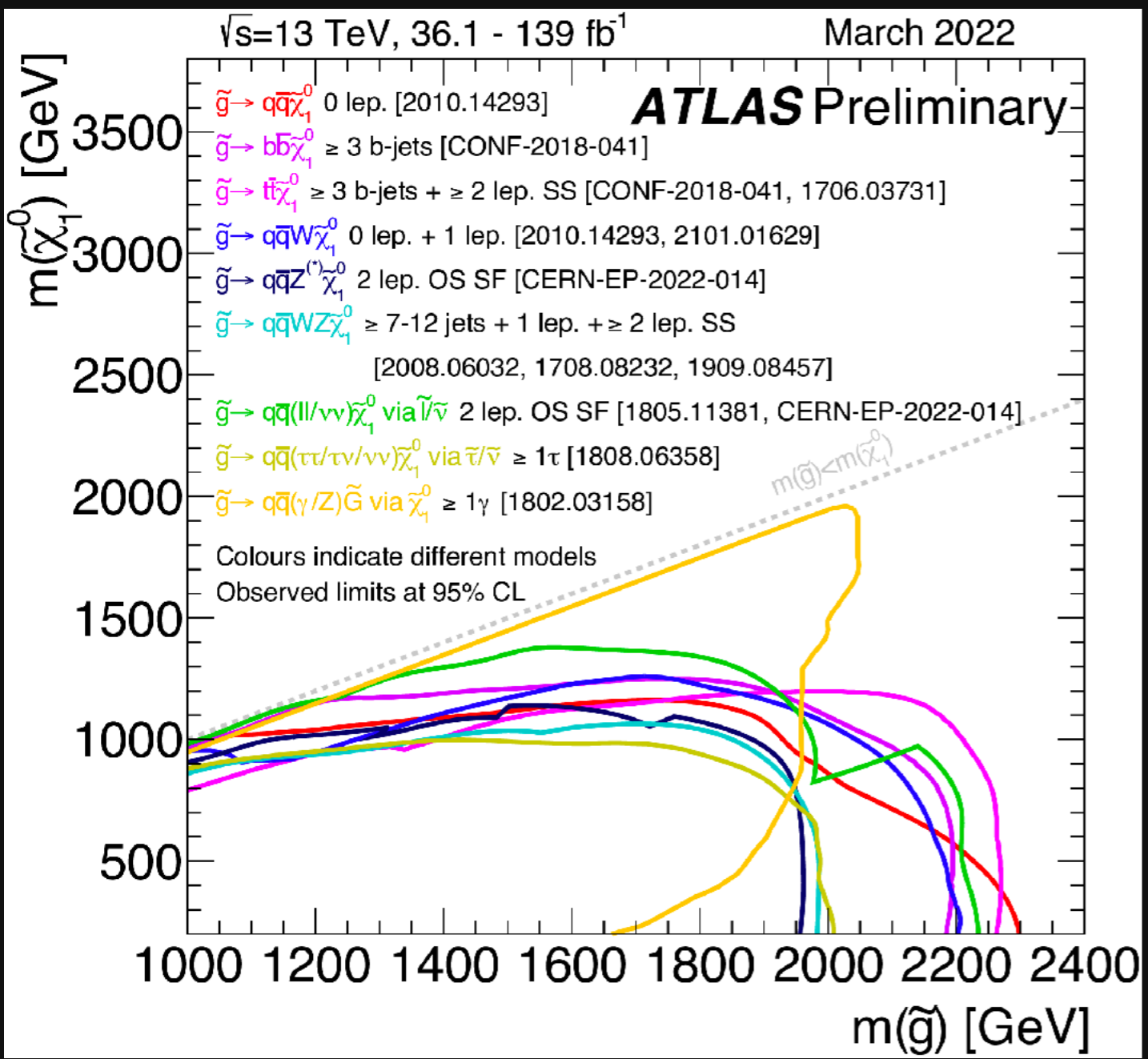
gluino-mediated stop pair production

- ✦ No SUSY found here. (but I finished my PhD!)

Limit on gluino mass
~2.20 TeV @ 95% CL

Search for strongly-produced sparticles!

And so far...



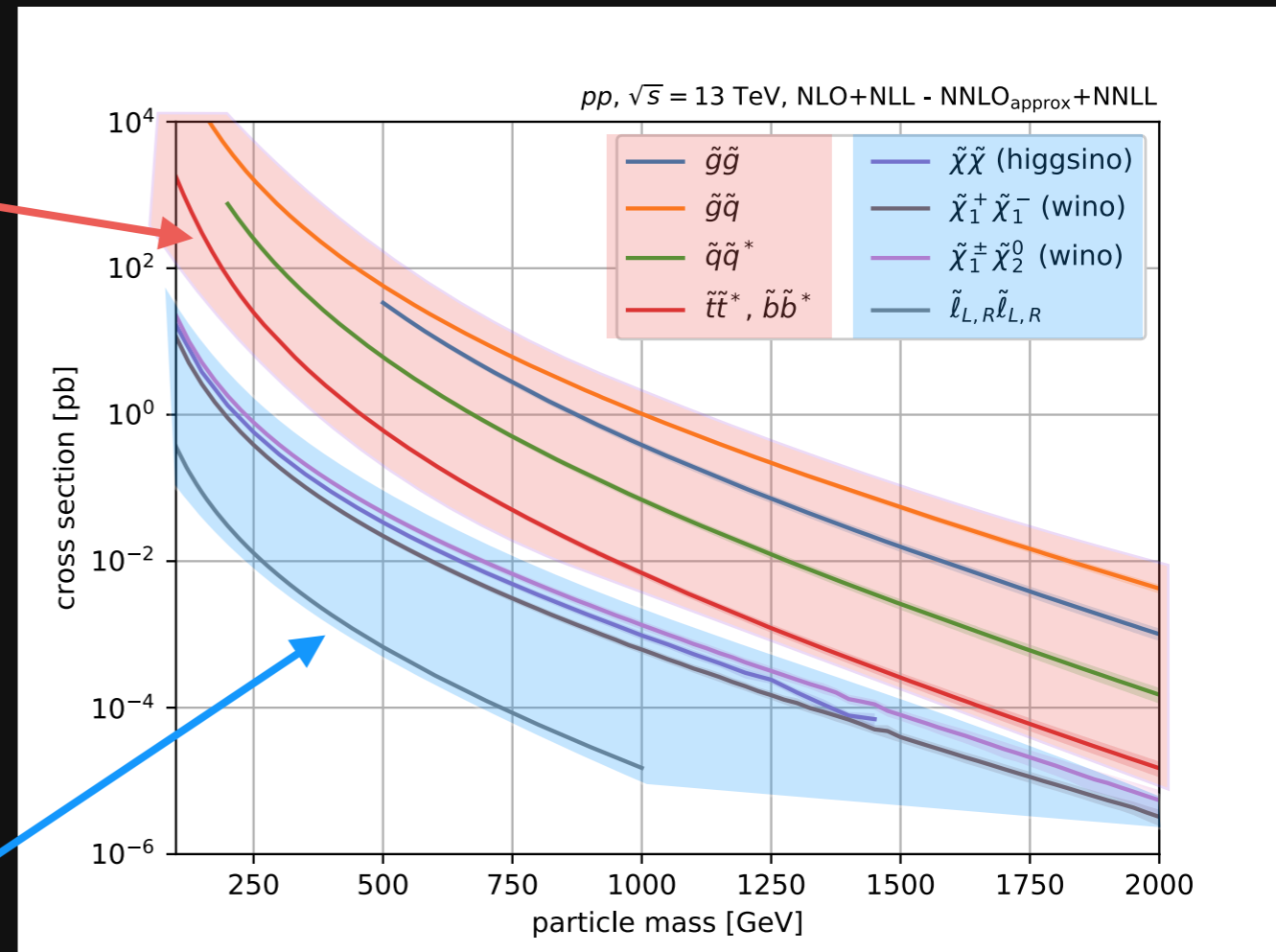
- ✦ No SUSY found here.
- ✦ Not really “light” anymore!
- ✦ Maybe we should look **somewhere** else!

each contour represents different interpretations of a SUSY model

Search for strongly-produced electroweakly-produced sparticles?

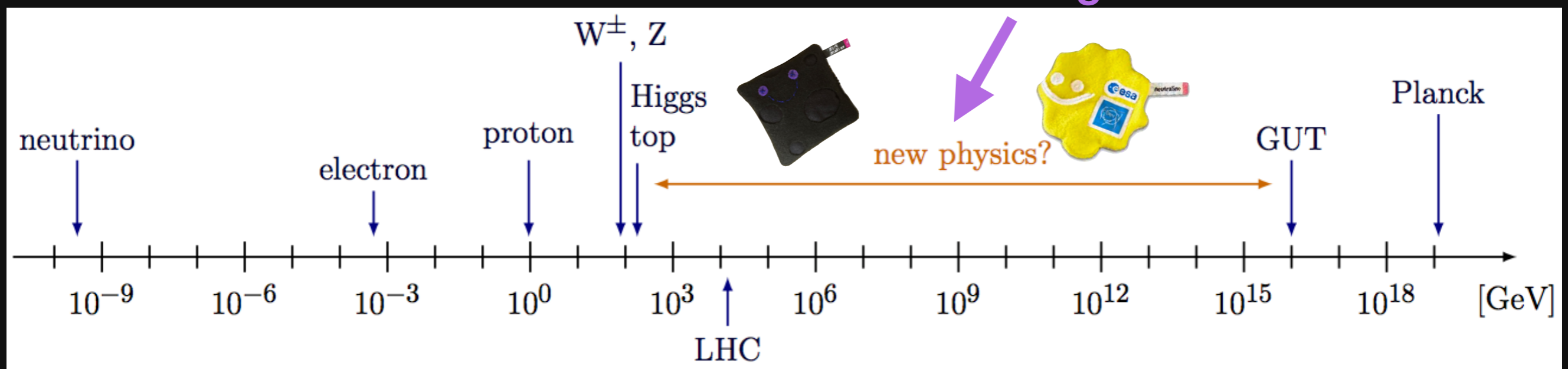
Questions so far

- ✦ Reaching the **energy limits** of our current machine searching for SUSY produced through strong interactions
- ✦ Bigger dataset → start hunting **rarer processes** to produce SUSY through electroweak interactions



Searching for EWK SUSY

Where is SUSY hiding?

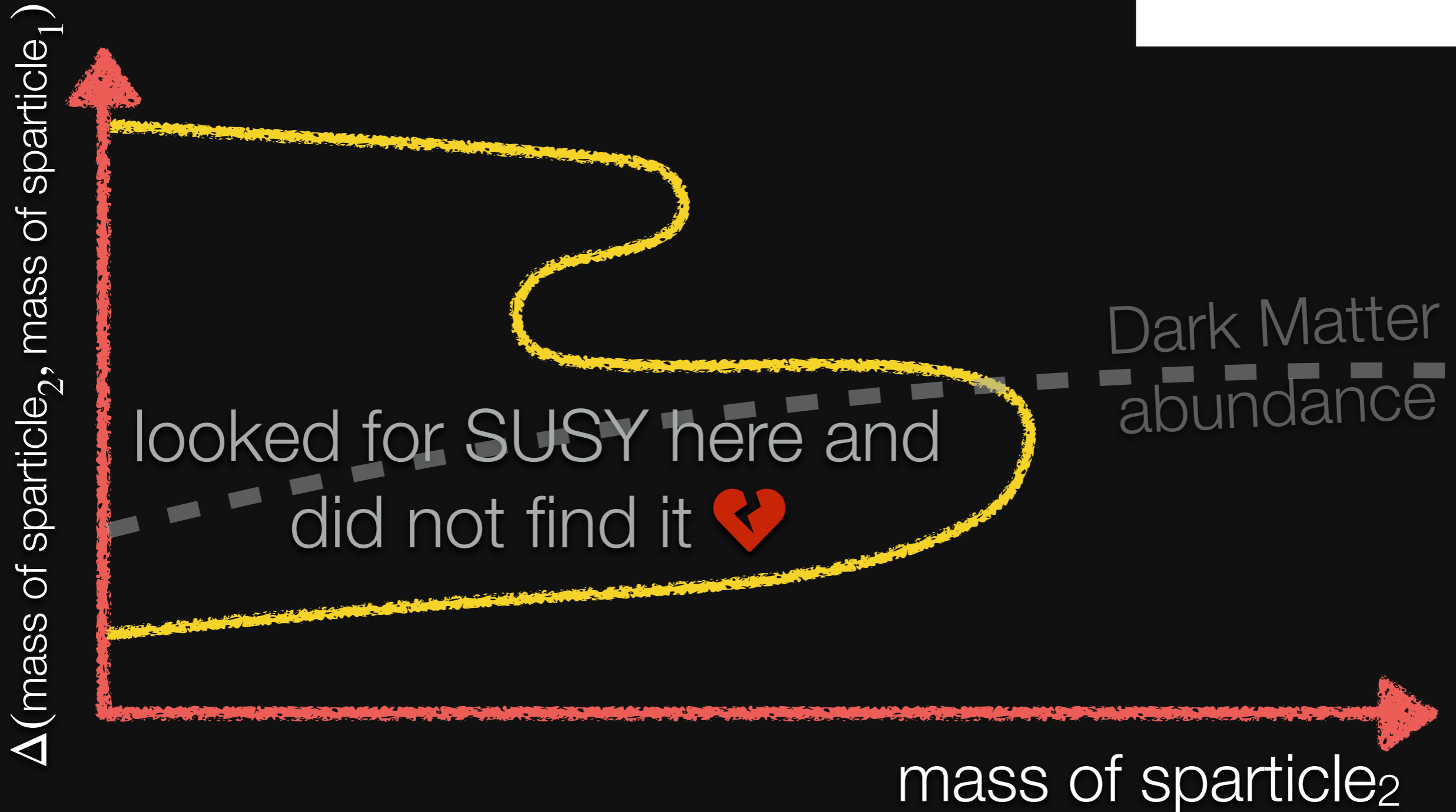
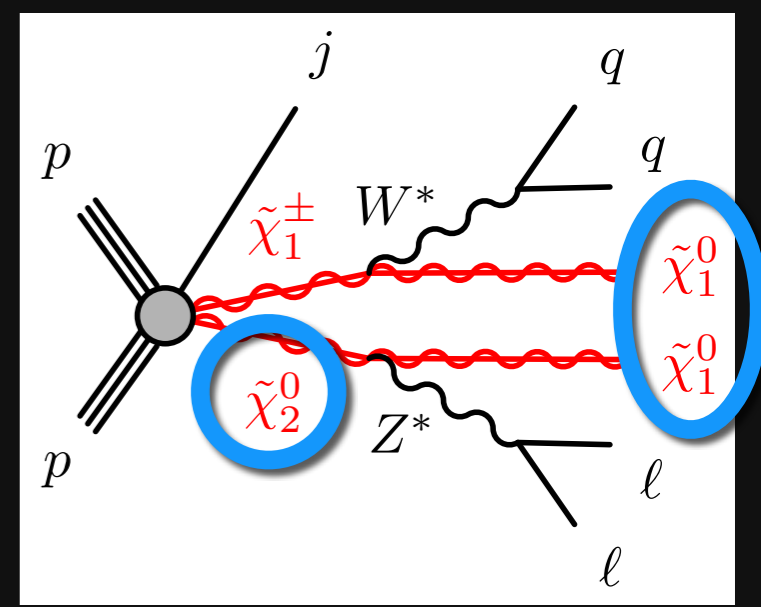


“SUSY is just around another corner.”

— Giordon Stark

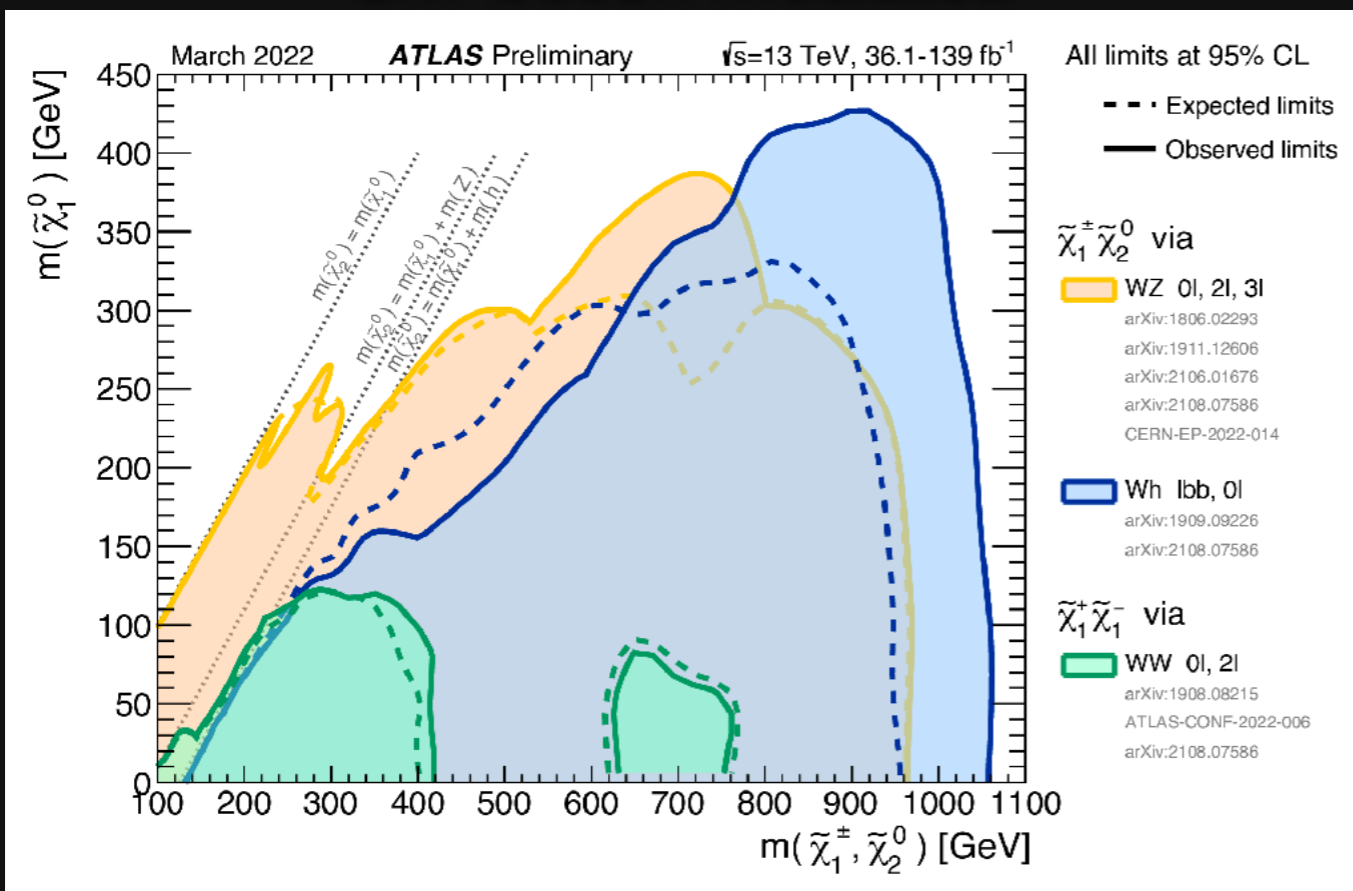


Interpreting Δm

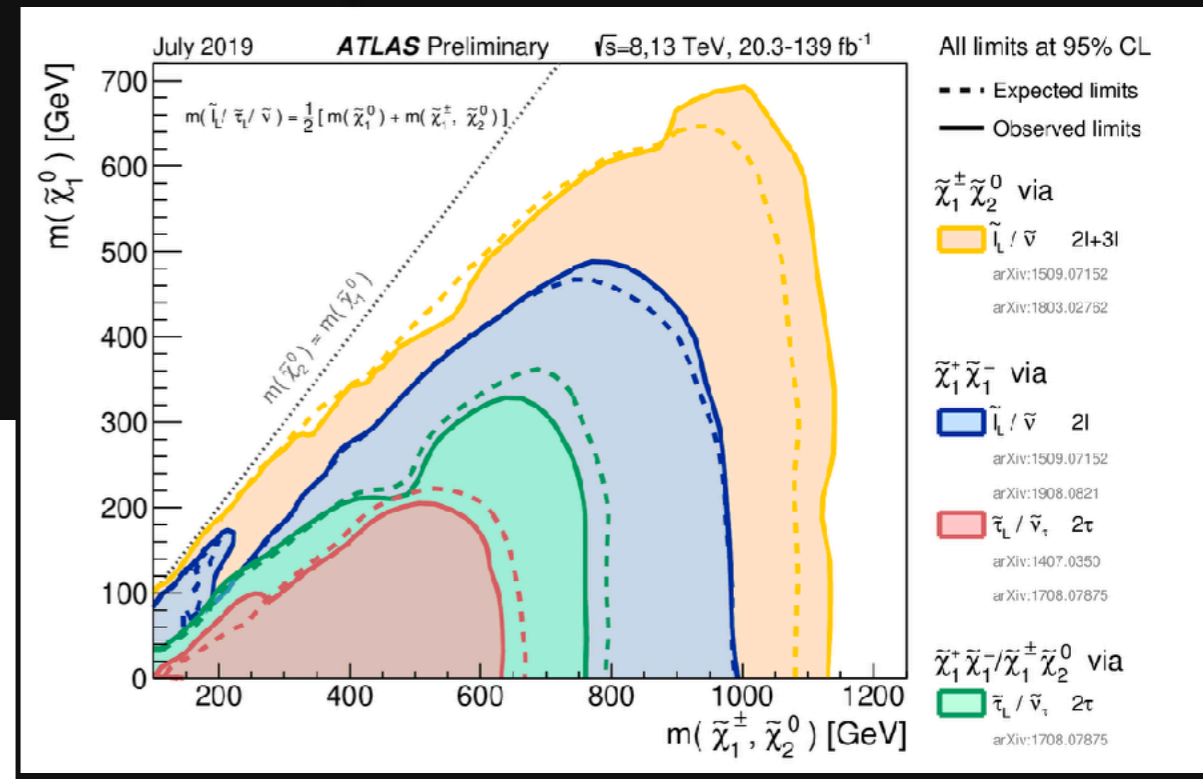


And so far...

