# Direct Detection and Collider Searches of Dark Matter Lecture 5

Graciela Gelmini - UCLA

#### **Content of Lecture 5**

- Introduction to search strategies at the LHC
- Complete and simplified DM models, their searches at the LHC and complementarity with other DM searches

Subject is very vast, so idiosyncratic choice of subjects + citations disclaimer

# The Large Hadron Collider (LHC)

- The most powerful particle accelerator in the world, 27 km around, 100 m below ground
- 1600 superconducting magnets operating at 1.9 K

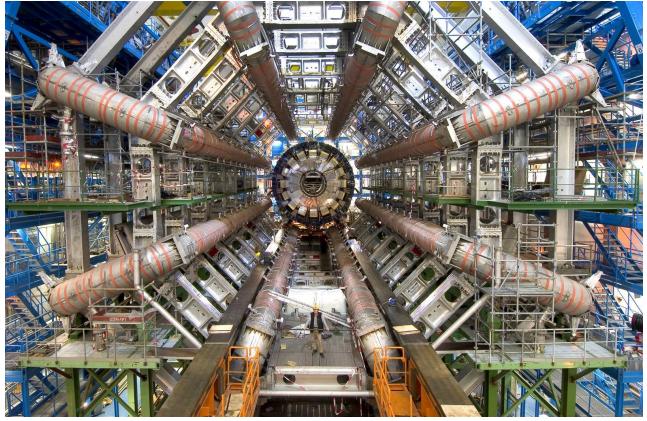


# DM search strategies at the LHC

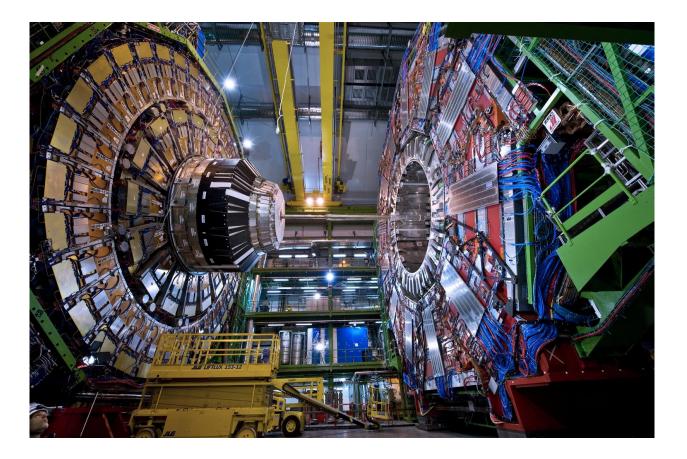
Dark Matter School, Lund, Sept. 26-30, 2016

.

### **ATLAS**



# CMS

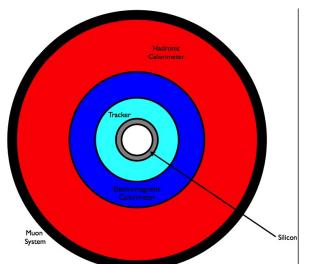


#### LHC multipurpose experiments: ATLAS and CMS



#### Are very large and very complicated detectors!!

#### Theorist version of an LHC detector



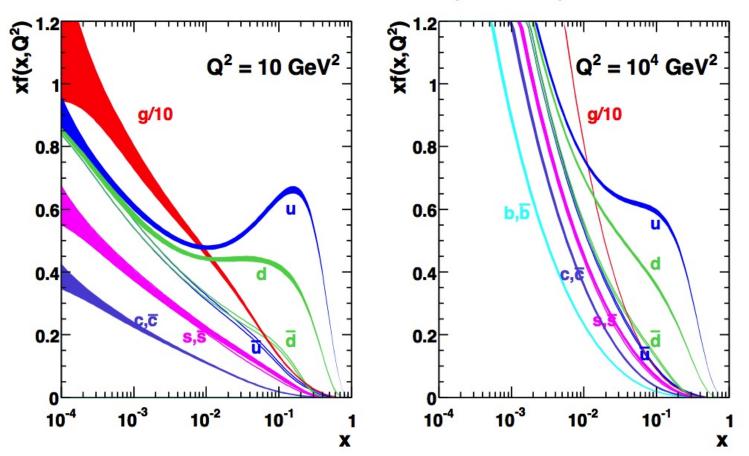
- Inner layer of silicon detectors: displaced vertices
- EM tracker: path of charged particles
- Calorimeters: stop e,  $\gamma$  and hadrons
- Outer radius of  $\mu$  detectors: muon momenta
- Magnetic fields bend charged particle paths: measure momentum

