# DARK MATTER

## Problem Sheet 5: Exam

This problem sheet brings together everything that you have learnt on the course.

#### 1. Classical evidence

- (a) Give three pieces of evidence for *dark matter* in the Universe.
- (b) A spherical galaxy has enclosed mass:

$$M(< r) = M_* + M_0 r^\alpha \tag{1}$$

where  $M_*$  is the total visible mass in stars;  $M_0$  and  $\alpha \ge 0$  are constants; and r is the radius from the centre of the galaxy.

(c) Balance the centripetal force and gravity to show that the velocity of a particle of mass m moving on a circular orbit in the galaxy at radius r is given by:

$$v(r) = \sqrt{\frac{G\left(M_* + M_0 r^\alpha\right)}{r}} \tag{2}$$

- (d) At large distance  $r \to \infty$  from the centre of the galaxy, the rotational velocity is observed to tend towards a constant value  $v(r) = v_0$ . Derive the value of  $\alpha$  in this case and show that  $v_0^2 = GM_0$ .
- (e) Show that for the above value of  $\alpha$ , the cumulative mass M(< r) is given by:

$$M(< r) = M_* + \frac{v_0^2}{G}r$$
(3)

(f) What do you think the above equation for the cumulative mass is telling us? Explain your reasoning.

#### 2. Cosmological Probes

The Friedmann equation is given by:

$$\left(\frac{\dot{a}}{a}\right)^2 = H_0^2 \left[\Omega_\Lambda + \Omega_m a^{-3} + \Omega_r a^{-4} - (\Omega_0 - 1)a^{-2}\right]$$
(4)

where  $\Omega_0 = \Omega_{\Lambda} + \Omega_m + \Omega_r$  and *a* is the dimensionless scale factor.

(a) Hansen et. al (2004) estimate an age of  $12.7 \pm 0.7$  Gyrs for the white dwarf stars in the star cluster M4<sup>1</sup>. Assuming  $\Omega_{\Lambda} = 0$  and  $H_0 = 70 \text{ km/s/Mpc}$ , can you construct a Universe of this age? Is this a reasonable model of the Universe? [Hint: you will need to include the curvature term; see the notes for the solution in this case.]

### 3. Baryonic effects

Suppose I have a galaxy of mass  $M_i$ . We will assume that it is a point mass with dark matter orbiting around it on circular orbits of radius  $r_i$ .

- (a) Suppose I slowly add a mass of  $M_b$  in baryons to the above galaxy. Derive the resulting radii of the dark matter particle orbits  $r_t$ .
- (b) Imagine now that I instantaneously remove the mass  $M_b$  (due to e.g. a galactic wind). Prove that the galaxy becomes unbound if  $M_b = M_i$ .

<sup>&</sup>lt;sup>1</sup>https://arxiv.org/abs/astro-ph/0205087.