SM-boson mediated



lepton mediated



$m(\tilde{\chi}_1^0)$

$$m(\tilde{\chi}_1^\pm, \tilde{\chi}_2^0)$$

- ✦ No SUSY found here.
- ✦ Maybe we should **combine** results!

each contour represents different interpretations of a SUSY model

What is our best understanding of the universe?

Questions so far

- ✦ ~~“Naturalness” isn’t that well-motivated here, so perhaps SUSY is possibly electroweakly-produced?~~
- ✦ 🔍 Keep searching using (upcoming) Run 3+ data
- ✦ How do we make sure that our analysis results are still **interpretable** with new phenomenology today?
- ✦ How do we **combine** different analysis results to constrain the allowed SUSY models?

Statistical Techniques

How do experimentalists count?

not a real
experimentalist

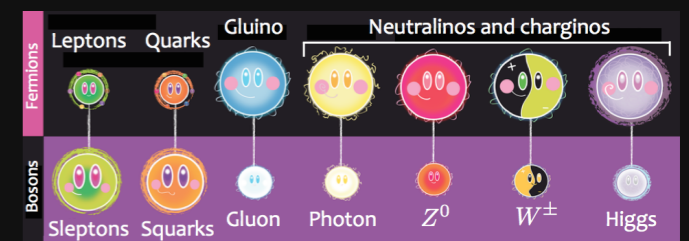
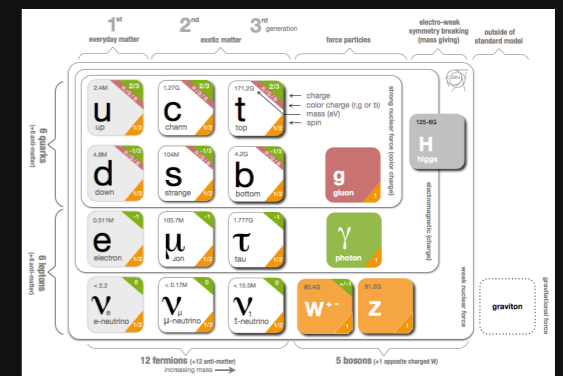
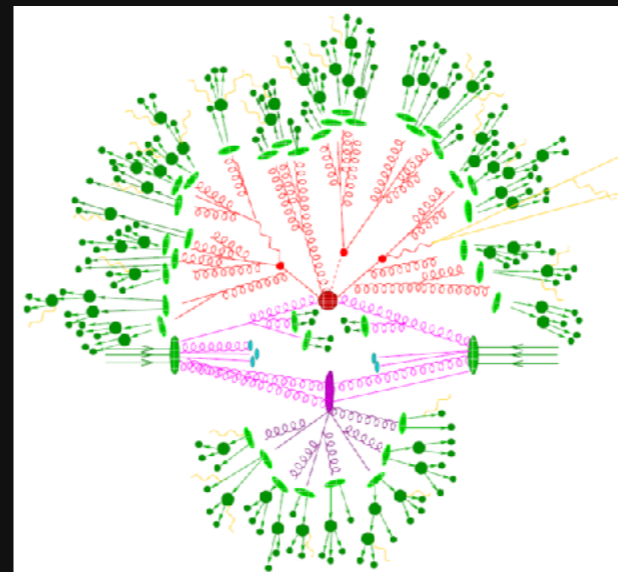
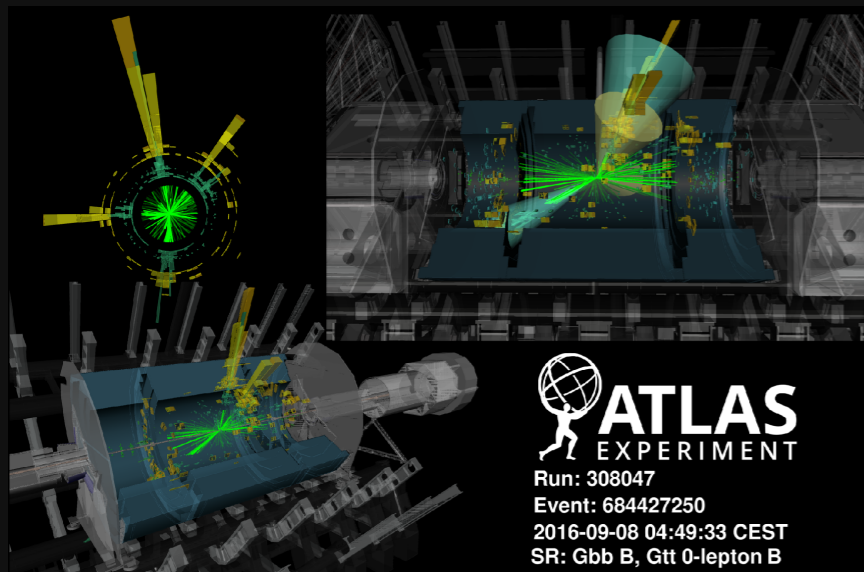


“Do you know why they call me the Count? Because I love to count! Ah-hah-hah!” — The Count (of Sesame Street)

The Big Picture

observations

model
(SM + SUSY)



$p(\text{data} = \text{observed})$

$p(\text{theory})$

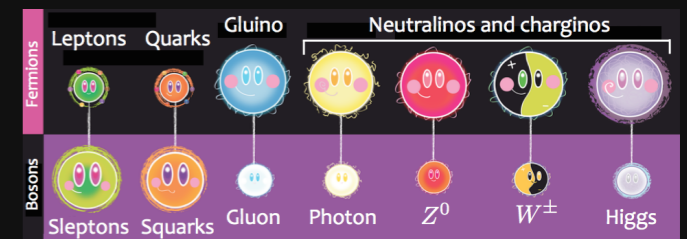
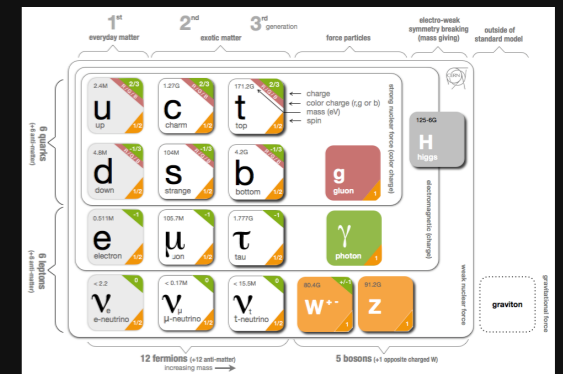
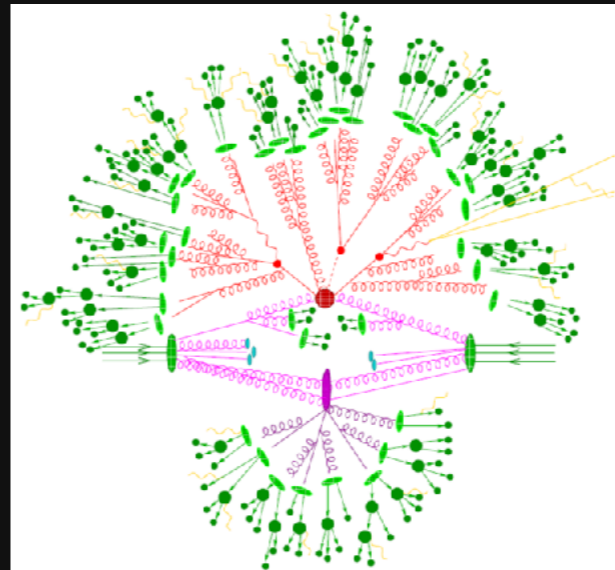
! A **likelihood function** encodes everything we know about the **detector**, the **theory**, and the **data**

! How good is our theory?

The Big Picture (I)

- ✦ Rely on our best **understanding of the theory**
 - ✦ Standard Model / Beyond the Standard Model (e.g. QED, QCD, MSSM)
 - ✦ Matrix Elements
 - ✦ Parton Distribution Functions
 - ✦ Finite Order in Perturbative Calculations (e.g. NLO, NNLO, etc...)
 - ✦ Parton Showering and Hadronization
- ✦ ... and best **simulation of our detector**
 - ✦ Material Interactions
 - ✦ In-time and out-of-time pile-up
 - ✦ Calorimeter efficiency
 - ✦ Tracking efficiency
 - ✦ Magnetic field mapping
 - ✦ Beamspot origin
 - ✦ etc...

model
(SM + SUSY)

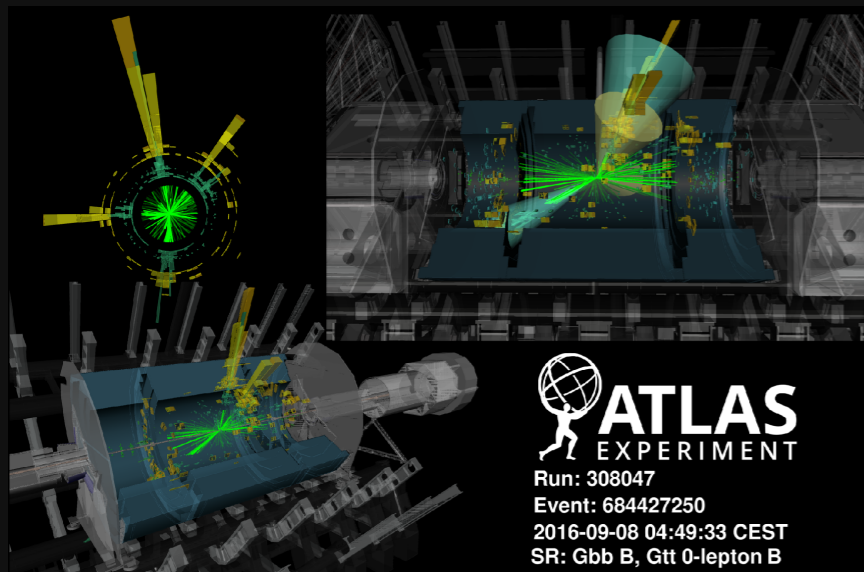


$p(\text{theory})$

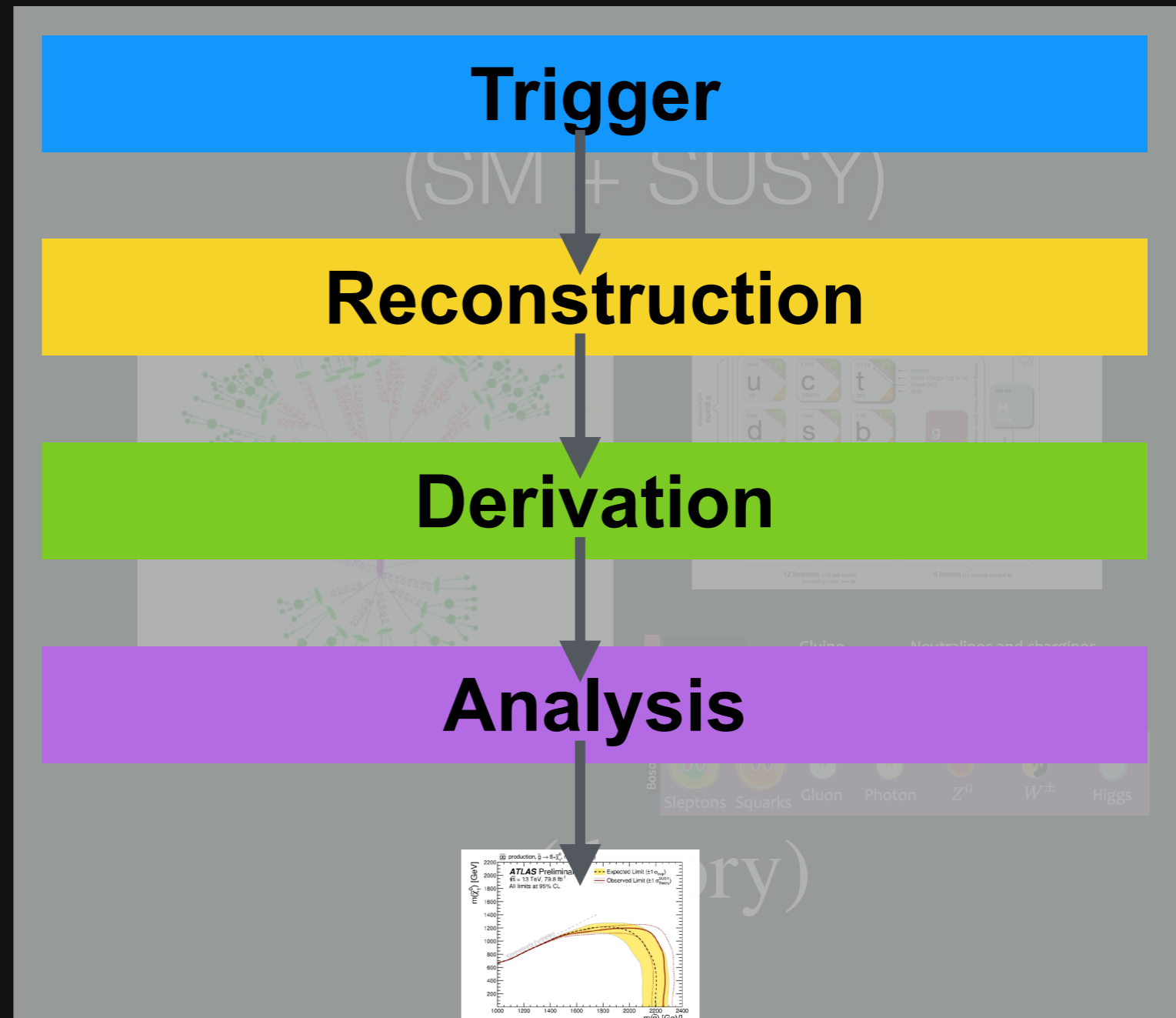
! How good is our modeling?

The Big Picture (II)

observations



$p(\text{data} = \text{observed})$



What is a statistical model?

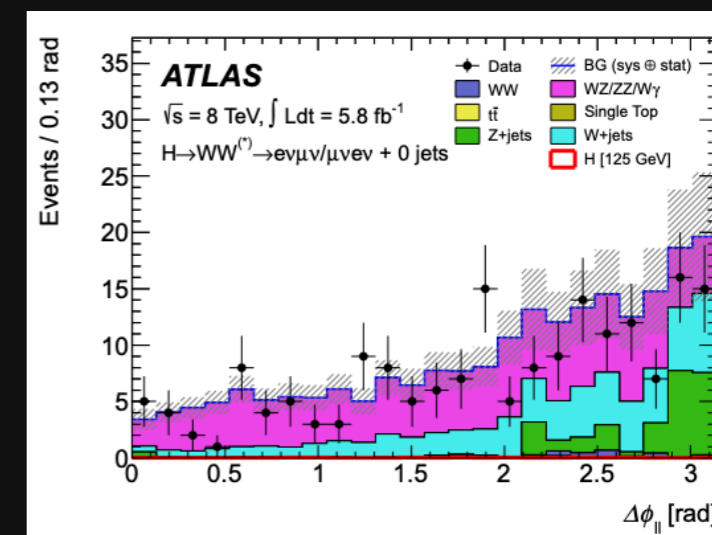
$$p(\text{data} \mid \text{theory})$$

- ✦ In a perfect world, we know the model.
- ✦ However, due to uncertainties and the high-dimensional phase-space, most analyses will only approximate the model, using a statistical model

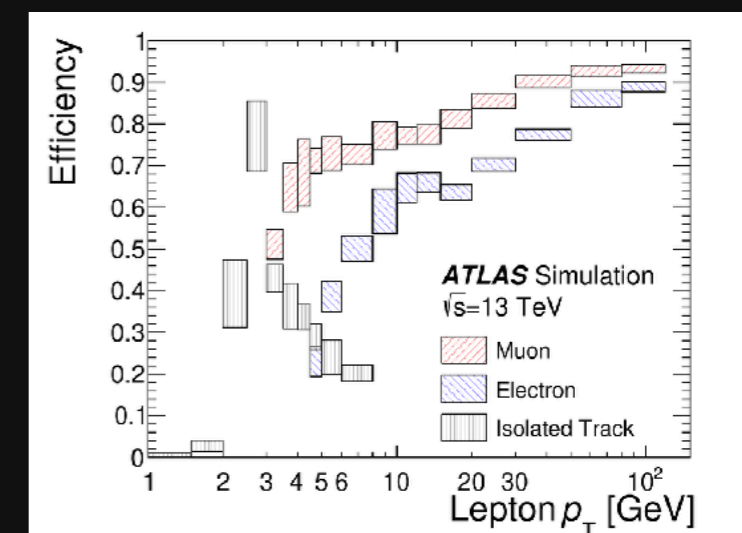
$$p_{\text{approx}} \equiv \hat{p}(q(x) \mid \text{theory})$$

- ✦ Make this more tractable by sampling from the true model and constructing an approximate likelihood

a projection of n -dimensional physics into one kinematic variable



how well can our ATLAS detector identify low p_T leptons?

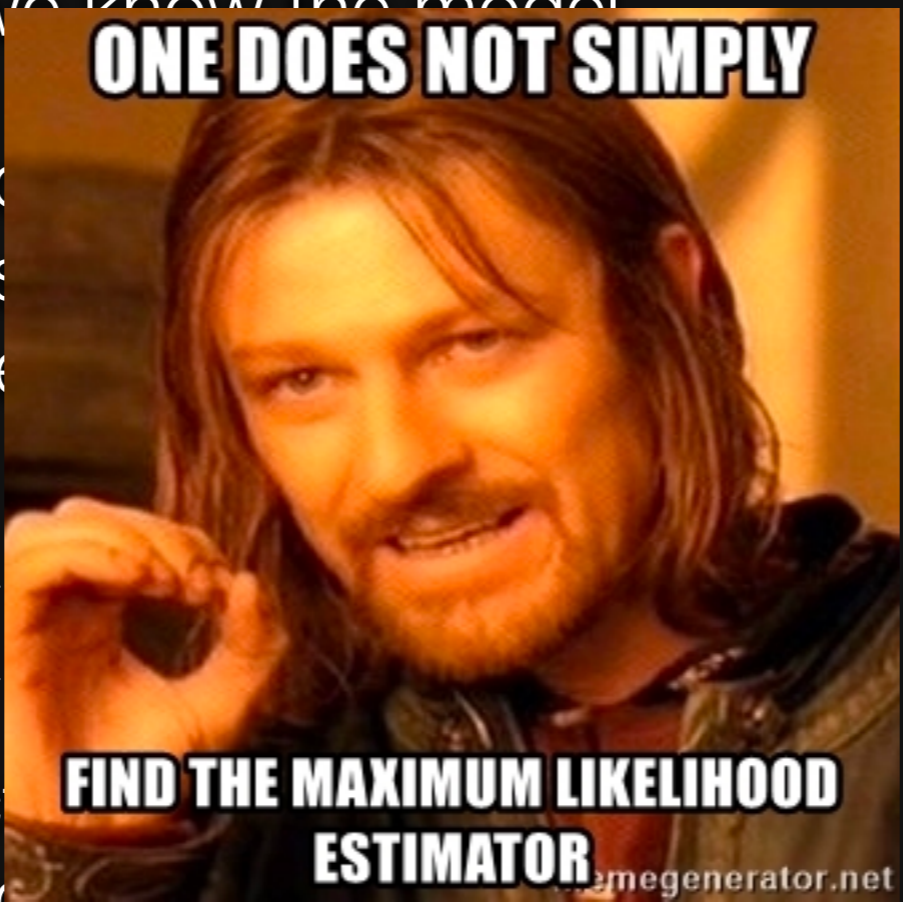


! A probability distribution function

What is a statistical model?

$$p(\text{data} | \text{theory})$$

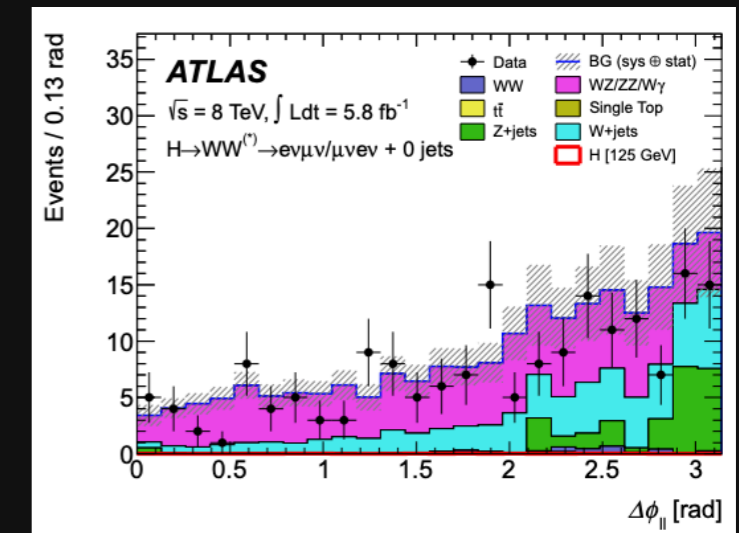
- In a perfect world, we know the model
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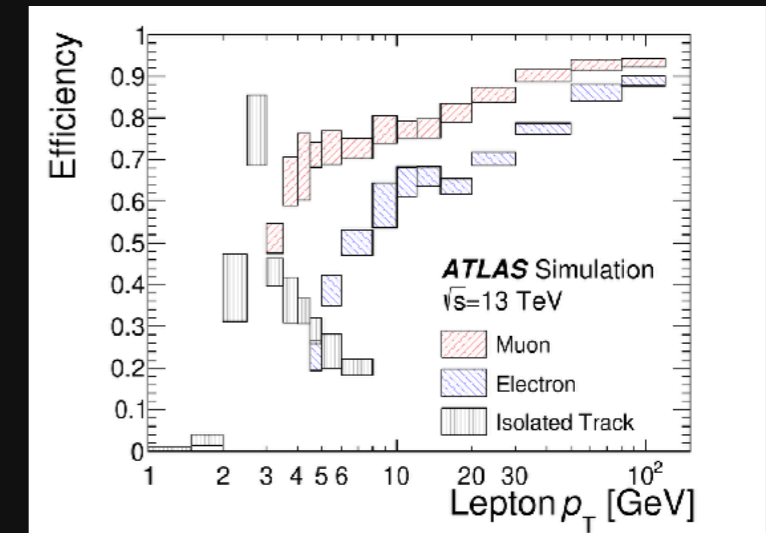
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- Make this more tractable by approximating the true model and constructing an approximate likelihood

a projection of n -dimensional physics into one kinematic variable



how well can our ATLAS detector identify low p_T leptons?