# There is no DM detector! DM signal is missing energy and momentum, actually MET $(p_T)$ . But so is for neutrinos!

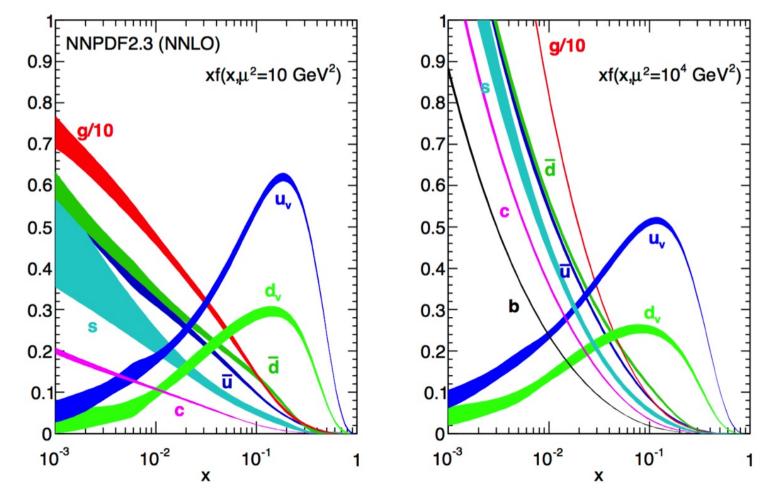
In hadron colliders, the initial momentum along the beam axis  $xp_{had}$  of the colliding partons is not known so the amount of TOTAL missing energy/momentum cannot be determined. However, the initial par tonic momentum transverse to the beam axis  $p_T = 0$ , so any net momentum in the transverse plane indicates Missing Transverse Energy (MET) really  $p_T$ (advantage of lepton colliders: can measure the total missing energy/momentum)

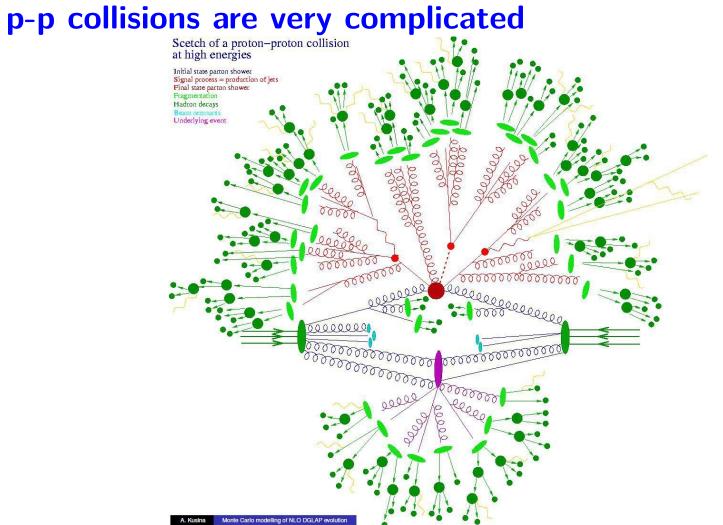
#### **Protons are bags of partons**

#### MSTW 2008 NLO PDFs (68% C.L.)









#### **Outgoing partons** Scetch of a proton-proton collision at high energies ragment and hadronize Initial state parton shower Signal process = production of jet Final state parton shower into jets. agmentat Hadron decays Underlying event Leegeeeeee Proton remnants leave 0000000 behind hadronic debris. 00000000000000 Protons come in as color singlets.

#### p-p collisions are very complicated Fig. from T. Tait

### DM searches at the LHC

#### Main signature:

DM particles escape detection at colliders, thus they are characterized by missing transverse energy (missing  $E_T$ , MET) in collider events.

#### Caveats:

- The DM particles may be too heavy to be produced (above a few TeV).

- A signal produced by a particle escaping the detectors with lifetime  $\simeq 100$  ns cannot be distinguished from one with lifetime  $> 10^{17}$  s as required for DM particles.

- Hadron colliders are relatively insensitive to DM that interacts only with leptons.
- A DM signal may be hidden by backgrounds.

#### Main backgrounds for DM MET search

• "QCD background"

Measuring MET is difficult because requires measuring accurately EVERYTHING VISIBLE. Miss-measurement of jet energies is a source of fake missing momentum.

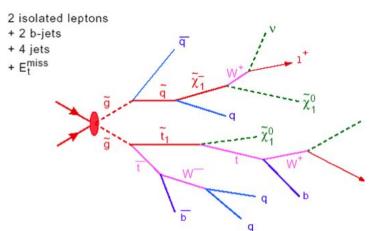
- Neutrinos are a background if they cannot be identified
- Z  $\rightarrow \nu \nu$  20% of the time look like DM MET.

- W  $\rightarrow v\ell$ , if the charged lepton  $\ell$  is missed, v cannot be identified and looks like DM MET.

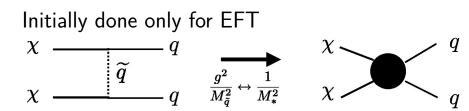
- Same for  $\tau$  decays, also produce  $\nu$ 's and  $\ell$ 's

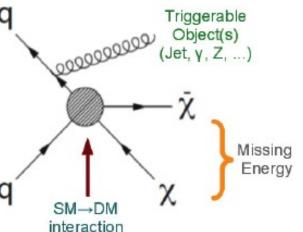
#### Searches at the LHC

Either in complete theories
 DM through known decay chain
 (specific UV complete models e.g. SUSY, or simplified topologies)