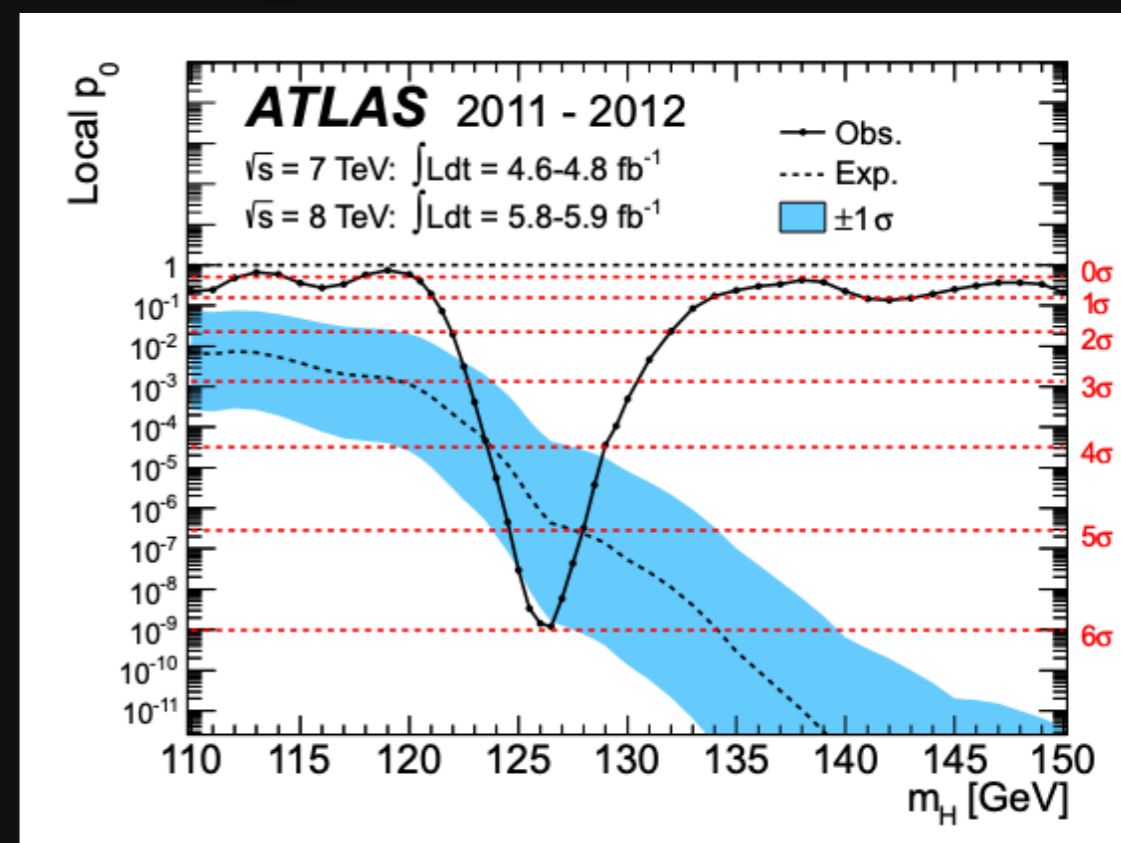
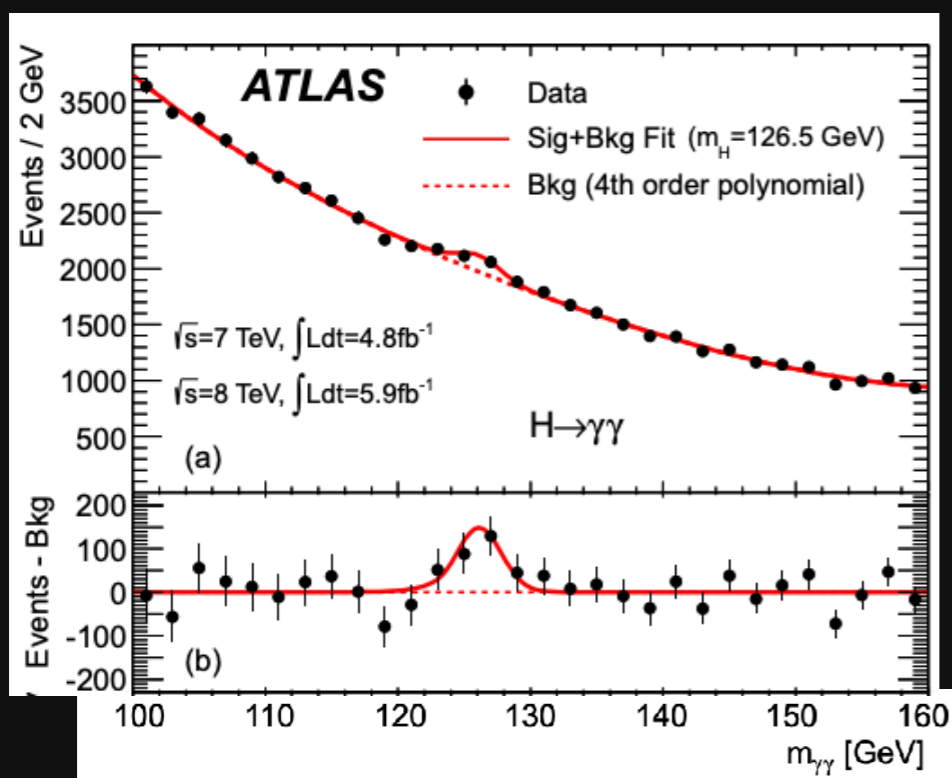
! A probability distribution function

What is a statistical model?

$$\hat{p}(\text{data} \mid \text{theory})$$

Higgs discovery p -value $< 10^{-7}$

$$m_h = 125.7 \pm 0.7 \text{ GeV}$$



Two statistical models in orange

- one representing “background” [SM, excluding Higgs]
- one representing “signal+background” [SM, including Higgs mechanism]

Using observations (black data points) and a statistical model, **inferences** include:

- p -values
- confidence intervals
- limits
- expected yields
- data/MC comparisons, etc...

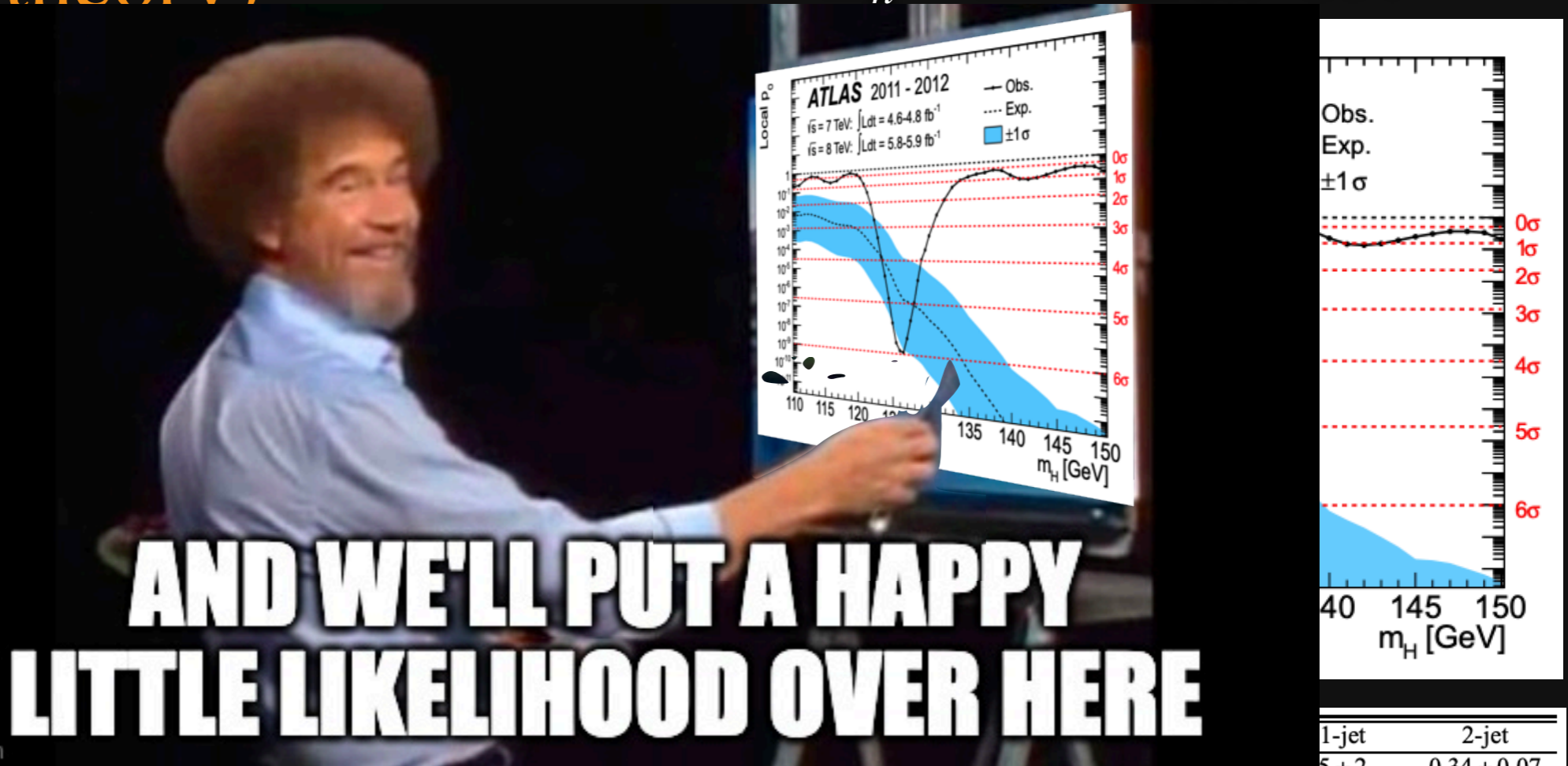
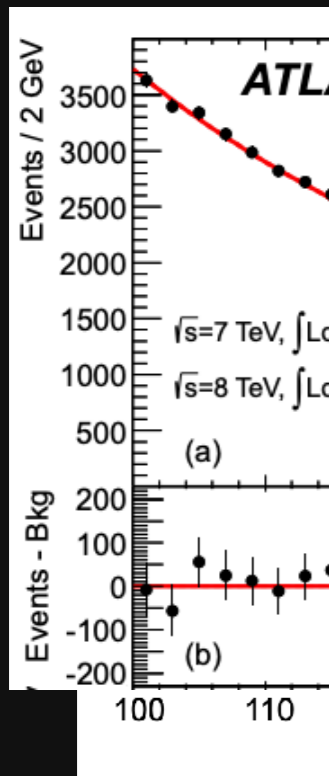
	0-jet	1-jet	2-jet
Signal	20 ± 4	5 ± 2	0.34 ± 0.07
WW	101 ± 13	12 ± 5	0.10 ± 0.14
$WZ^{(*)}/ZZ/W\gamma^{(*)}$	12 ± 3	1.9 ± 1.1	0.10 ± 0.10
$t\bar{t}$	8 ± 2	6 ± 2	0.15 ± 0.10
$tW/tb/tqb$	3.4 ± 1.5	3.7 ± 1.6	-
$Z/\gamma^* + \text{jets}$	1.9 ± 1.3	0.10 ± 0.10	-
$W + \text{jets}$	15 ± 7	2 ± 1	-
Total Background	142 ± 16	26 ± 6	0.35 ± 0.18
Observed	185	38	0

What is a statistical model?

Higgs discovery p -value $< 10^{-7}$

$$m_h = 125.7 \pm 0.7 \text{ GeV}$$

$$\hat{p}(\text{data} | \text{theory})$$



Two statistical

- one representing “background” [SM, excluding Higgs]
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- limits
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	1-jet	2-jet
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WW	101 ± 13	12 ± 5
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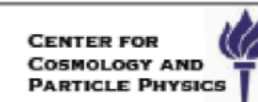
12 years ago...

he was also at
this workshop



Jason Nielsen
SCIPP director

The situation 10 years ago...



Origins I: The First “Statistics in HEP” conference

WORKSHOP ON CONFIDENCE LIMITS

CERN, Geneva, Switzerland
17–18 January 2000

CERN 2000-005

Massimo Corradi

Does everybody agree on this statement, to publish likelihoods?

Louis Lyons

Any disagreement? Carried unanimously. That’s actually quite an achievement for this Workshop.

...[Fred James wants to be able to calculate coverage, Don Groom wants to be able to calculate goodness of fit]...

Cousins

I thought the point of unanimity was that publishing the likelihood function was a *necessary* condition, not a sufficient condition.

But a practical problem remained: How to communicate multi-D likelihood?

<http://indico.cern.ch/conferenceDisplay.py?confId=100458>

⚠ ATLAS reminded everyone that we all agreed in 2000 to publish likelihoods!

he was also at
this workshop



Jason Nielsen
SCIPP director



⚠️ ATLAS reminded everyone that we all agreed in 2000 to publish likelihoods!

So we (3) did it...



G. Stark



M. Feickert



L. Heinrich



New open release streamlines interactions with theoretical physicists

The ATLAS Collaboration has released the first open likelihoods from an LHC experiment.

12th December 2019 | By [Katarina Anthony](#)



Explore ATLAS open likelihoods on the HEPData platform. (Original image: Ahmet Anil Sen/Behance)



Courtesy of CERN

ATLAS releases 'full orchestra' of analysis instruments

01/14/21 | By Stephanie Melchor

The ATLAS collaboration has begun to publish likelihood functions, information that will allow researchers to better understand and use their experiment's data in future analyses.

Meyrin, Switzerland, sits serenely near the Swiss-French border, surrounded by green fields and the beautiful Rhône river. But a hundred

<https://atlas.cern/updates/news/new-open-likelihoods>

<https://www.symmetrismagazine.org/article/atlas-releases-full-orchestra-of-analysis-instruments>

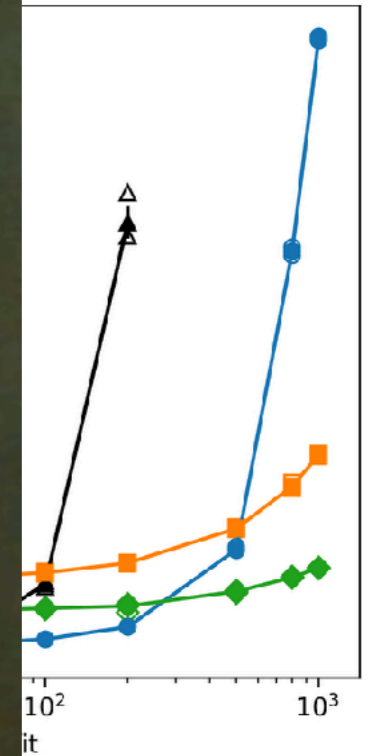
Series

TO DO STATS

unch binder




Reproducing s
ATLAS experi
st




ONE MUST LEARN ROOT

Serialized and Published!

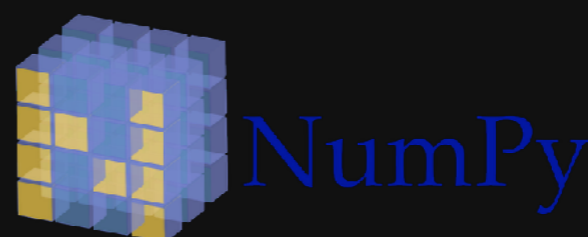
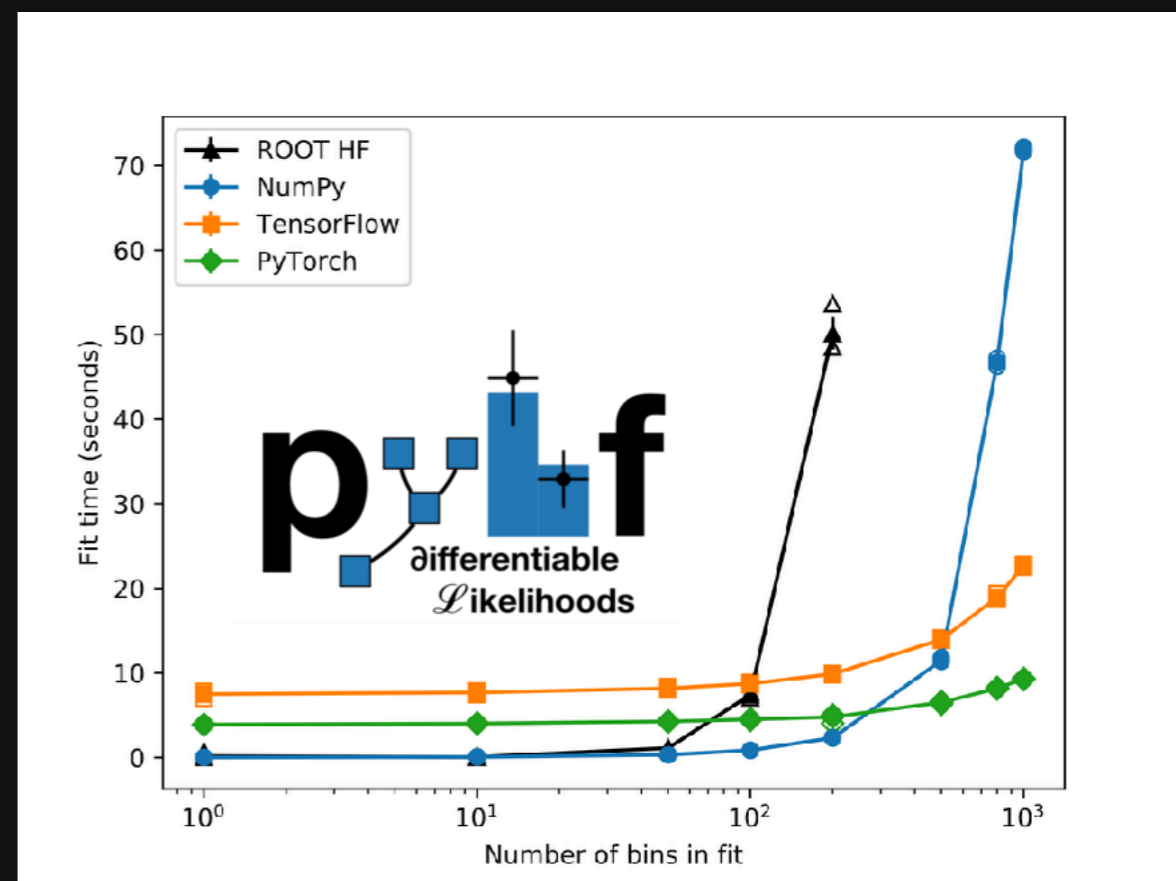



ATLAS PUB Note
ATL-PHYS-PUB-2019-029
21st October 2019



Reproducing searches for new physics with the ATLAS experiment through publication of full statistical likelihoods

The ATLAS Collaboration

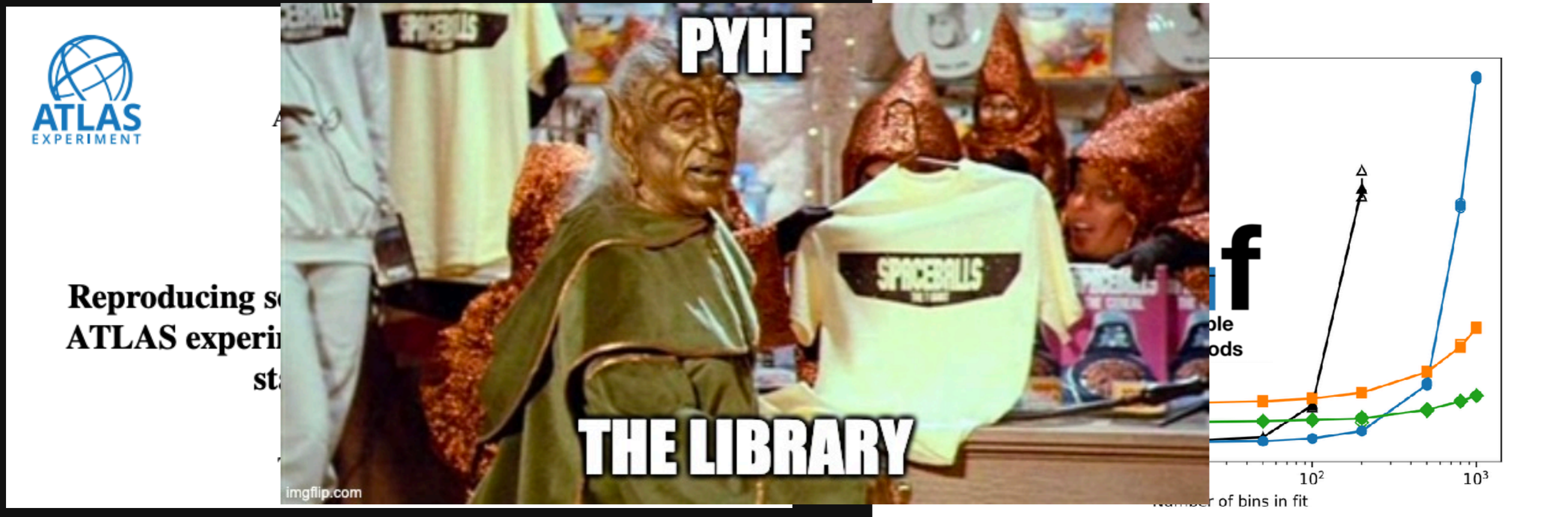


```
“python -m pip install pyhf”
```

 **orders-of-magnitude faster inference**

Serialized and Published!

 launch binder



ATLAS EXPERIMENT

Reproducing s
ATLAS exper
st

PYHF

SPACEBALLS

THE LIBRARY

f

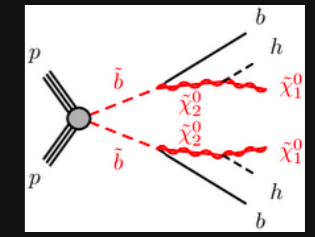
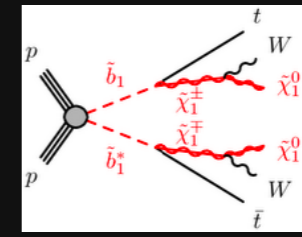
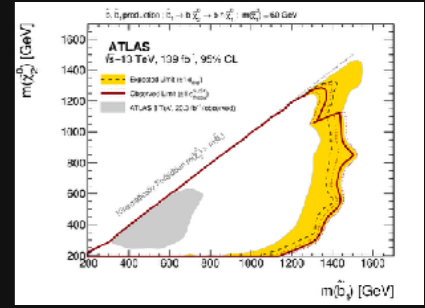
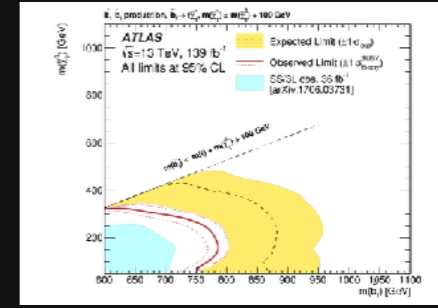
number of bins in fit



```
“python -m pip install pyhf”
```

 orders-of-magnitude faster inference

More public l'hoods! (since 2021!)