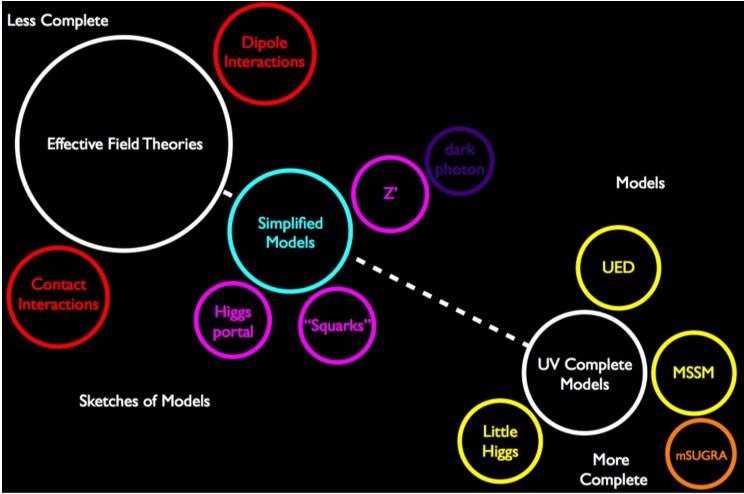


Or direct DM production plus a visible particle either in effective field theories (EFT) or simplified DM models photon or gluon ("monophoton" or "monojet" signal) or mono-W's (leptons), mono-Z's (dileptons), or mono-Higgses.





#### Spectrum of DM Theory Space Fig. from T. Tait



# Complete and simplified DM models, searches at the LHC and complementarity with other DM searches

Dark Matter School, Lund, Sept. 26-30, 2016

.

Graciela Gelmini-UCLA SUSY **Effective Field Theory**  $ar{q}$ q 1000000000  $\tilde{\chi}_{2}^{0}$ ã ĝ  $\bar{\chi}$  $\widetilde{\chi}_1^0$ ĩ W<sup>+</sup> 1<sup>+</sup> W χ qq b a

#### UV complete models - E.g. Supersymmetry (SUSY) Most studied model

- Symmetry between bosons and fermions.
- Models are completely calculable
- Hierarchy: maintains EW scale  $\ll$  GUT scale
- Requires two Higgs doublets minimum.
- Every known particle has supersymmetric partner(s)

Fermions:	Bosons:
SM fermions: $\ell$ , $q$	sfermions: $\widetilde{\mathscr{C}}$ , $\widetilde{q}$
gauginos: $\widetilde{B}, \widetilde{W}, \widetilde{g}$	SM gauge bosons: <i>B</i> , <i>W</i> , <i>g</i>
Gravitinos: $\widetilde{G}$	graviton
higssinos: $\widetilde{H}$	Higgs bosons

• *R*-parity=  $(-1)^{3B+L+2S}$  is  $P_{SM} = +1$ ,  $P_{SUSY} = -1$  distinguishes SM particles and SUSY partners

#### Many versions of SUSY- Many parameters

## **MSSM**

- Minimum number of particles (SUSY partners+ two Higgs doublets)
- Number of parameters: those of the SM + 106!!!
- Parameter reduction:
  - pMSSM: simplified weak-scale MSSM: SM + 19 p.
  - CMSSM: constrained MSSM: SM+5 parameters
  - $(m_0, A_0, m_{1/2}, \tan\beta, \mu)$
  - mSUGRA: minimal supergravity: SM+5 parameters  $(m_0, A_0, m_{1/2}, \tan\beta)$ , sign of  $\mu$ )

#### NMSSM

• Non Minimum number of particles (extra singlet Higgs, etc)

#### Many versions of SUSY- Many parameters

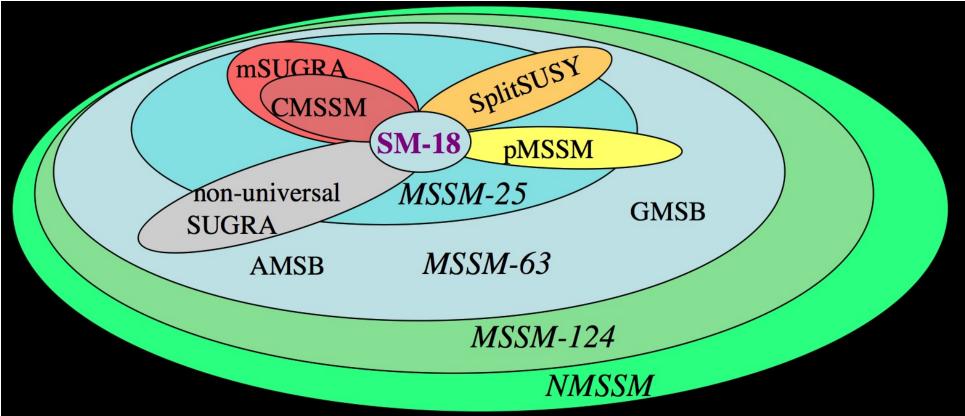


Fig. from P. Gondolo

### **SUSY models**

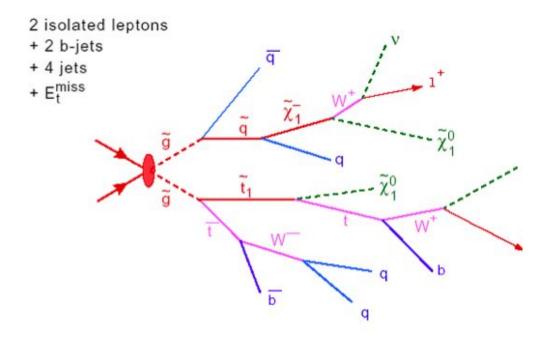
- If R-parity is conserved, the Lightest Supersymmetric Partner (LSP) is stable, thus a good WIMP dark matter candidate (if neutral and colorless):
   ν sneutrino, G Gravitino (partner of graviton), a axino (partner of the axion) or X<sup>0</sup> neutralino (gaugino/ higgssino, partner of neutral gauge boson/Higgs boson)
- In the MSSM the usual LSP is the lightest neutralino. In the basis  $\widetilde{B}$ ,  $\widetilde{W_3}$ ,  $\widetilde{H_1^0}$ ,  $\widetilde{H_2^0}$  the mass matrix is  $\Gamma$   $M_1$   $\Omega$   $\Pi$

15	$\int M_1$	0	$-M_Z c_\beta s_W$	M <sub>Z</sub> s <sub>β</sub> s <sub>W</sub> ⊺	
	0	$M_2$	MzcBcw	$-M_Z \mathbf{s}_{\boldsymbol{\beta}} \mathbf{c}_{\boldsymbol{W}}$	
	$-M_Z c_\beta s_W$	Mzcβcw	0	$-\mu$	
	$M_Z s_\beta s_W$	$-M_Z s_\beta c_W$	$-\mu$	0	

 $\tan \beta = v_2/v_1$ ,  $M_1$ : Bino mass,  $M_2$ : Wino<sub>3</sub> mass,  $\mu$ : mixes  $H_1$   $H_2$ 

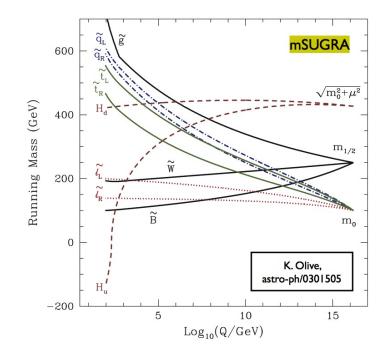
- If one stage unification of fundamental forces:  $M_2 = 2M_1$ , if  $M_1 < |\mu|$ , LSP=  $\widetilde{B}$  typical **cMSSM**, if  $M_1 \simeq |\mu|$ , LSP = mixed  $\widetilde{B}$ - $\widetilde{H}$  OK
- If R-parity is conserved, an unstable SUSY particle decays into another SUSY particle and SUSY particles are only produced in pairs from SM particles.

### LHC typical SUSY decay chain



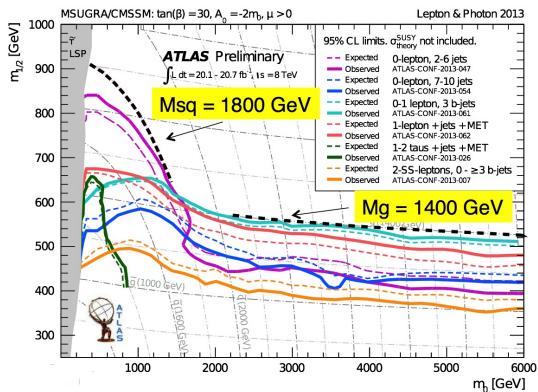
- Typical topology:
- missing energy
- multiple hadronic jets
- number of leptons

**mSUGRA** Start at the GUT scale with few parameters and use renormalization group equations to get to the electroweak scale.



**pMSSM** (phenomenological MSSM) Has 19 SUSY parameters directly defined at the electroweak scale:  $\mu$ ,  $m_A$ , tan  $\beta$ ,  $A_b$ ,  $A_t$ ,  $A_{\tau}$ ,  $M_1$ ,  $M_2$ ,  $M_3$ ,  $m_{Q1}$ ,  $m_{Q3}$ ,  $m_{u1}$ ,  $m_{d1}$ ,  $m_{u3}$ ,  $m_{d3}$ ,  $m_{L1}$ ,  $m_{L3}$ ,  $m_{e1}$ ,  $m_{e3}$ 

#### mSUGRA/cMSSM were in serious troubles by 2013



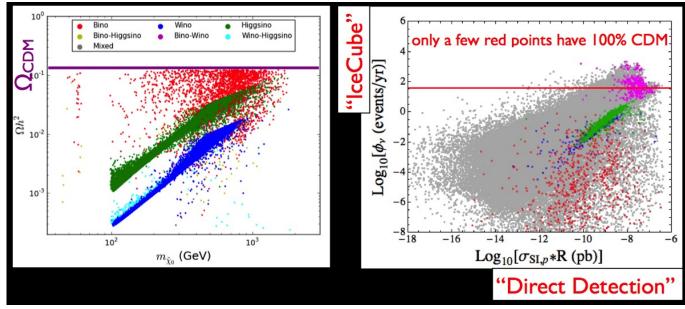
SO:• Use more parameters, pMSSM, or more particles NMSSM or

 $\bullet$  Simplified SUSY model spectrum (SMS) with 100% BR for decay chain considered