SUSY-2018-41

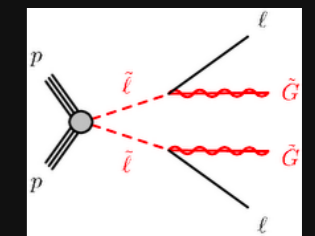
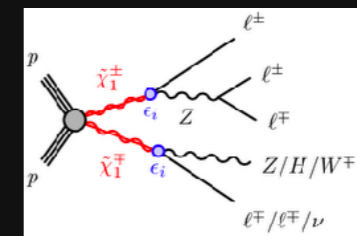
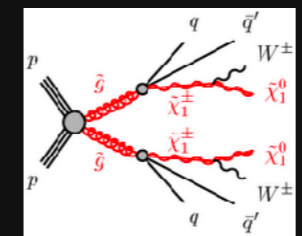
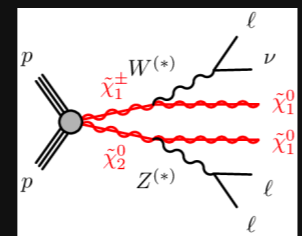
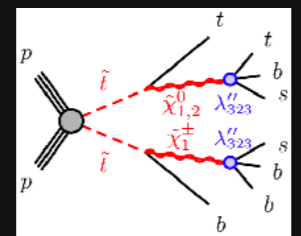
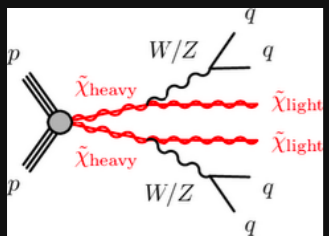
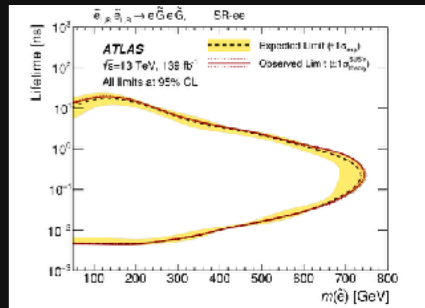
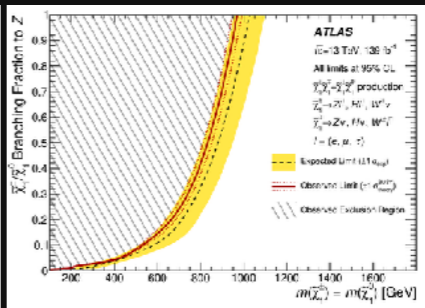
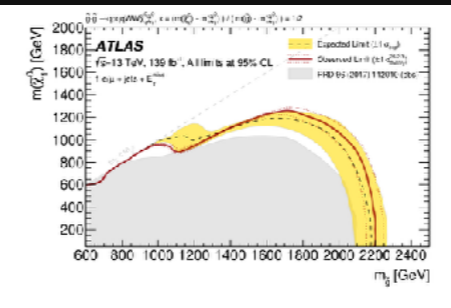
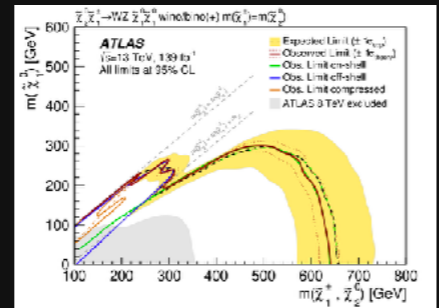
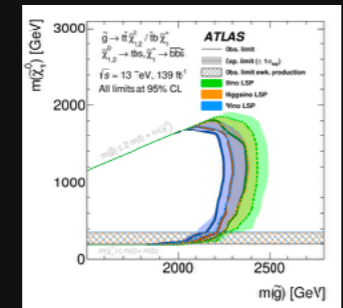
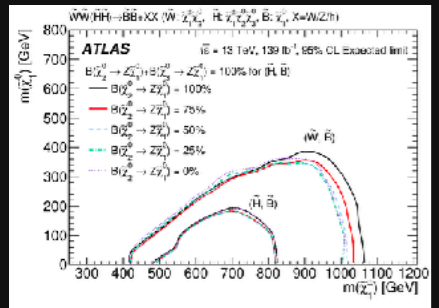
SUSY-2019-04

SUSY-2019-09

SUSY-2018-10

SUSY-2018-36

SUSY-2018-14



SUSY-2018-22

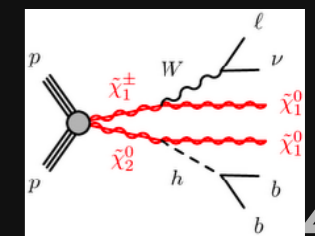
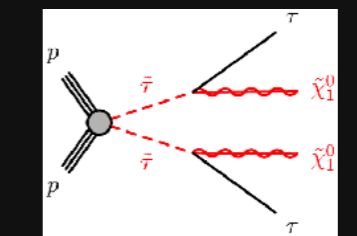
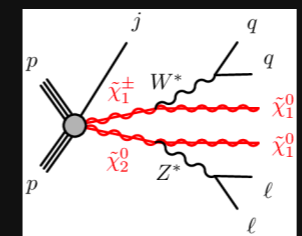
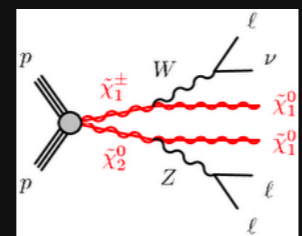
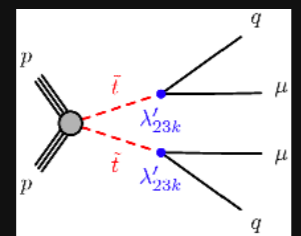
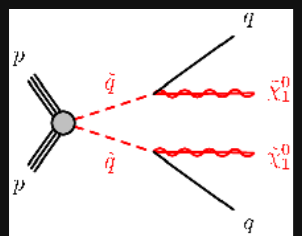
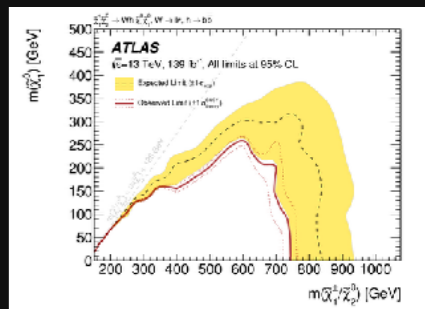
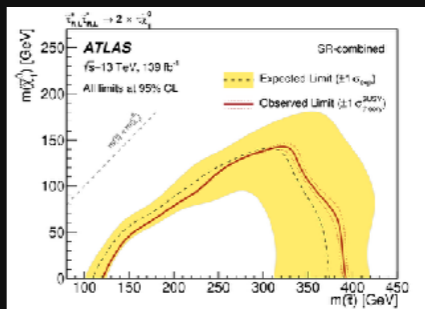
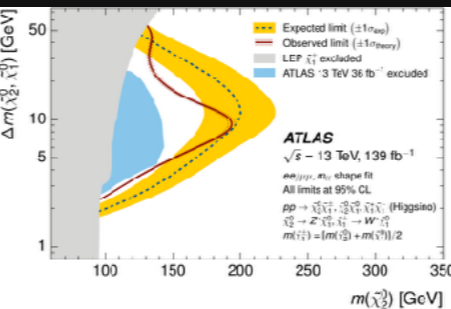
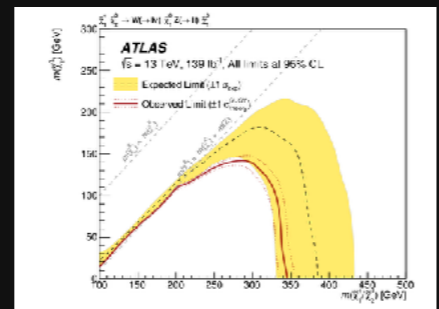
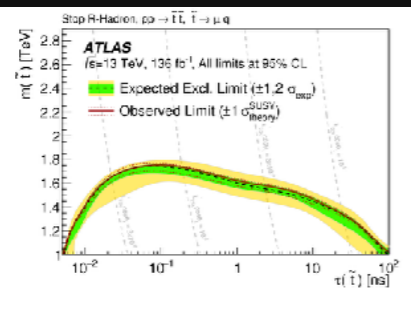
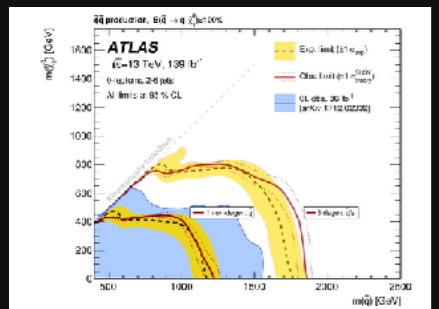
SUSY-2018-33

SUSY-2018-06

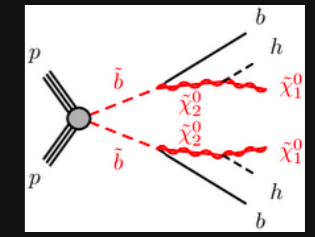
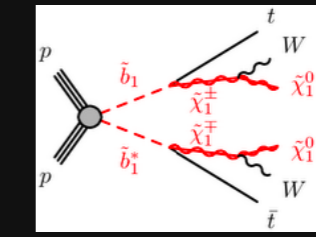
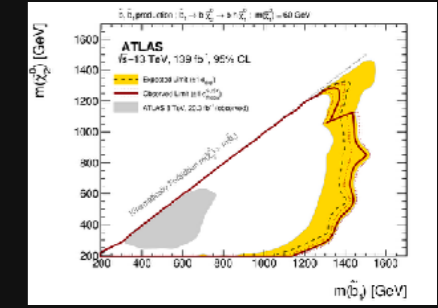
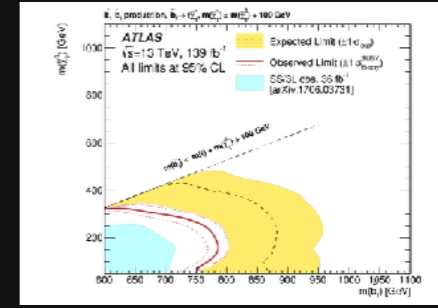
SUSY-2018-16

SUSY-2018-04

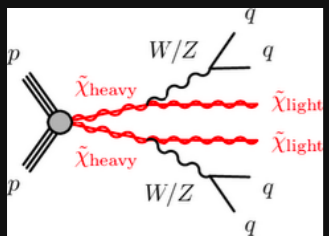
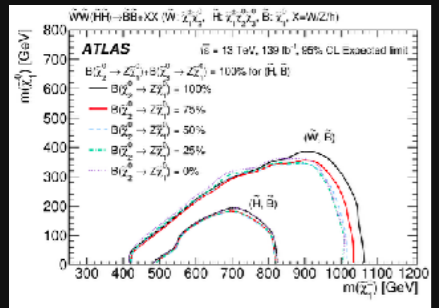
SUSY-2019-08



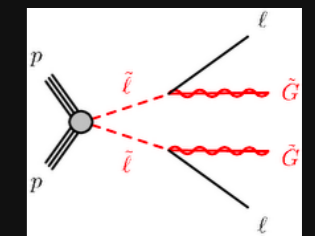
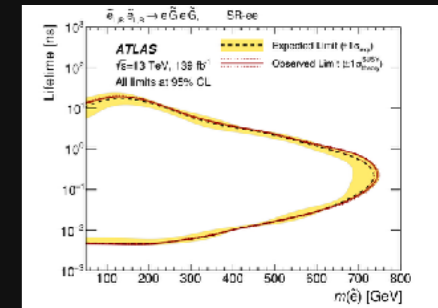
More public l'hoods! (since 2021!)



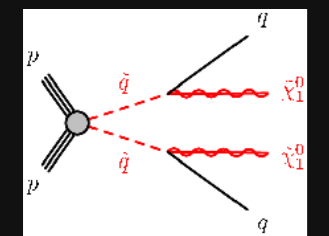
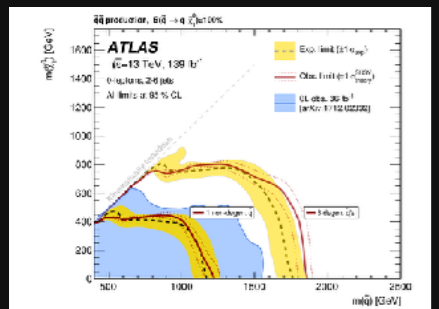
SUSY-2018-41



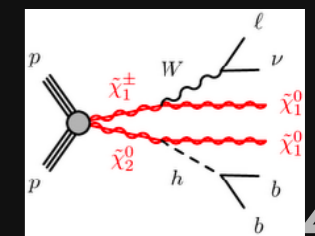
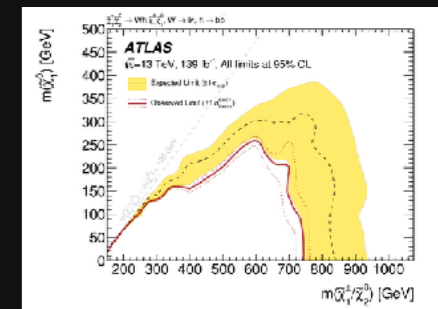
SUSY-2018-14



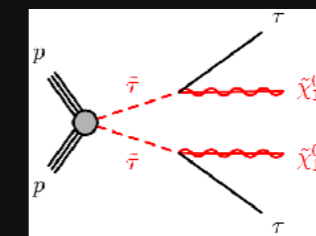
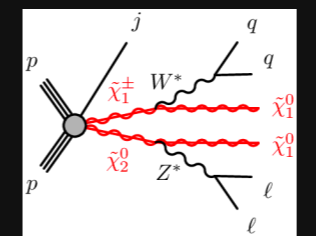
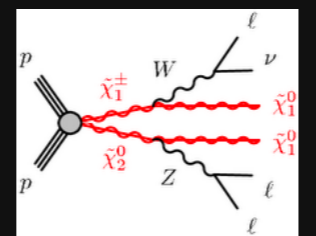
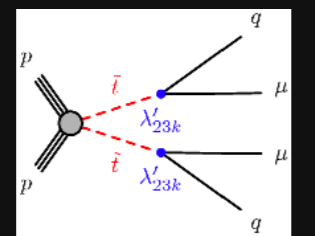
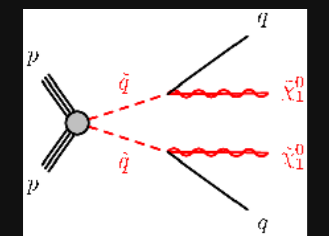
SUSY-2018-22



SUSY-2019-08



STOP TRYING TO MAKE LIKELIHOODS HAPPEN



! Builds on top of my work with pyhf

Integration into theory tools

“if you build it, they will come”



 [arXiv:2009.01809](https://arxiv.org/abs/2009.01809)

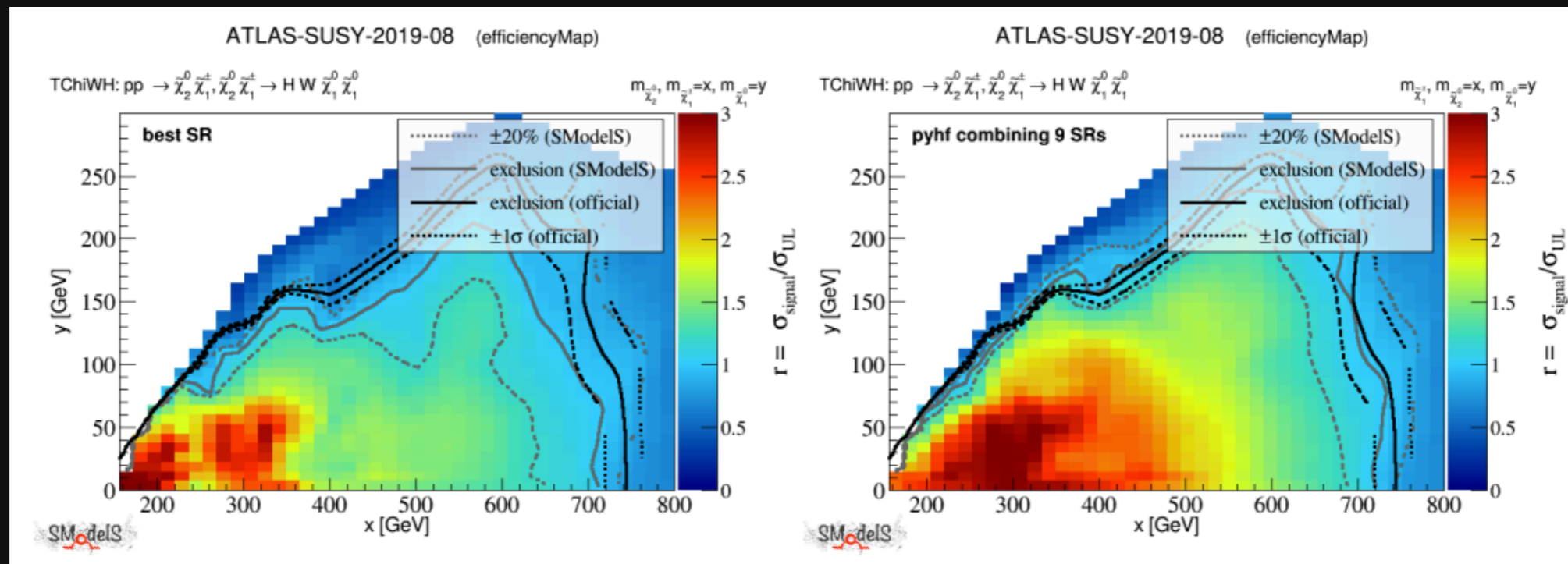
A SModelS interface for pyhf likelihoods

Gaël Alguero^a, Sabine Kraml^a, Wolfgang Waltenberger^{b,c}

^aLaboratoire de Physique Subatomique et de Cosmologie, Université Grenoble-Alpes, CNRS/IN2P3, 53 Avenue des Martyrs, F-38026 Grenoble, France

^bInstitut für Hochenergiephysik, Österreichische Akademie der Wissenschaften, Nikolsdorfer Gasse 18, 1050 Wien, Austria

^cUniversity of Vienna, Faculty of Physics, Boltzmanngasse 5, A-1090 Wien, Austria



...and other experiments

LHC QCD

How to discover QCD Instantons at the LHC¹

Simone Amoroso^a Deepak Kar^b Matthias Schott^{2c}

^aDESY, Homburg, Germany

^bUniversity of Witwatersrand, South Africa

^cJohannes Gutenberg-University, Mainz, Germany

E-mail: matthias.schott@cern.ch

arXiv:2102.06176

Charged Lepton Flavor Violation at the EIC

EIC

Vincenzo Cirigliano, Kaori Fuyuto, Christopher Lee, Emanuele Mereghetti, and Bin Yan

Theoretical Division, Los Alamos National Laboratory, Los Alamos, NM 87545, U.S.A.