## pMSSM (phenomenological MSSM) LHC bounds

Cahill-Rowell et al 1305.6921

Only a few Bino LSP models with coannihilation can still constitute all the DM

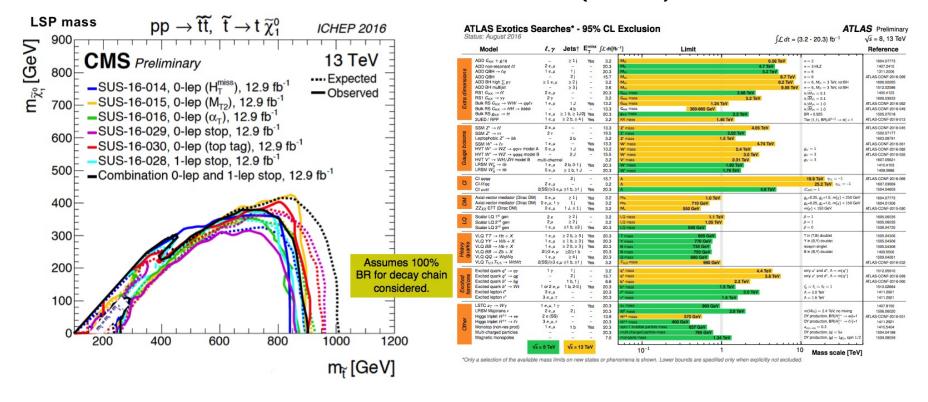


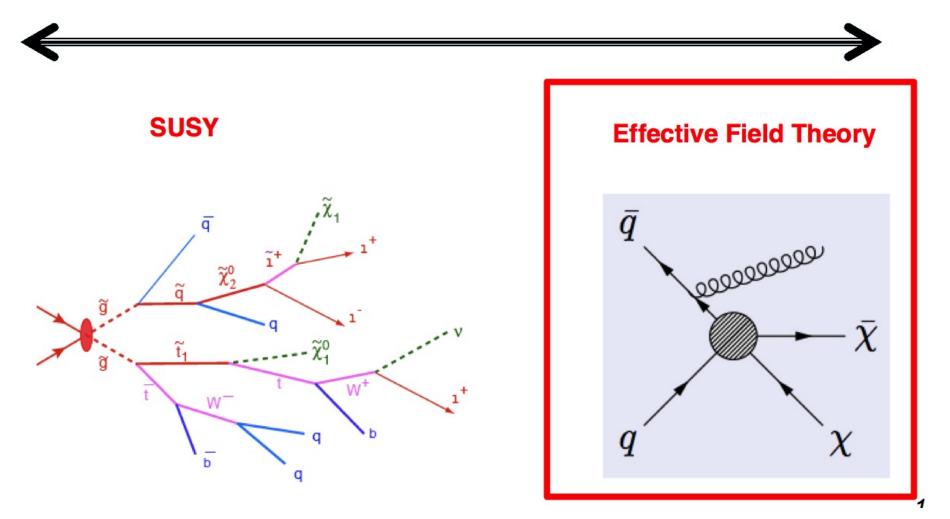
LSP  $\widetilde{B}$ -like (typical in CMSSM) is overdensed - or fine-tunned ( $\sigma_{annih}$  into  $f\bar{f}$  through  $\widetilde{f}$  exchange is helicity suppressed  $\sim m_f$ ) LSP  $\widetilde{H}$ -like and  $\widetilde{W}$ -like are underdensed unless  $m \simeq \text{TeV}$ 's (large  $\sigma_{annih}$  to W<sup>+</sup>W<sup>-</sup>, ZZ or  $f\bar{f}$ )

#### **pMSSM** - **Complementarity of Searches** neutralino in the phenomenological MSSM with 19 free parameters, 50 GeV<m< 4 TeV, 2 10<sup>5</sup> points, each a model- Notice $R = \Omega_{\chi}/\Omega_{DM}$ : fraction of DM density. $R\sigma$ in vert.axis. Cahill-Rowley et al. 1405.6716

Excluded by DD and ID XENON1T Models will be tested by:  $10^{-5}$ Survives DD, ID, and LHC Excluded by ID but not DD Excluded by LHC but not DD or ID Excluded by DD but not ID • direct detection (green) - black  $10^{-7}$ line is reach of LZ indirect detection with (qd)  $10^{-9}$ FermiLAT and future CTA (red) • other indirect detection (blue) 2 escape all other searches and  $10^{-13}$ will be tested only at the LHC (magenta). 10-15 • Gray regions will survive all searches in the near future.  $10^{-17}$  $10^{3}$  $10^{2}$  $m( ilde{\chi}^0_1)$  (GeV)

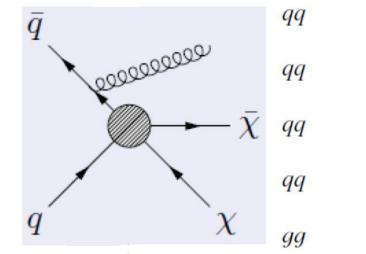
#### SUSY Simplified Models Spectrum (SMS)

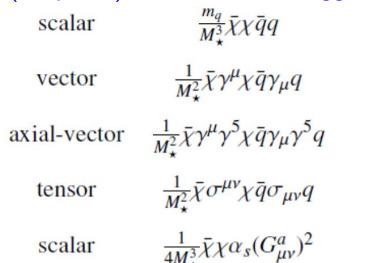




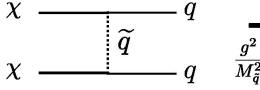
#### At the other extreme of a complete theory is Effective Field Theory. "MONO"-Searches at the LHC

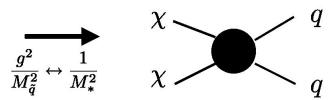
Direct DM production plus a photon or gluon ("monophoton" or "monojet" signal) or mono-W's (leptons), mono-Z's (dileptons), or even mono-Higgses.

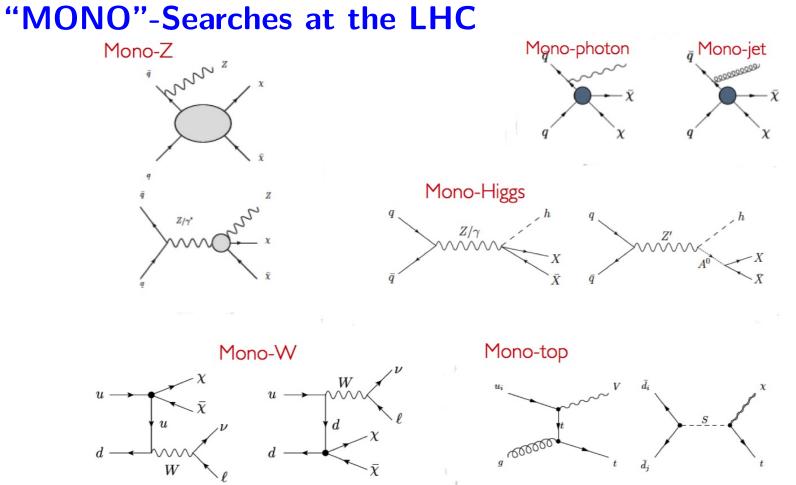




Initially done only for CONTACT INTERACTIONS (EFT couplings)

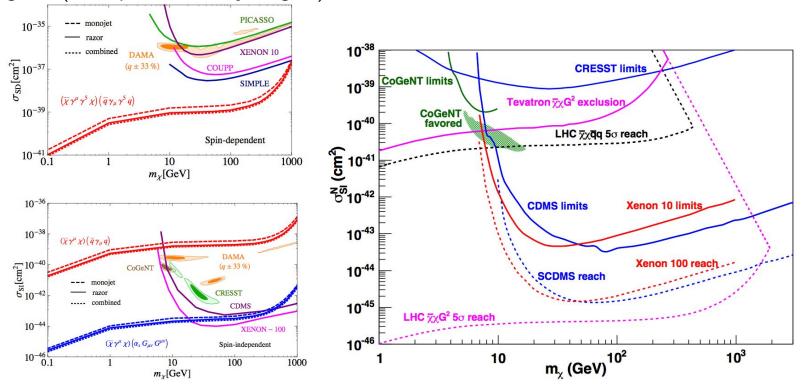






#### EFT "Mono"- searches at the LHC direct production plus a photon or

gluon (monophoton monojet signal) for CONTACT interactions Beltran et al 1002.4137; Fox et al 1203.1662



CAVEAT: in direct DM detection "contact interaction" if mediator M> q > MeV's but at LHC q much larger,M> 100's GeV- Compare with care!

#### "Mono"- searches at the LHC - direct detection limits must be compared with care.

The approach using single effective operators with contact interactions, is limited because of absence of possible interference between different operators, effect of lighter mediators than those necessary to have a contact interaction at the LHC...

A mediator heavy at the partonic LHC energies is also heavy in direct detection BUT THE OPPOSITE IS NOT TRUE. If the mediator is light for the LHC, could itself be produced (and at LEP etc) so limits change.

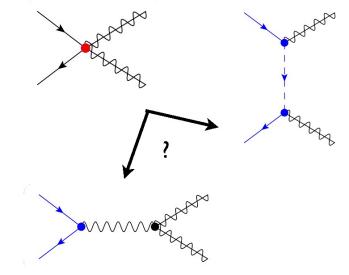
Now trying to use **"simplified models"** and classify classes of mediators in the s-channel and the t-channel, or by the way the DM relic density could occur (e.g. proposed "Benchmarks for DM detection at the LHC" De Simone, Giudice & Strumia 1402.6287)

. Lots of work to do in this direction...

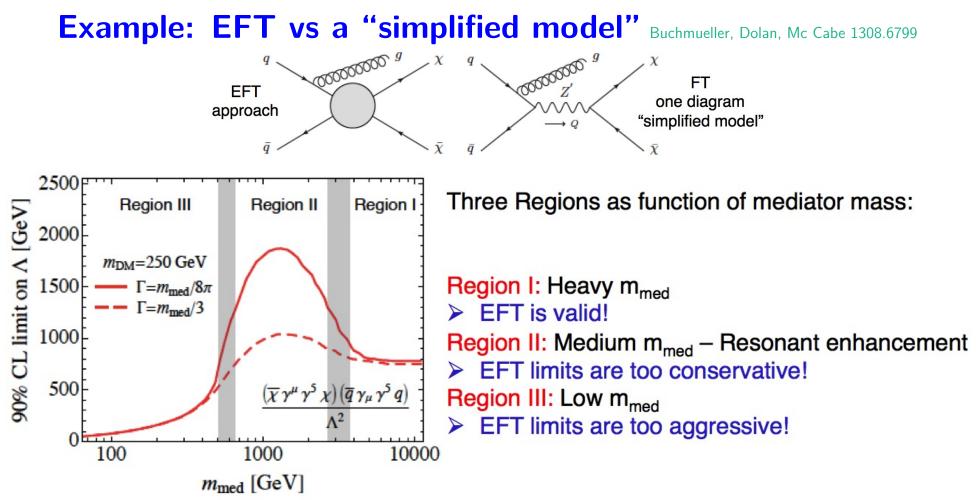
"LHC Dark Matter Working Group" recommendations: 1603.04156

#### EFT vs "simplified model"

• Many different FT couplings have the same EFT limit! Many, many models to study. How to classify them?