E-mail: cirigliano@lanl.gov, kfuyuto@lanl.gov, clee@lanl.gov,

emereghetti@lanl.gov, binyan@lanl.gov

Sensitivity of Future Hadron Colliders to Leptoquark Pair Production in the Di-Muon Di-Jets Channel

B. C. Allanach¹, Tyler Corbett², Maeve Madigan^{a,1}

¹DAMTP, University of Cambridge, Wilberforce Road, Cambridge, CB3 0WA, United Kingdom

²The Niels Bohr International Academy, Blegdamsvej 17, University of Copenhagen, DK-2100 Copenhagen, Denmark

arXiv:1911.04455

FCC

Search for new phenomena in events with two opposite-charge leptons, jets and missing transverse momentum in pp collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector

The ATLAS Collaboration

arXiv:2102.01444

ATLAS

SEARCH FOR $B^+ \rightarrow K^+ \nu \bar{\nu}$ DECAYS WITH AN INCLUSIVE TAGGING METHOD AT THE BELLE II EXPERIMENT

On the single leptoquark solutions to the B -physics anomalies

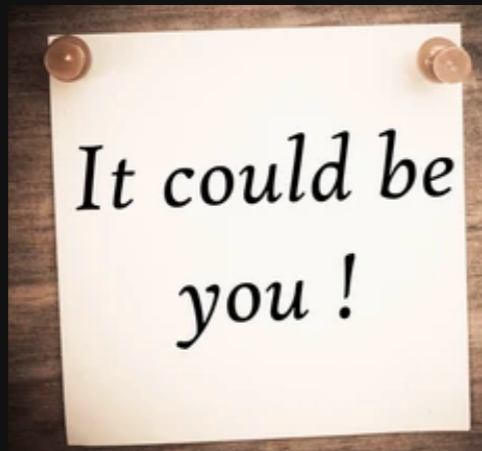
Andrei Angelescu,^{1,*} Damir Bečirević,^{2,†} Darius A. Faroughy,^{3,‡} Florentin Jaffredo,^{2,§} and Olcyr Sumensari^{2,¶}

¹Max-Planck-Institut für Kernphysik, Saupfercheckweg 1, 69117 Heidelberg, Germany

arXiv:2103.12504

arXiv:2105.05754

Belle II



ALICE

μ -collider

Hunting wino and higgsino dark matter at the muon collider with disappearing tracks

Rodolfo Capdevilla,^{a,b} Federico Meloni,^c Rosa Simoniello,^d Jose Zurita^e

^aDepartment of Physics, University of Toronto, Canada

arXiv:2102.11292


Published Models (I)

- With a new JSON format, user-friendly tools, and increasing experimental support — key inference results can be reproduced in minutes — for physics including
 - Parton distribution functions
 - Higgs boson measurements at the LHC
 - Searches for new particles at the LHC
 - Searches for dark matter
 - Heavy flavor physics
 - World averages
 - Global fits



[arXiv:2109.04981](https://arxiv.org/abs/2109.04981)

Publishing statistical models: Getting the most out of particle physics experiments



Kyle Cranmer^{1*}, Sabine Kraml^{2†}, Harrison B. Prosper^{3§} (editors),
Bechtle⁴, Florian U. Bernlochner⁴, Itay M. Bloch⁵, Enzo Canonero⁶, Marcin
Cz⁷, Andrea Coccaro⁸, Jan Conrad⁹, Glen Cowan¹⁰, Matthew Feickert¹¹,
erreiro Iachellini^{12,13}, Andrew Fowlie¹⁴, Lukas Heinrich¹⁵, Alexander Held¹,
a Kuhr^{13,16}, Anders Kvellestad¹⁷, Maeve Madigan¹⁸, Farvah Mahmoudi^{15,19},
Dundas Morá²⁰, Mark S. Neubauer¹¹, Maurizio Pierini¹⁵, Juan Rojo⁸, Sezen
22, Luca Silvestrini²³, Veronica Sanz^{24,25}, Giordon Stark²⁶, Riccardo Torre⁸,
Thorne²⁷, Wolfgang Waltenberger²⁸, Nicholas Wardle²⁹, Jonas Wittbrodt³⁰

Applications of Likelihoods

Hold your breath. Make a wish. Count to three.
Come with me. And you'll be. In a world of...

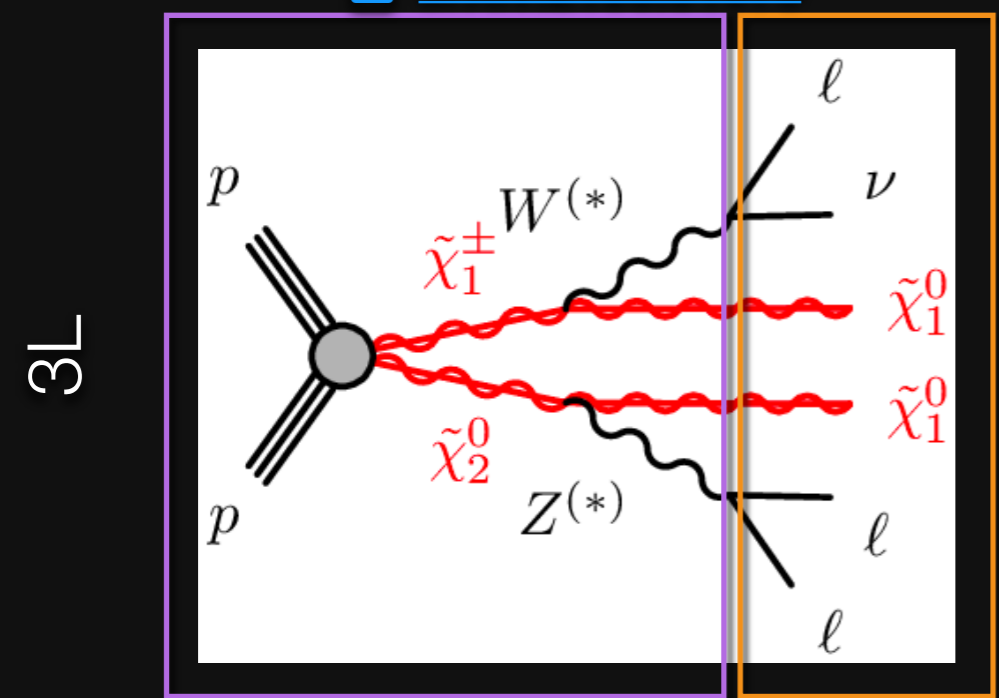
"Pure imagination."

— Willy Wonka

E.G.: Stat. Combination (I)

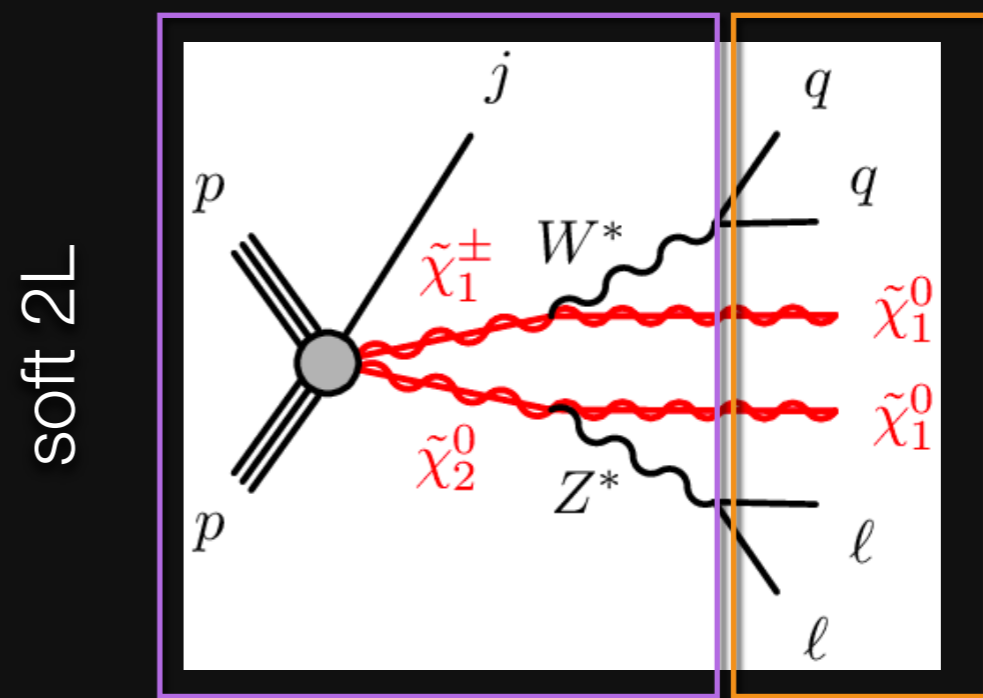
- Multiple analyses with different **signatures** can still target the same **model**
 - See Zach Marshall's talk on differences between signature and model: <https://indico.cern.ch/event/1023573/contributions/4400586/>

[ATLAS-CONF-2020-015](https://atlas.conf.cern.ch/2020/015)
[arXiv:2106.01676](https://arxiv.org/abs/2106.01676)



off-shell: $\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) \in [5, 90]$ GeV
 on-shell: $\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) \geq 90$ GeV
 signature: $3\ell + 0j + E_T^{\text{miss}}$

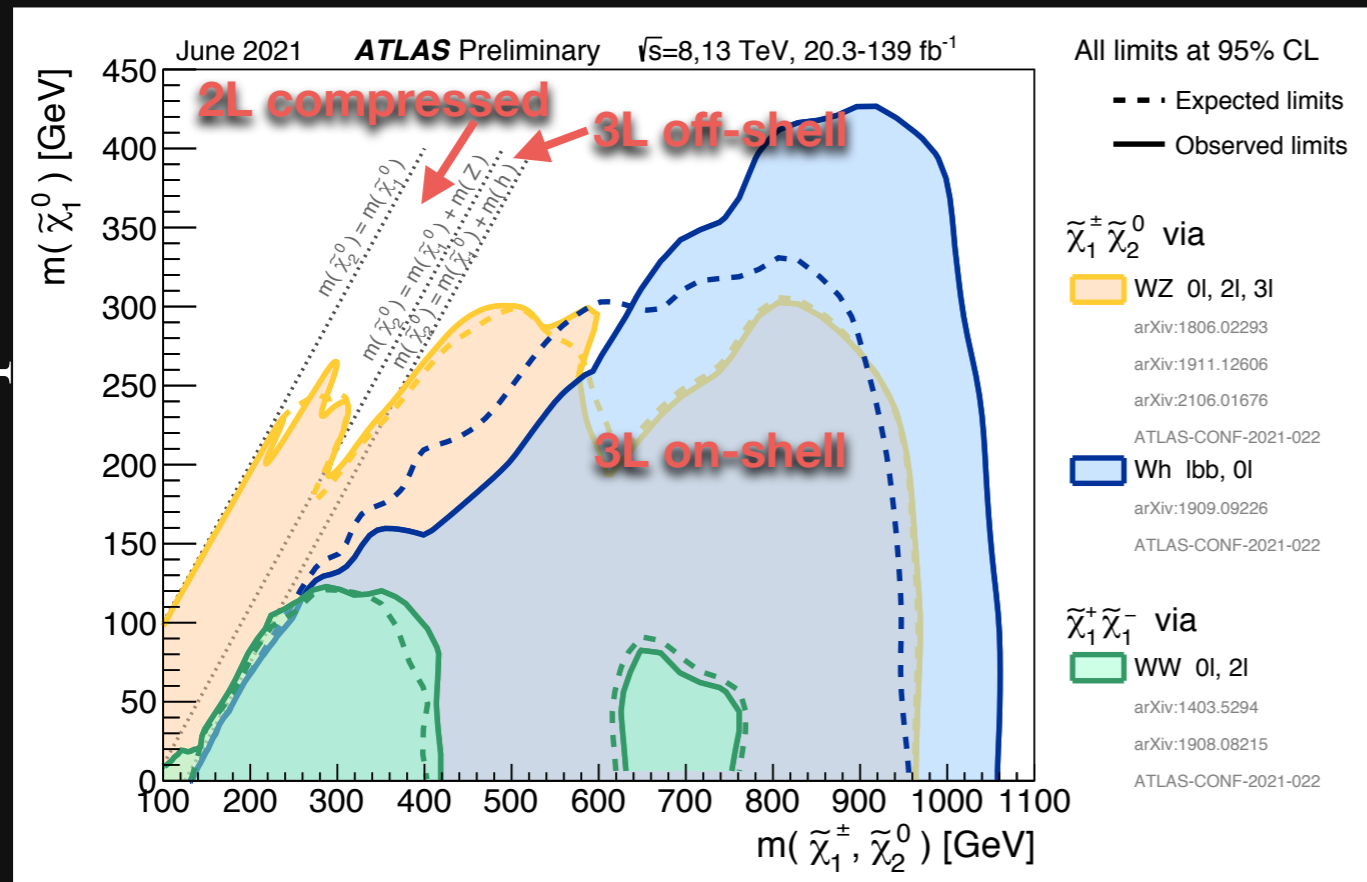
[arXiv:1911.12606](https://arxiv.org/abs/1911.12606)



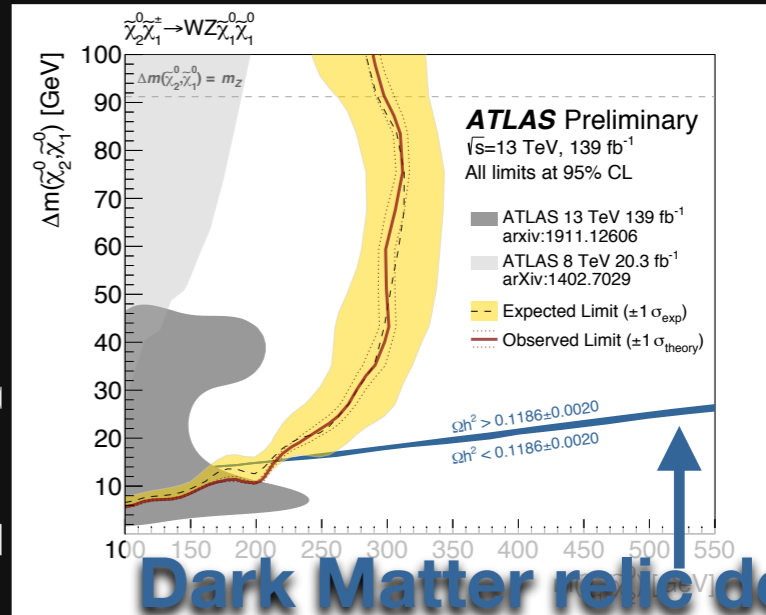
soft 2L: $\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) \in [1, 60]$ GeV
 signature: $2\ell + 3j + E_T^{\text{miss}}$

E.G.: Stat. Combination (II)

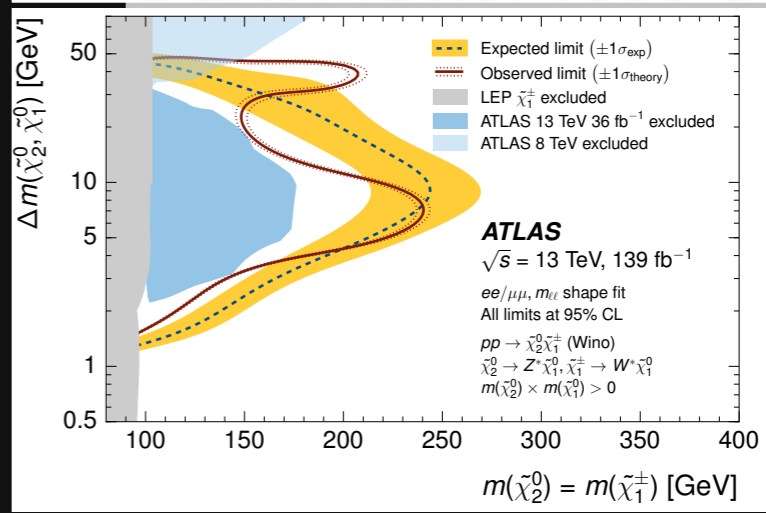
- Goal: **combine multiple searches** to paint a tapestry of our sensitivity to Higgsino/Wino-Bino models (production) which decay to on-shell/off-shell Standard Model bosons (W/Z)



$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0)$



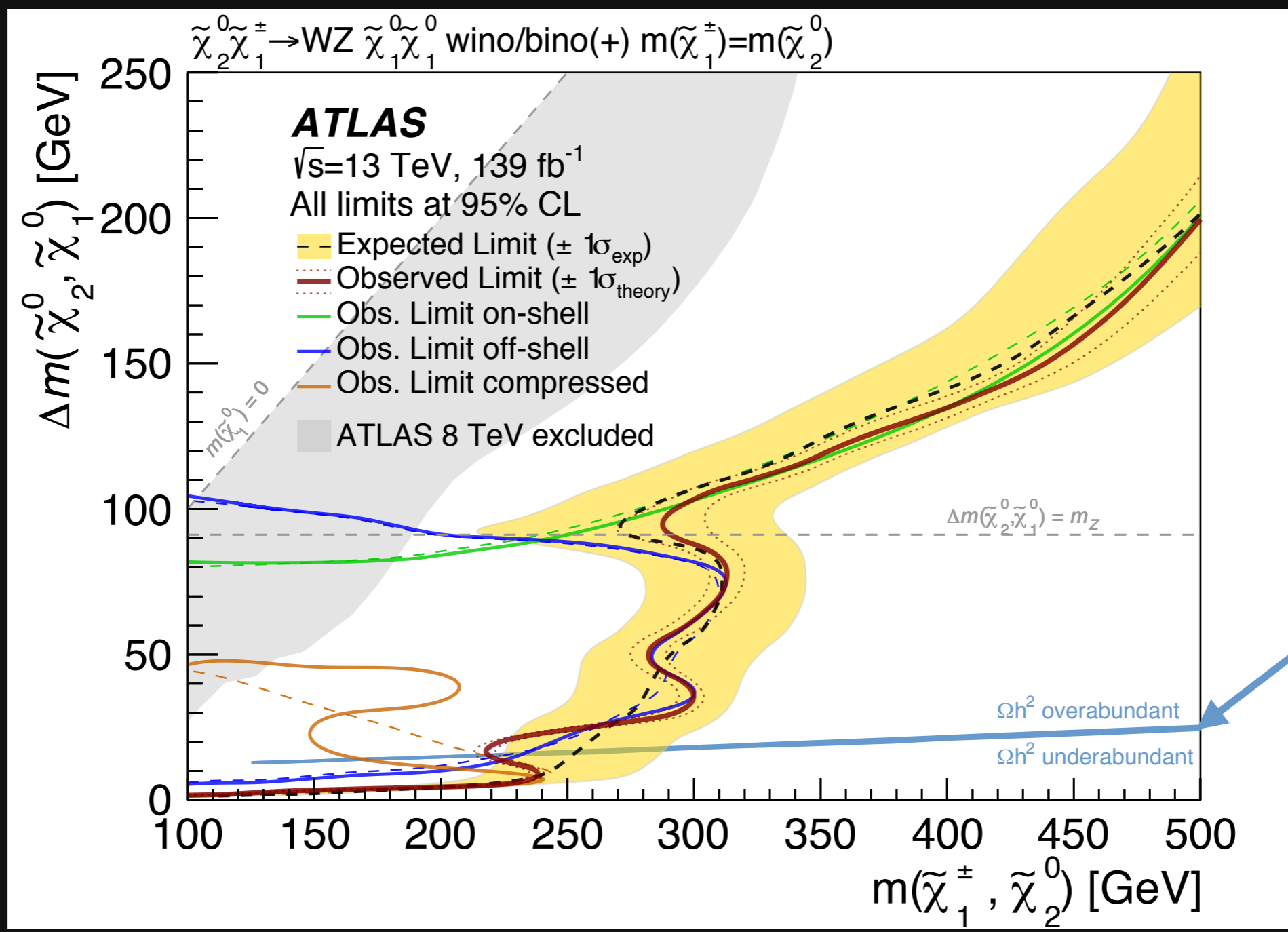
3L



2L

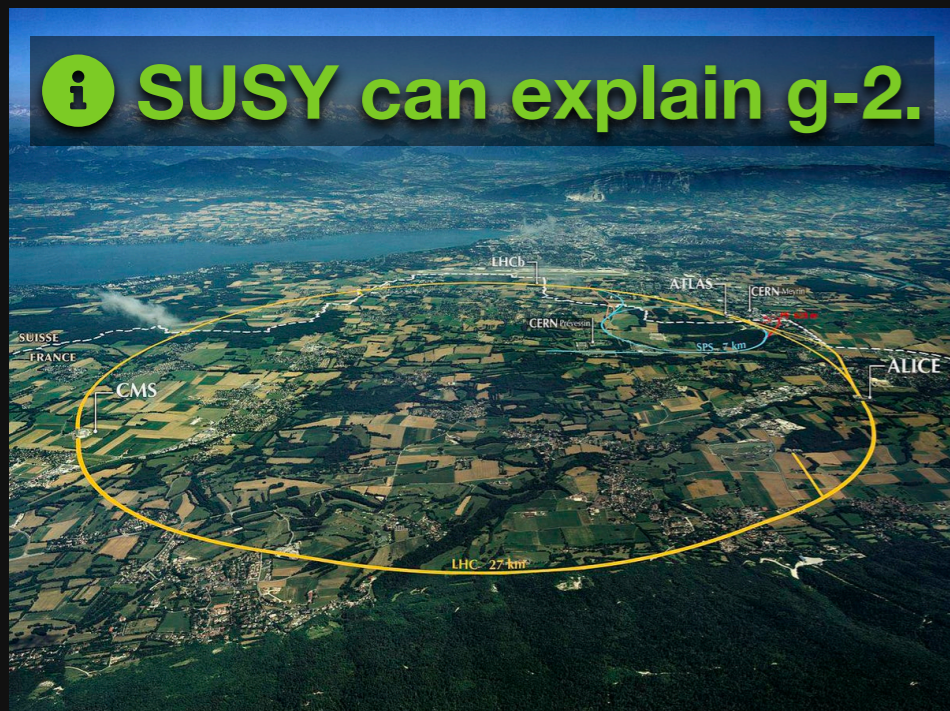
$m(\tilde{\chi}_1^+, \tilde{\chi}_2^0)$

E.G.: Stat. Combination (III)