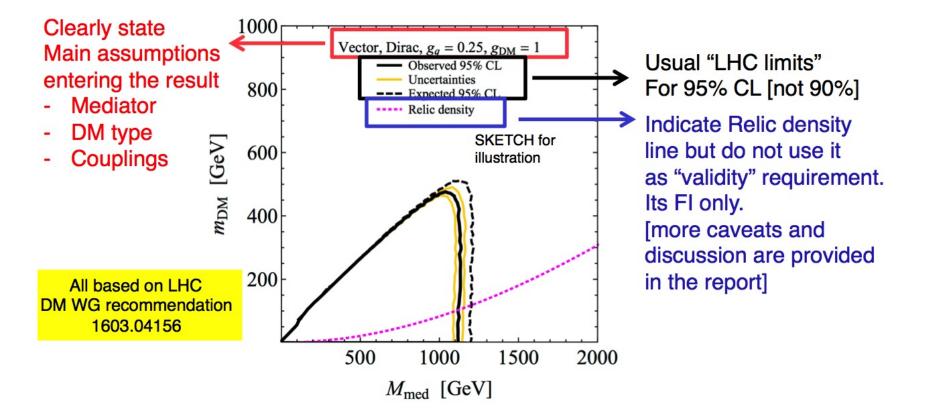
• For light enough mediators, limits on them are important. The mediators themselves can be created at the LHC and other accelerators.



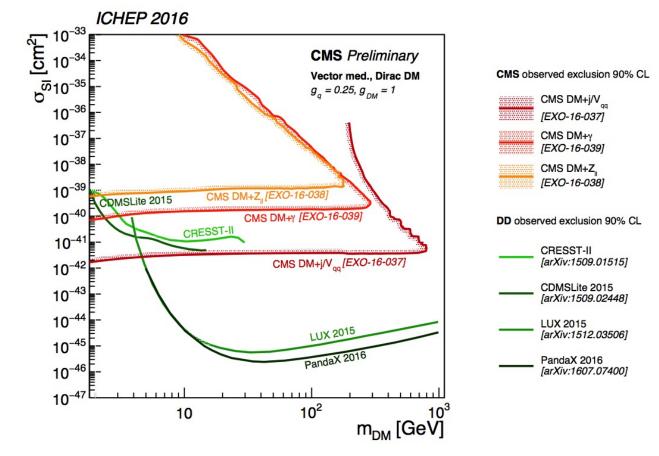
### How to present "simplified models"

"LHC Dark Matter Working Group" recommendations, 1603.04156

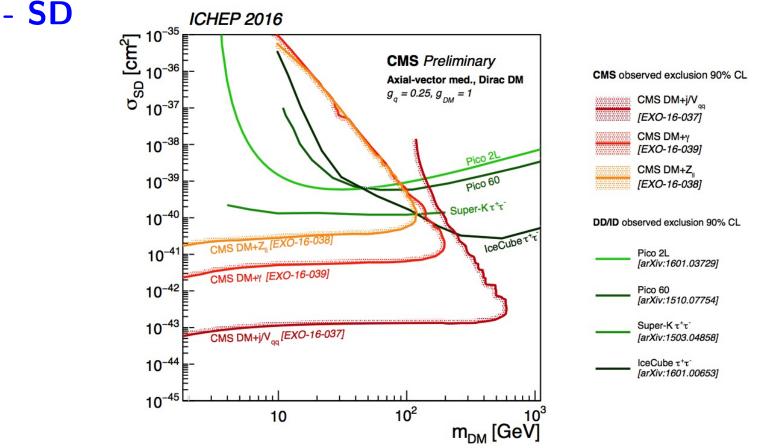
Main result of the interpretation of collider search in simplified model



#### "Mono" searches - Simplified Model - Vector mediator-SI



# "Mono" searches - Simplified Model Axial-Vect. mediator



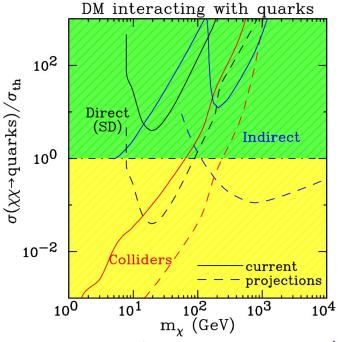
How would these results compared if for this particle  $\Omega/\Omega_{DM} = f = 0.01$ ?