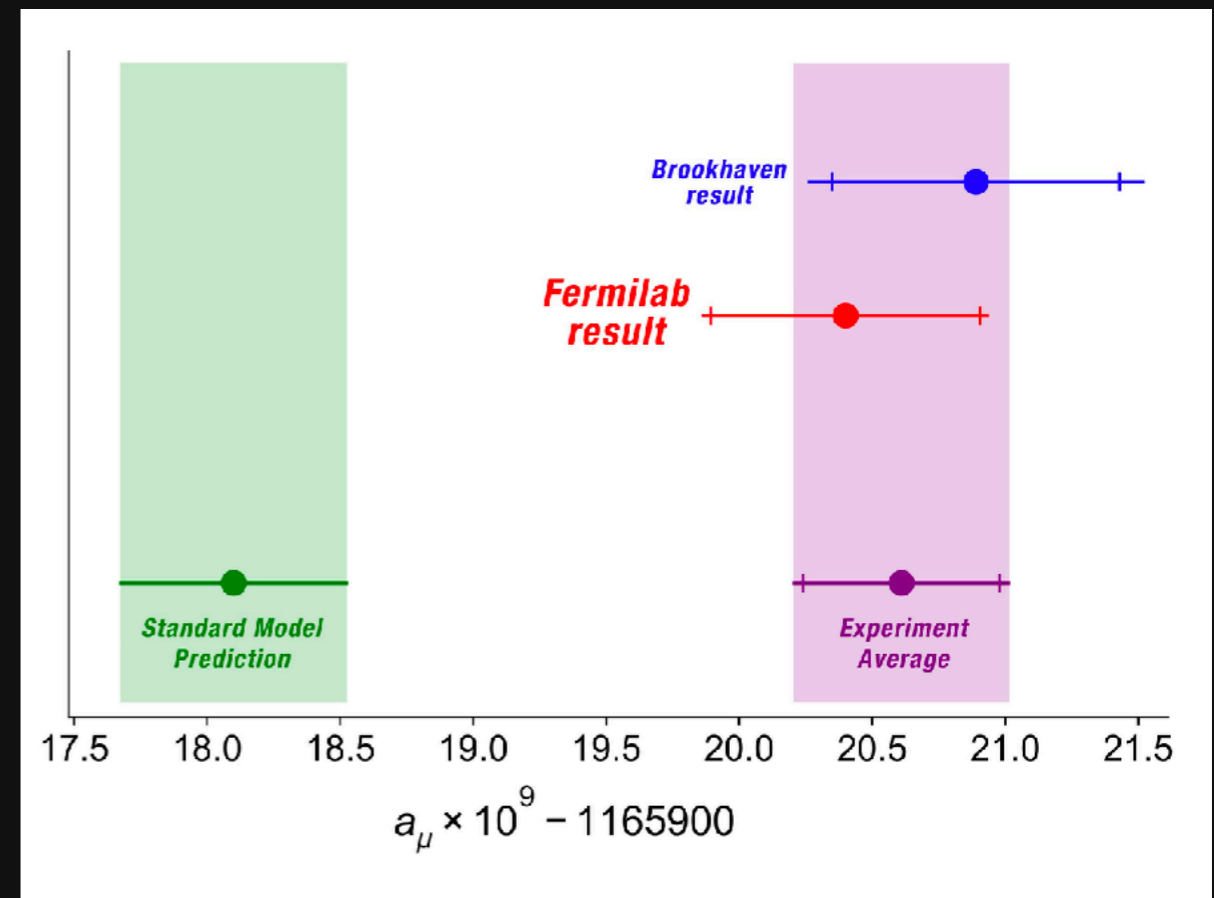


Dark Matter relic density

Example: global fit (g-2) (I)



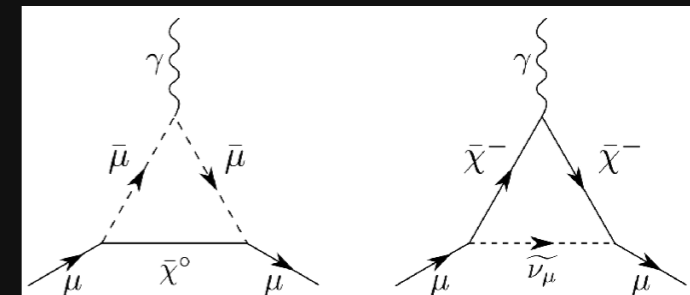
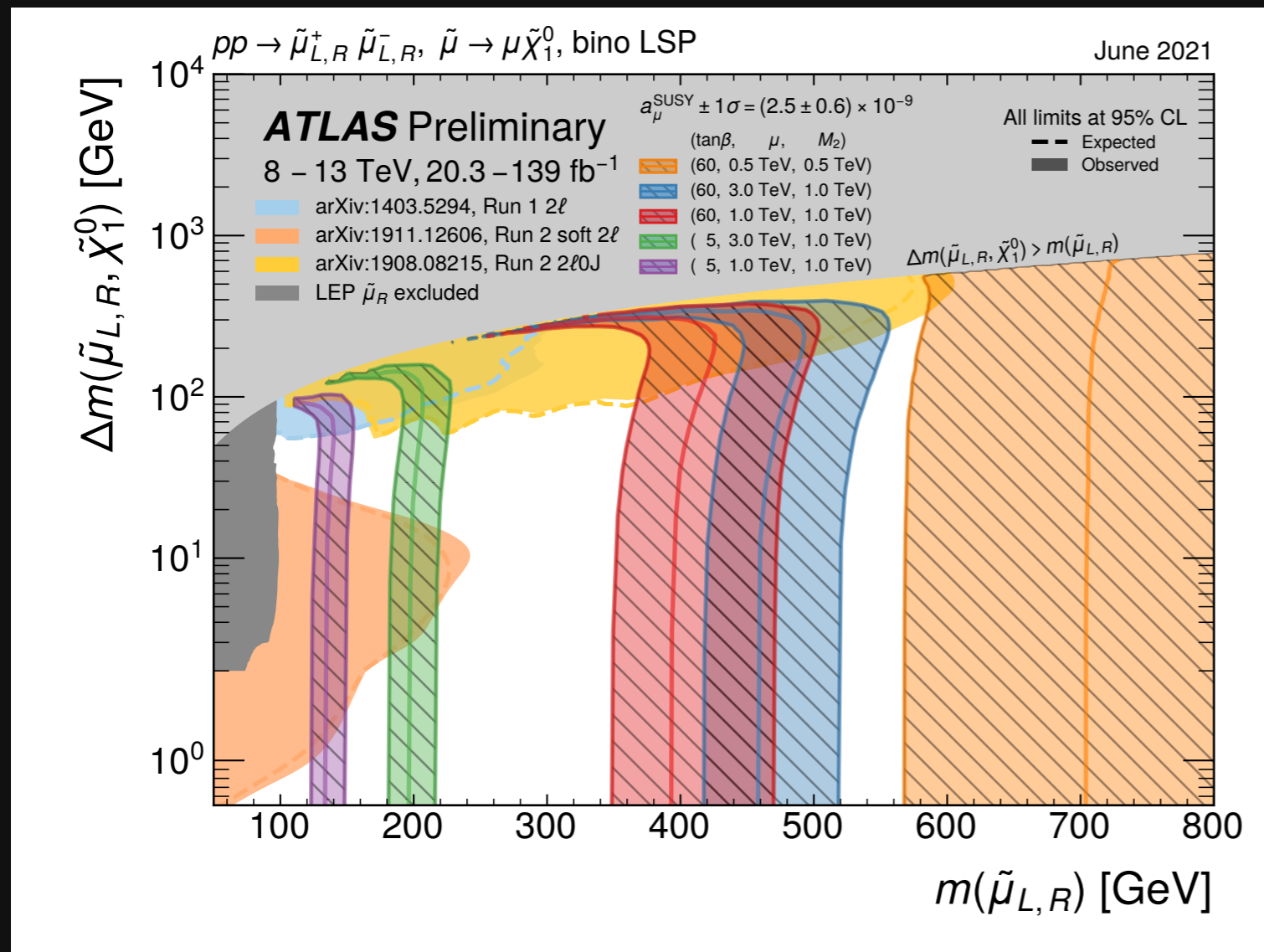
combine!



April 27th, 2021

? Can we use experimental measurements of g-2 to inform our ATLAS physics program?

Example: global fit (g-2) (II)



$$\tilde{\mu}^\pm / \tilde{\chi}_1^0 \text{ or } \tilde{\chi}_1^\pm / \tilde{\nu}_\mu$$

- Compute exclusion limits in the $(\tilde{\mu}, \tilde{\chi}_1^0)$ mass plane for different analyses that are sensitive to the smuon model with muon/bino-like $\tilde{\chi}_1^0$ signature
- Hatched bands are compatible with the observed g-2 anomaly measured by Fermilab and BNL experiments to $\pm 1\sigma$, for a variety of pMSSM parameters

center line = combined g-2 measurement

hatching = $\pm 1\sigma$ uncertainty on experimental measurement

Light smuon/light neutralino needed!

Outreach

Making physics more approachable



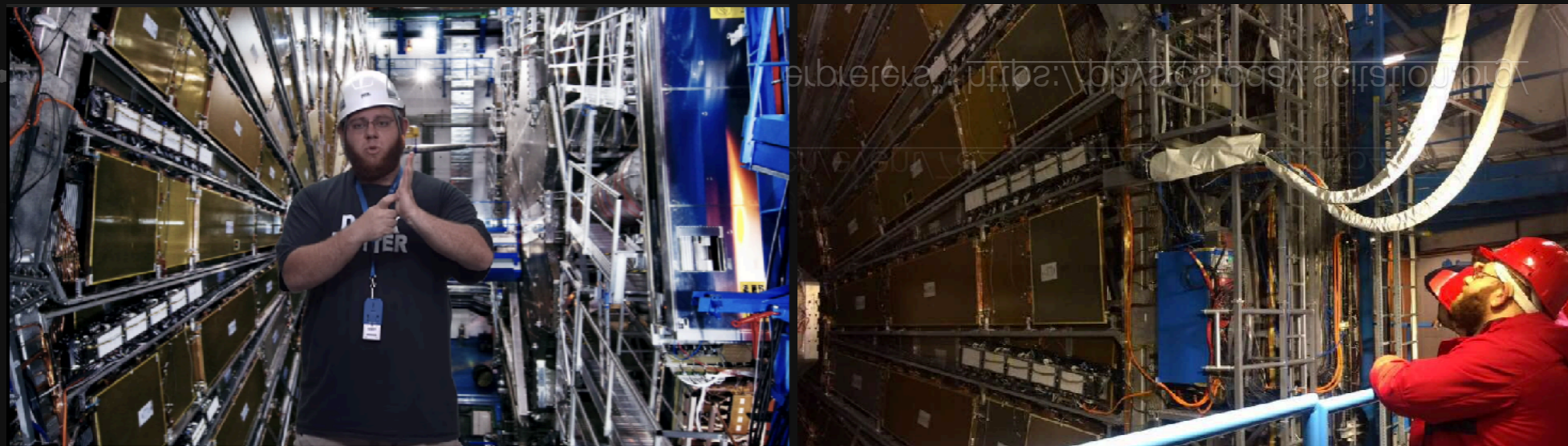
“Places such as CERN become ever more important: places where people from around the world come together to show what can be achieved when people overcome their differences to work towards common goals that ultimately bring benefit to all of humanity..”

— Fabiola Gianotti

Outreach

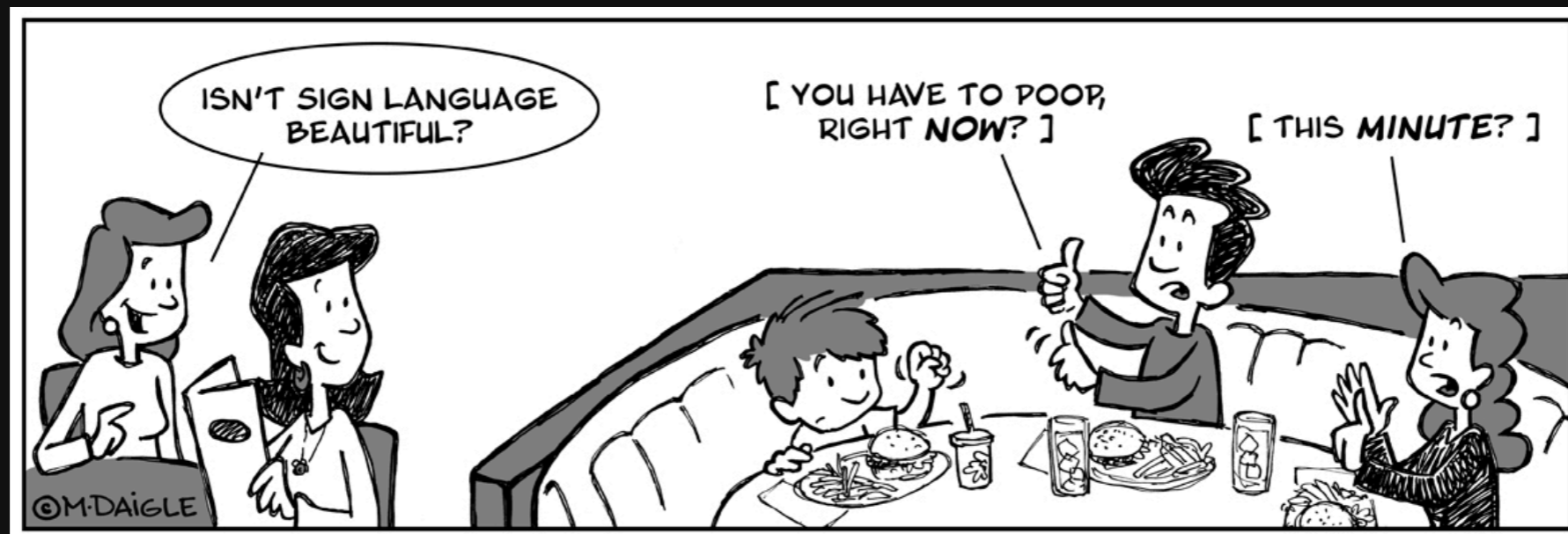


- Exhibit in ASL at the CERN museum (Microcosm): <https://www.youtube.com/watch?v=BaGjAruqFec>
- My life as a Particle Physicist (in ASL): <https://www.youtube.com/watch?v=3sESUT1UO6E>
- Fermilab Publication: “A Matter of Interpretation”: <https://www.symmetrymagazine.org/article/a-matter-of-interpretation-asl-physics>
- "PARTY CALL PHYSICS": Presentation during American Physical Society meeting about physics and accessibility: <https://indico.cern.ch/event/782953/contributions/3454898/>
- Physics Today: “Deaf Scientists thrive with interpreters”: <https://physicstoday.scitation.org/doi/10.1063/PT.6.4.20210723a/full/>



Signed Language

Making physics more accessible



Sign Language

70+ million people in the world **use sign language**

- ✦ Different from **country to country**
 - ✦ ASL, LSF, DGS, LIS, NZSL, Auslan, BSL, etc...
- ✦ Different **Grammar**
 - ✦ English (Subject-Object-Verb) vs ASL (Time-Subject-Verb-Object)
 - ✦ EXAMPLE: “The boy threw the ball.” vs “BALL, BOY THROW” / “BOY THROW BALL”
 - ✦ TOPICALIZATION: “She gave me money.” vs “MONEY? she-GIVE-me”
- ✦ Each sign is composed of **five parameters**. Altering one parameter changes the entire meaning
 - ✦ Handshape (HS), Palm Orientation (PO), Location, Movement, and Facial Expressions (NMS)
 - ✦ EXAMPLE: cool/apple/Bronx/ask/need/must/manage/scar/etc...

Phonology and Parameters

- ✦ Elementary particles are the **building blocks of matter**.
- ✦ The five parameters are the **building blocks of signs**.
- ✦ **Matter** is made up of **elementary particles**
 - ✦ **proton**: up, up, down
 - ✦ **neutron**: up, down, down
- ✦ **Signs** are built up using **parameters**
 - ✦ **mother**: HS5, PO-left, chin
 - ✦ **father**: HS5, PO-left, forehead



Examples

- **PARTICLES:** Closed small C, champagne flick multiple
- **STANDARD MODEL:** NDH O PO side, DH 4 GRID-TABLE
- **RADIATE/RADIATION:** 2H A HS TOGETHER to 5 wiggle outwards
- **ENERGY:** bent-L, shake
- **LUMINOSITY:** NDH B PO up, DH AND HS open-twist to 5 HS (make sure DH palm always touch)
- **COLLISION (Quantum Physics):** Moving FLAT O to Open-Hand 5 twist (make sure back of hands are touching at the end of the sign)

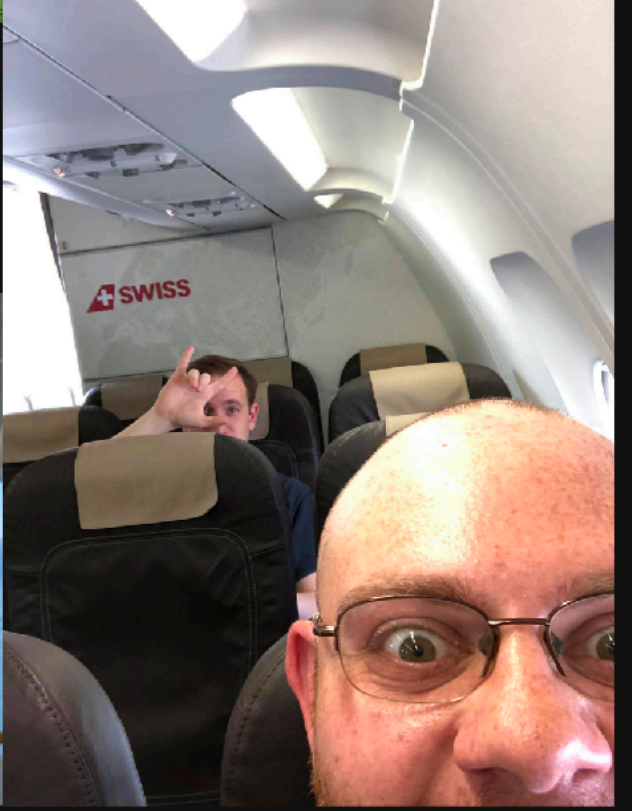


Expansion Example

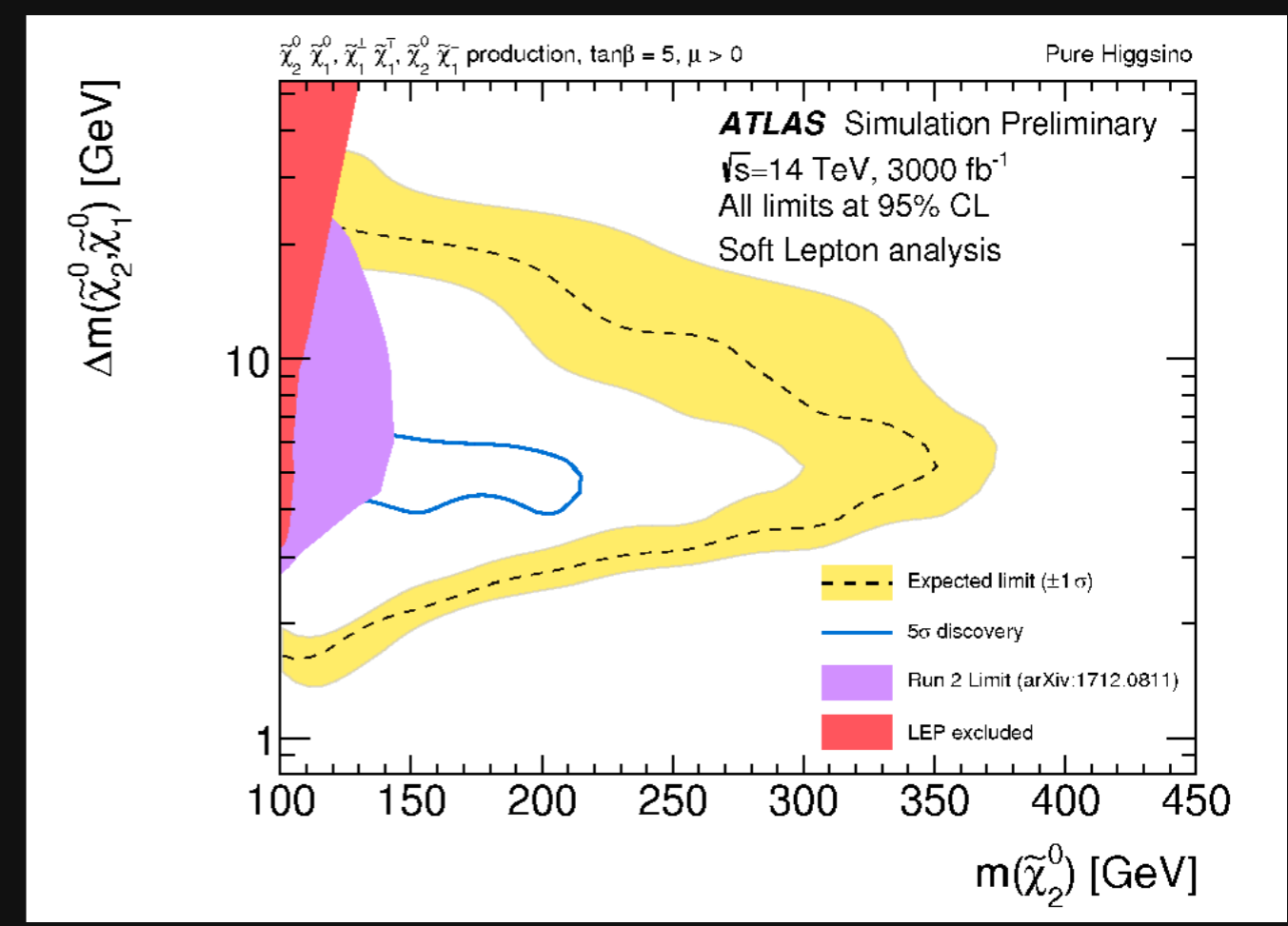
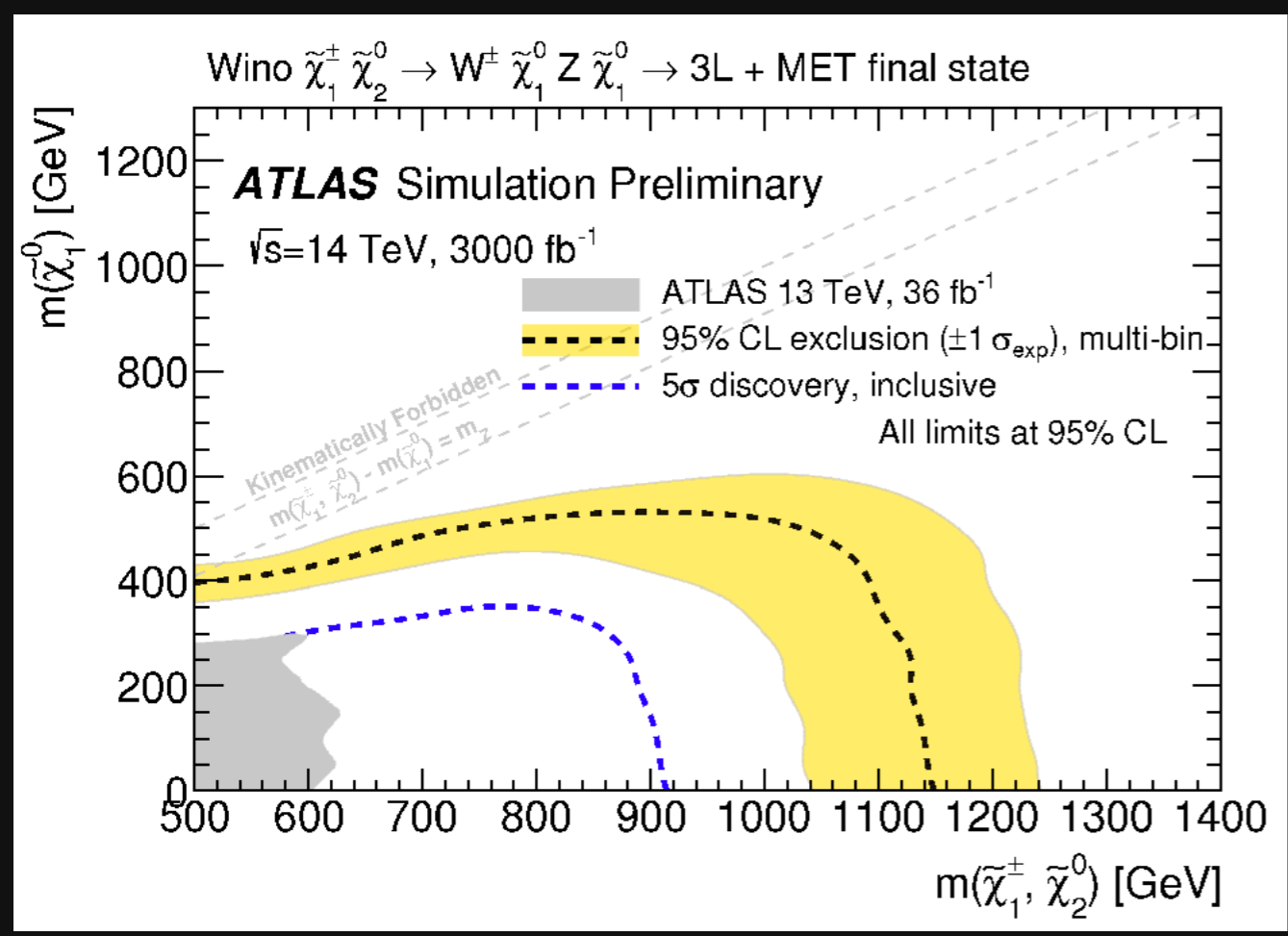


- ✦ What are atoms? Atoms are atomic particles that have many properties and are composed of three different subatomic particles: electrons, protons, and neutrons. The electrons form the atom's "cloudy" atmosphere. At the center of an atom is the nucleus where all the protons and neutrons live. Are protons and neutrons elementary particles? Nope, they can be broke up into even smaller particles. For example, the proton is made of three quarks: up, up, down.

Key features: "cinematic ASL"

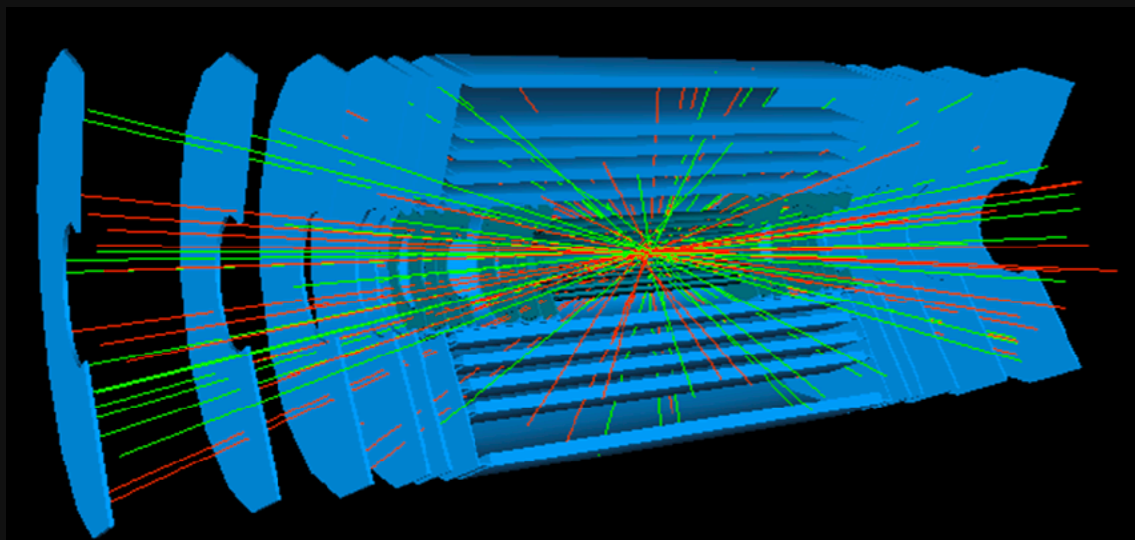


The Future of SUSY

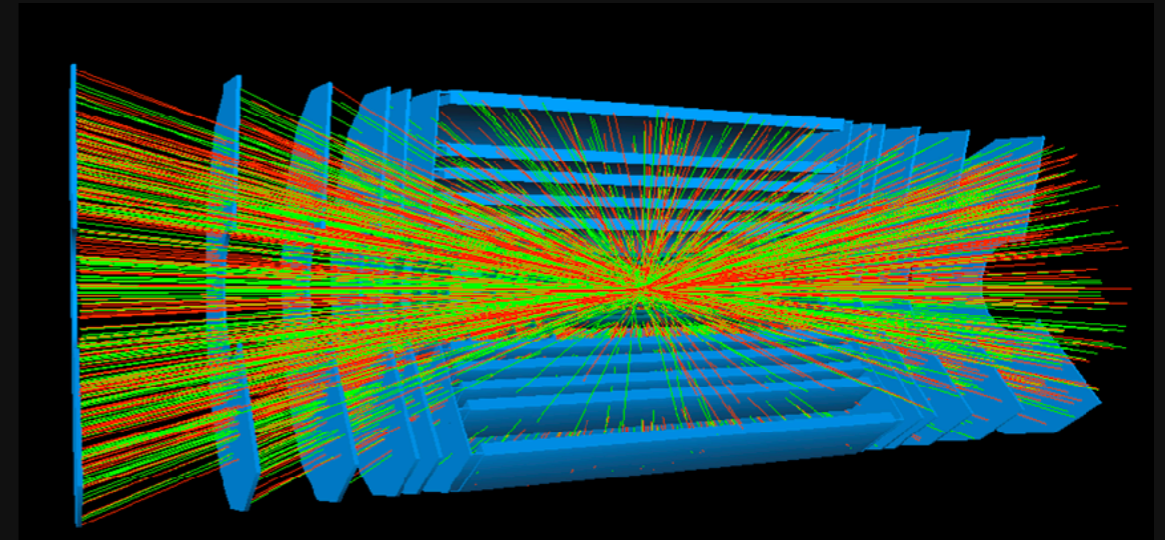


❓ How do we get there?

The High-Luminosity LHC



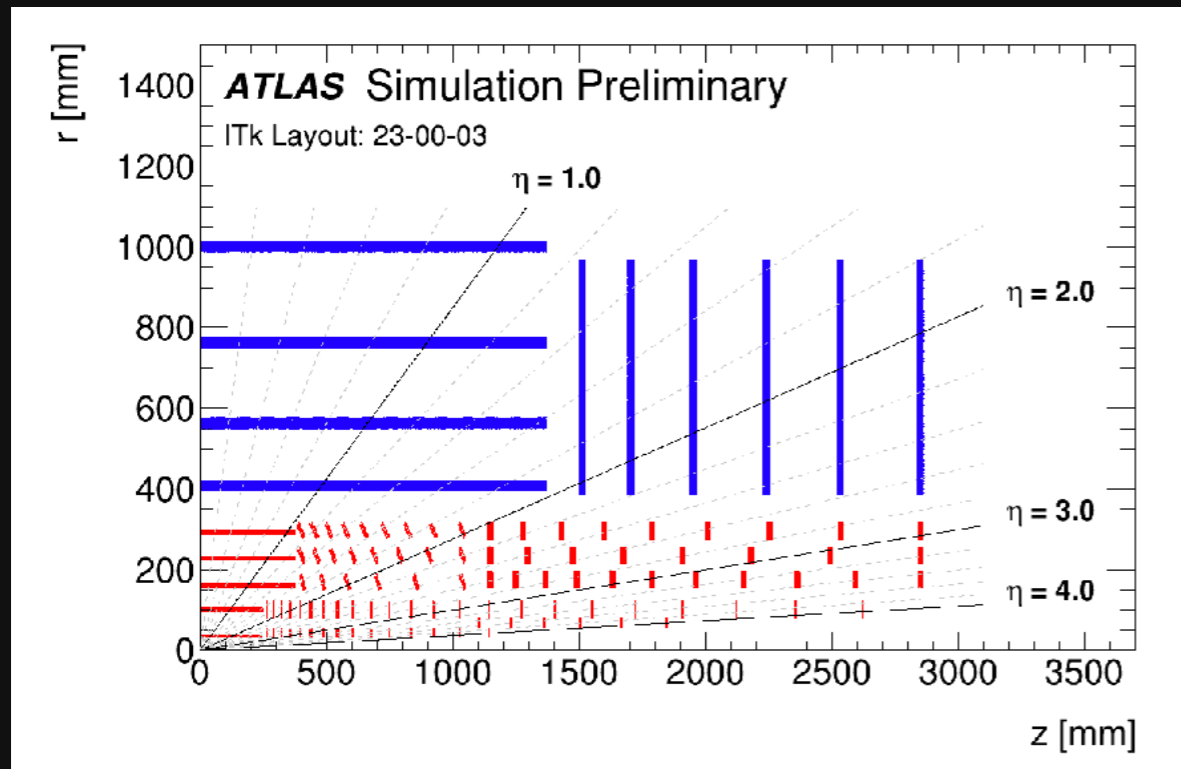
$$\langle \mu \rangle \approx 20$$



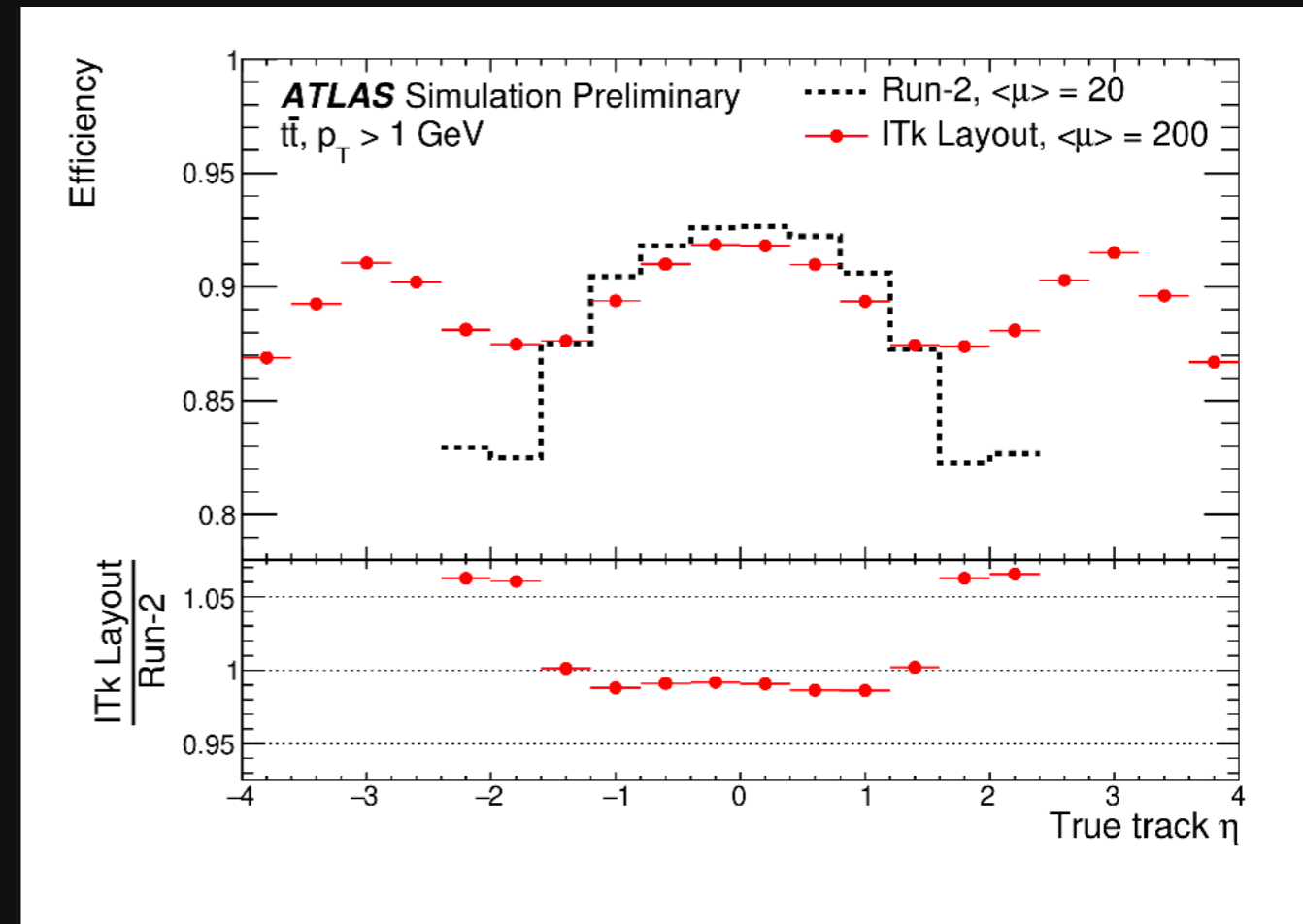
$$\langle \mu \rangle \approx 200$$

- ✦ Starting 2029, expect **200 simultaneous collisions**
- ✦ Requires high-throughput hardware, an **efficient trigger system**, and **radiation-tolerant tracking instrumentation** to survive

Inner Tracker (ITk)



schematic of layout



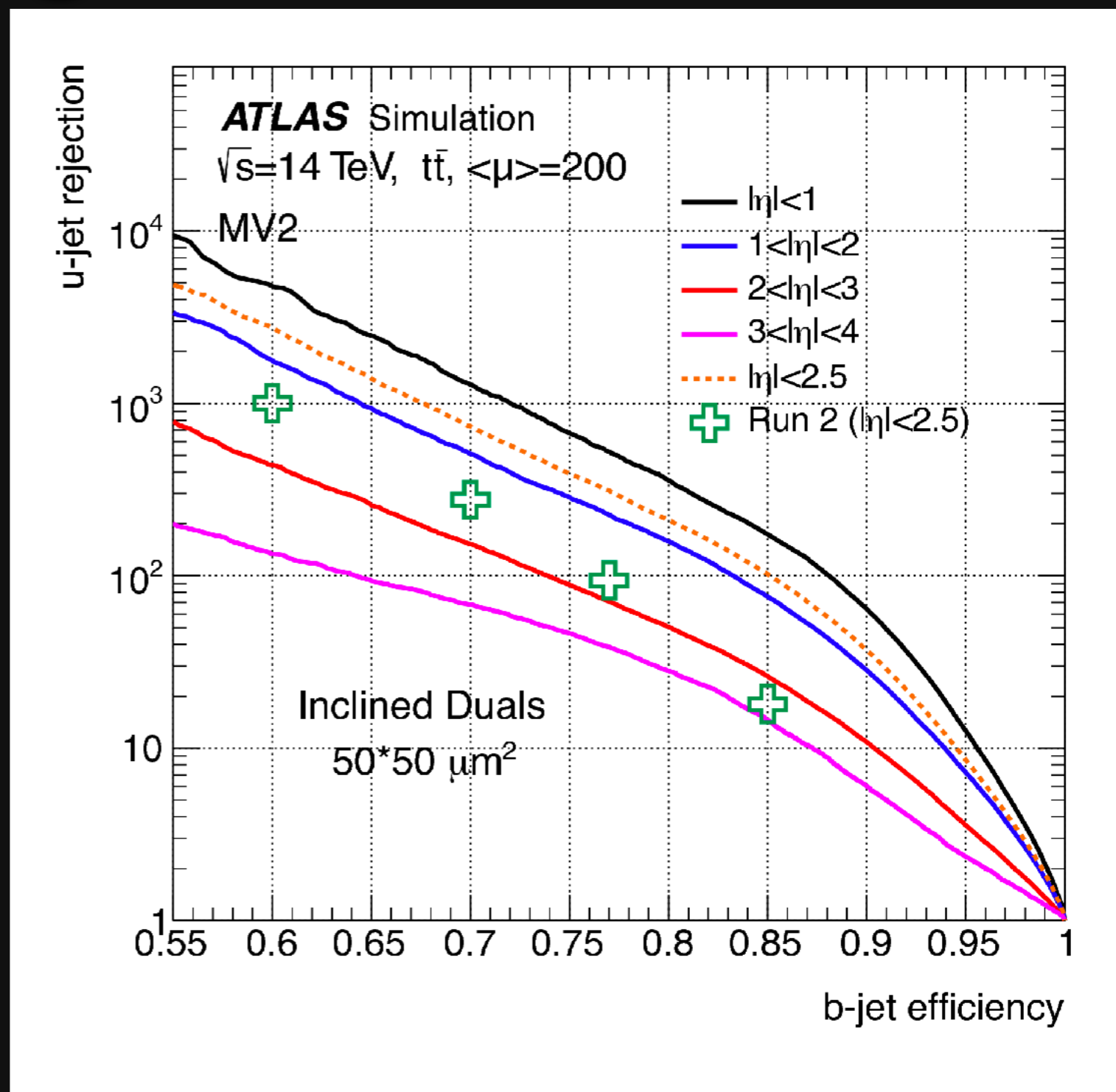
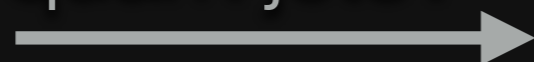
track reconstruction efficiency for $t\bar{t}$ events

- After Run 3 (2026), **completely replace** the inner detector of the ATLAS experiment
- We need **efficient tracking** with **excellent vertex resolution**

Tracking Upgrades

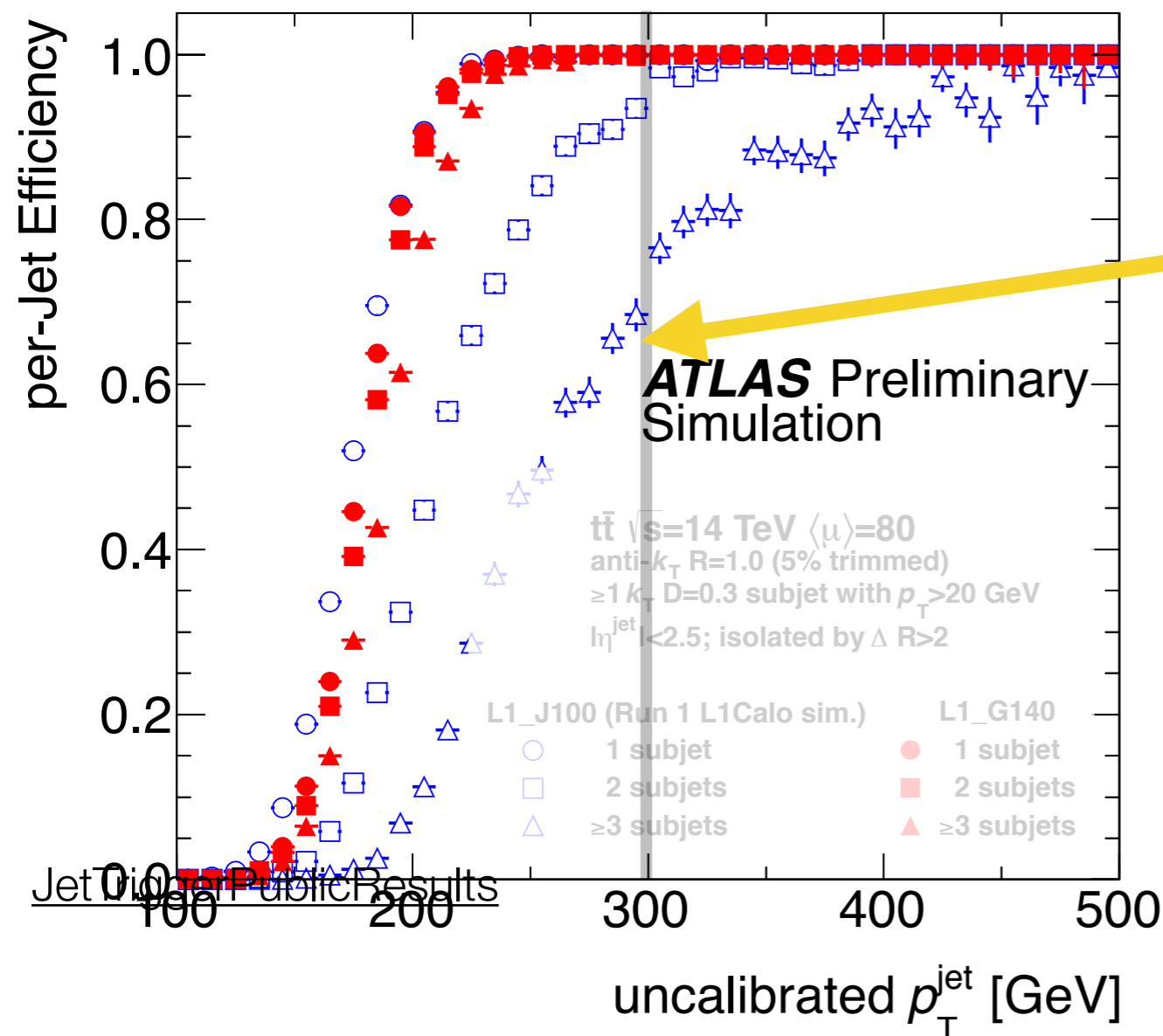
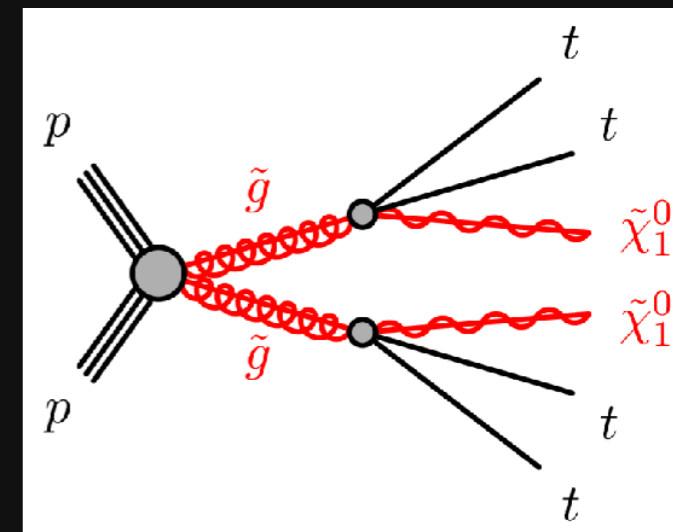
- Requires improved tracking hardware to at least **maintain** these efficiencies, and **survive** the harsh environment

- ❓ “How well can we identify *b*-quark jets?”



Trigger Upgrades

Blue=Current Trigger @ 100 GeV



a top quark at 300 GeV

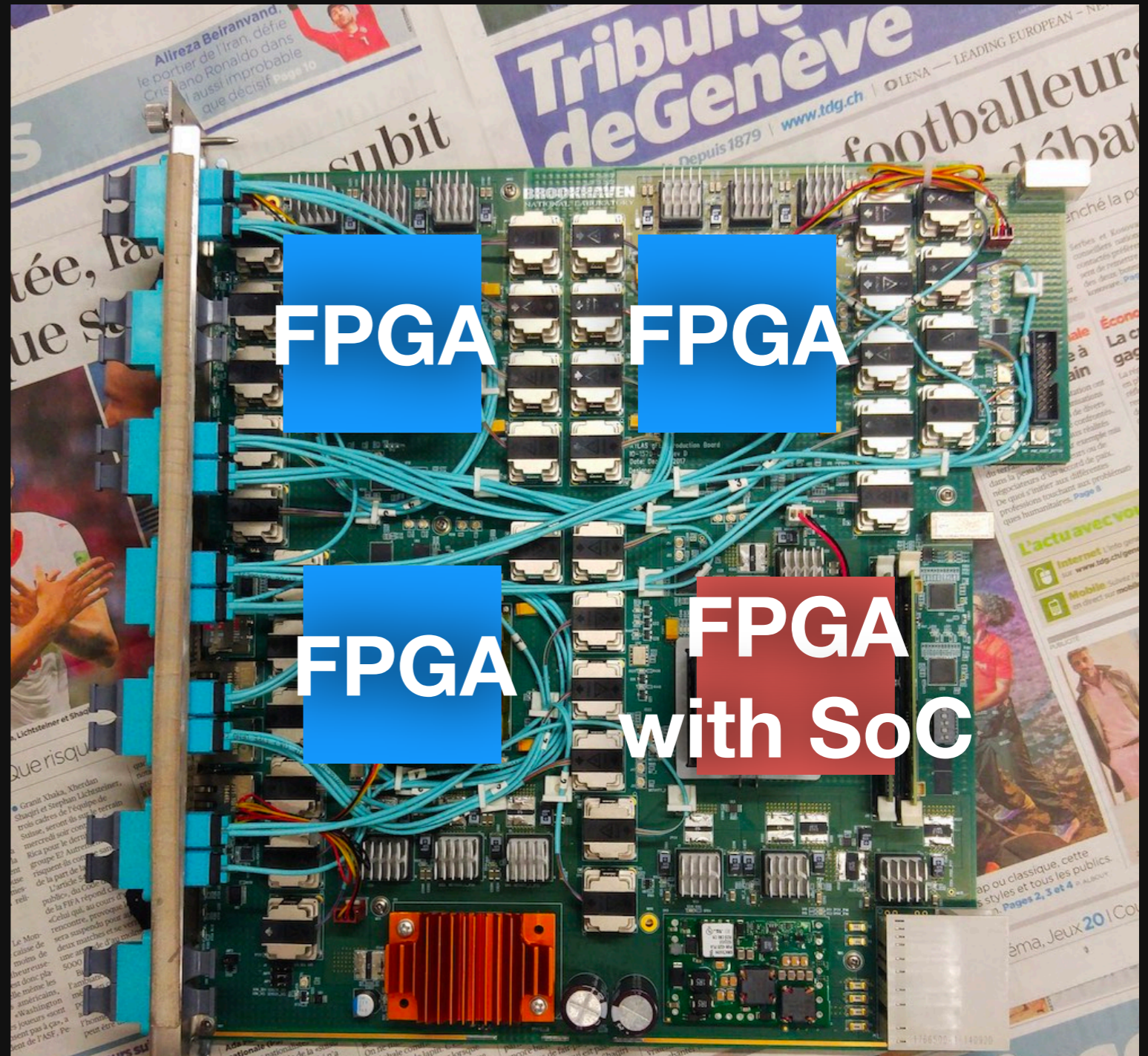
- Many analyses in ATLAS are sensitive to boosted objects with substructure
 - would like a trigger that does not cut them away
- gFEX maintains a flat trigger efficiency here

Red=gFEX Trigger @ 140 GeV

✓ gFEX recovers trigger efficiency for jets with substructure (and MET)! 63

gFEX Board (commissioning!)

- Must process everything in 5 bunch crossings - 125 nanoseconds!
- **3 processor FPGAs receiving calorimeter information and running algorithms**
- **An embedded System-on-Chip for slow control, monitoring, and slower but more advanced physics algorithms**
- Started in 2014 and now installed and commissioning since October 2021!

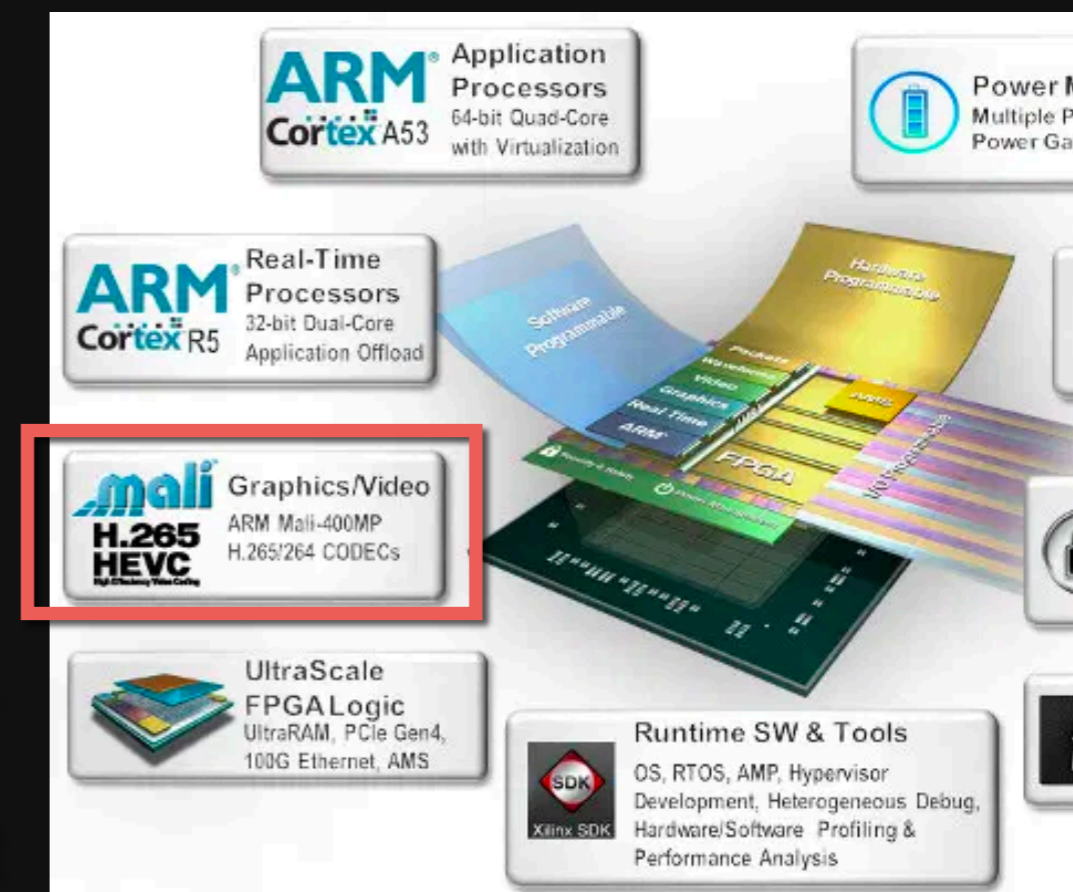


! full calorimeter information on a single board

Trigger Thoughts

- gFEX is using bleeding-edge technology (UltraScale+ MPSoC with ARM Mali-400 MP2 GPU)
- **What can we do with GPUs** in the trigger/detector?
 - Exploit tracking algorithms? Vertexing?
 - More global object reconstruction?
 - Parallelize algorithms?
- Could **information from the ITk readout chips** be used in the trigger?
 - Requires low latency (bandwidth-limited?)
 - Possibly reduce front-end electronics in the detector

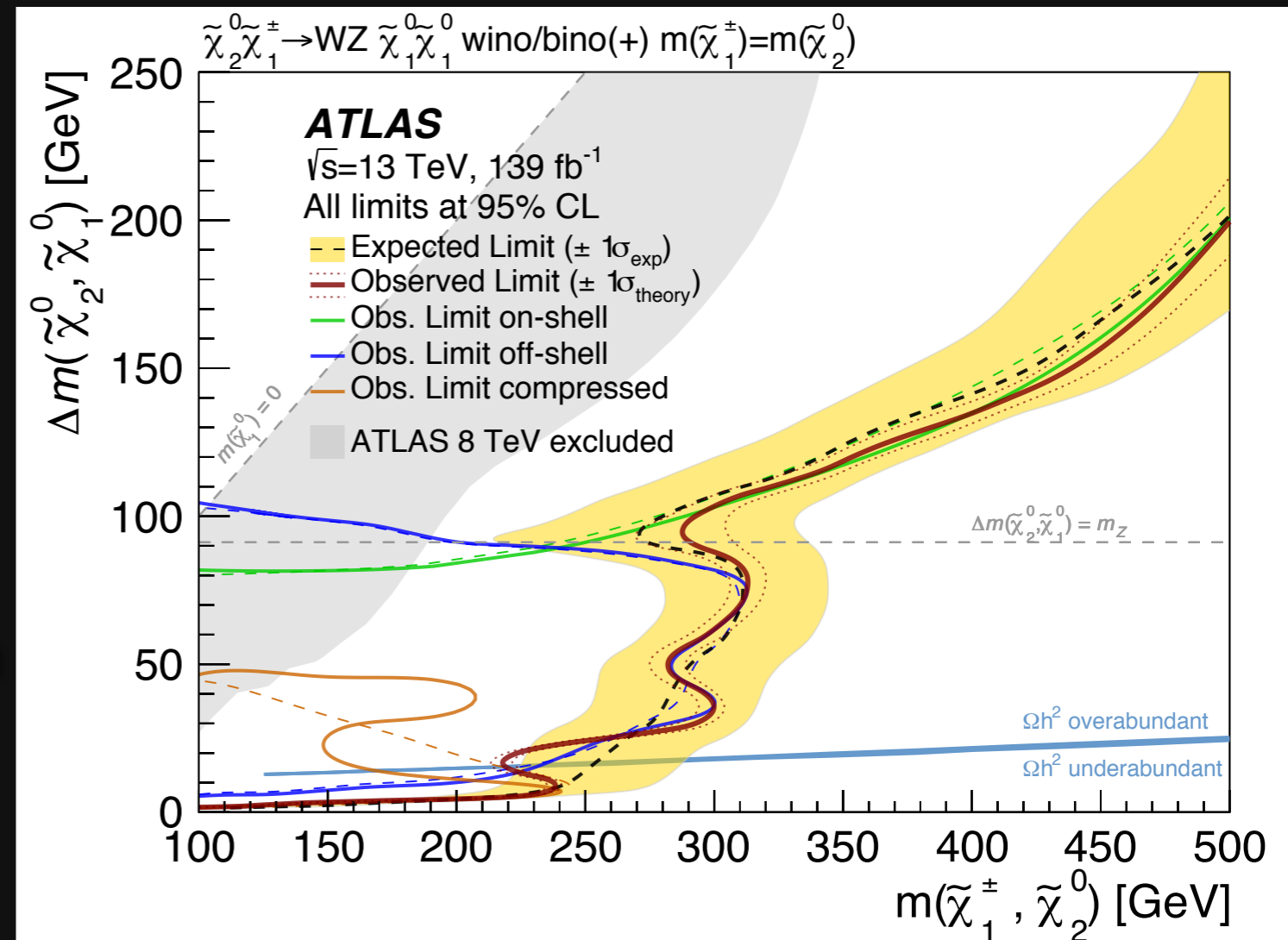
Research in this area could be exploited not only for HL-LHC, but other experiments such as μ -collider, EIC, FCC, DUNE, etc...



Conclusion



- What will the future bring?
 - Run 3 is now!
 - Run 4+ in 2029
 - Other colliders?
- A rich physics and hardware program
 - Lots of opportunities for students to work on instrumentation, data analysis, and software development
- Continue to improve accessibility through preservation/reproducibility as well as teaching/promoting inclusive practices
- Identify uncovered areas and find new rocks to look under for SUSY



Conclusion



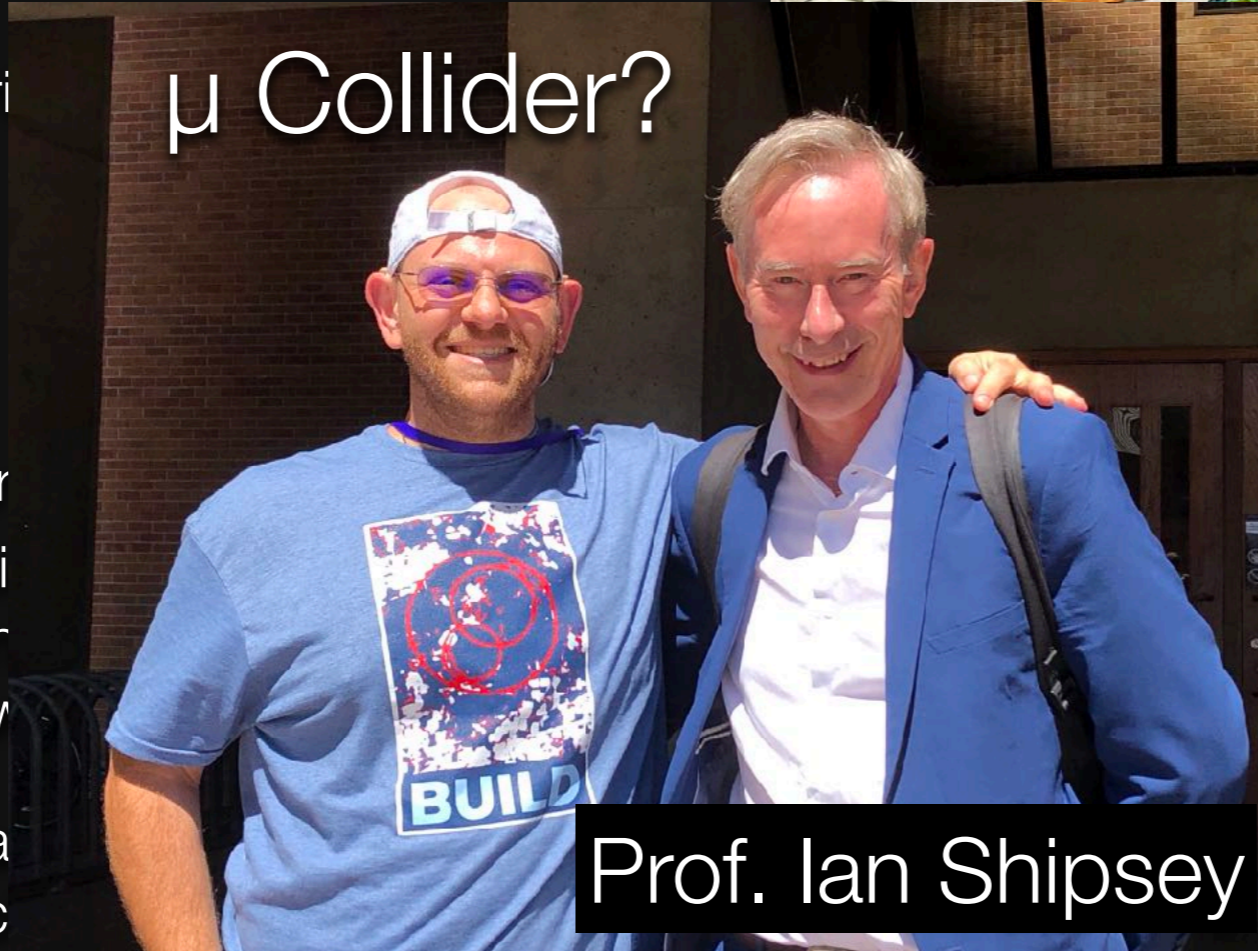
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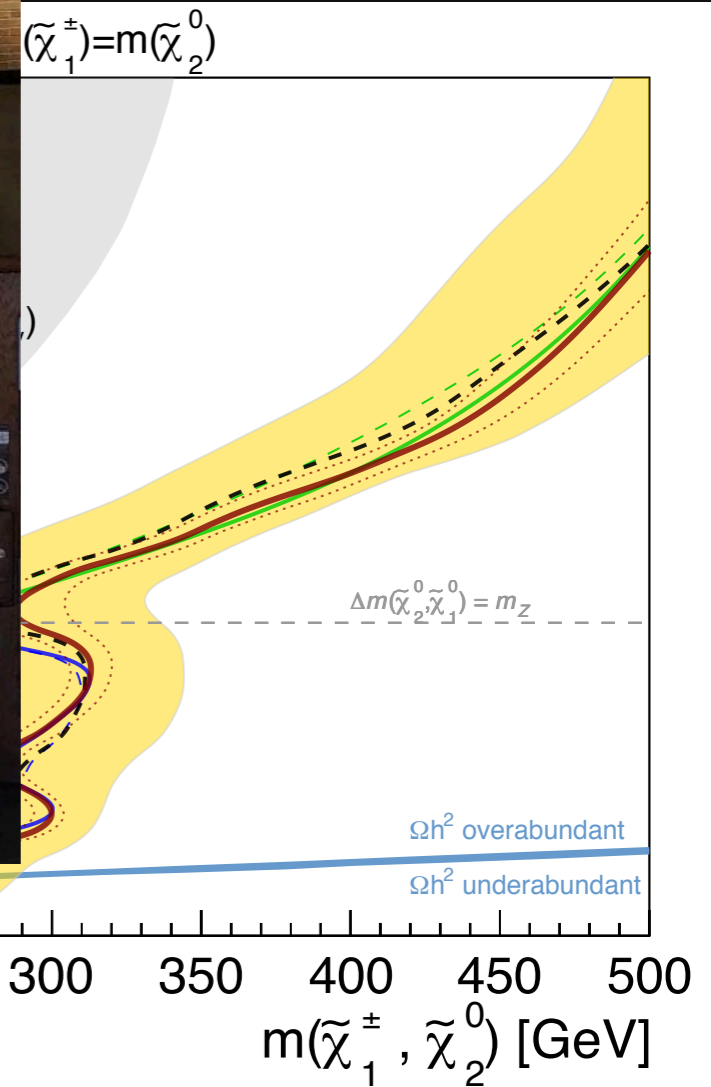
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- Identify uncovered areas and find new rocks to look under for SUSY

μ Collider?



Prof. Ian Shipsey





Seattle Snowmass — July 25th, 2022



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