**Example of complementarity of DM searches** Simple contact interaction model considered for the Snowmass 2013 study (Bauer et al. 2013)

$$\mathscr{L} = \frac{1}{M_q^2} \,\bar{\chi} \gamma^{\mu} \gamma_5 \chi \sum_q \bar{q} \gamma_{\mu} \gamma_5 q + \frac{\alpha_S}{M_g^3} \,\bar{\chi} \chi G^{a\mu\nu} G^a_{\mu\nu} + \frac{1}{M_\ell^2} \,\bar{\chi} \gamma^{\mu} \chi \sum_{\ell} \bar{\ell} \gamma_{\mu} \ell \,\,,$$

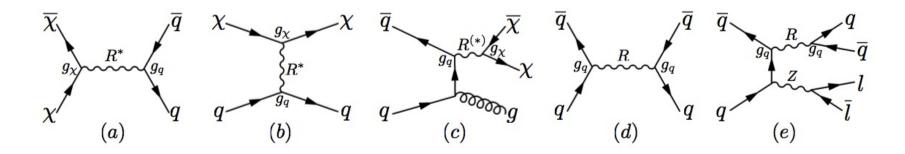
The interactions with quarks mediate SD direct signals (\*), and those with gluons mediate SI direct signals.  $M_q$ ,  $M_g$ , and  $M_\ell$  are chosen so that the relic density is exactly that of the DM though thermal production. (\* You will prove that  $\gamma^{\mu}\gamma_5$  leads to SD scattering).

Notice that: the limits would be very different if the DM candidate accounted for only a fraction  $\Omega/\Omega_{DM} = f < 1$  of the DM (\*)



(\* In exercise 10 you will explore this - assuming the indirect limits come from DM annihilation).

#### Important complementarity of dark sector particles and mediators searches More complete Simplified Model Chala et al. 1503.0591

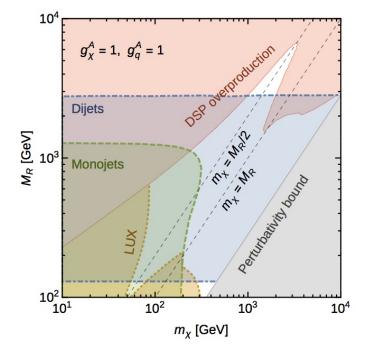


Dark Sector Particle (DSP)  $\chi$  and mediator R in all channels:

- (a) DM annihilation which sets the relic abundance,
- (b) DM scattering in direct detection experiments,
- (c) monojet signatures, in this case due to initial state radiation of a gluon,
- (d) LHC Dijet resonance signatures purely through mediator-quark couplings and
- (e) dijet associated production (at UA2, CDF, CMS, and ATLAS)

#### Example of combined constraints (at 95% CL): everything excluded!

Chala et al. 1503.0591



This shows that for each LHC DM Simplified Models there should be a complete study of the mediator particles in all experiments too.

#### A lot of work remains to do at the LHC!

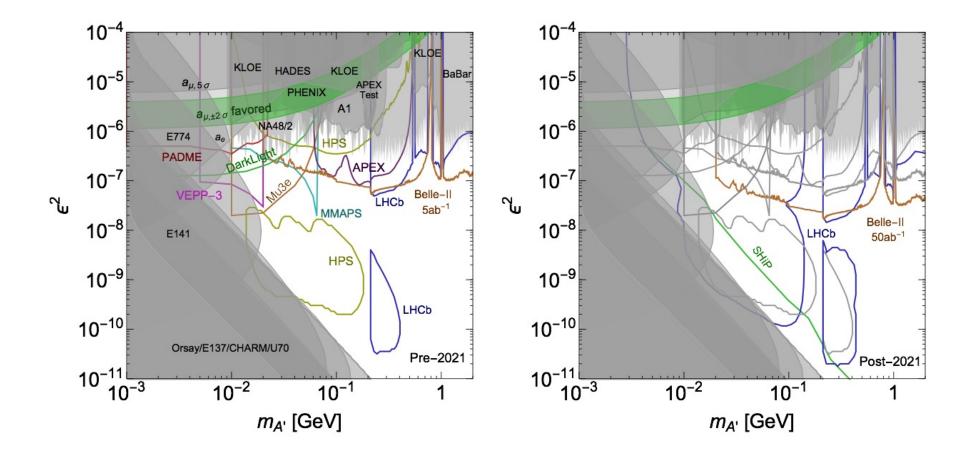
#### **Brief comment on accelerator dark sector searches** Dark Sectors Workshop 1608.08683

Dark sectors include one or more mediator particles coupled to the SM via a portal. The portal relevant for dark sector-SM interactions depends on the mediator spin and parity: it can be a scalar  $\varphi$ , a pseudoscalar a, a fermion N, or a vector A' (dark or hidden photon).

$$\mathcal{L} \supset \begin{cases} -\frac{\epsilon}{2\cos\theta_W} B_{\mu\nu} F'^{\mu\nu} , & \text{vector portal} \\ (\mu\phi + \lambda\phi^2) H^{\dagger}H , & \text{Higgs portal} \\ y_n LHN , & \text{neutrino portal} \\ \frac{a}{f_a} F_{\mu\nu} \widetilde{F}^{\mu\nu} , & \text{axion portal.} \end{cases}$$

Many proposals to detect dark sector particles in all types of accelerators (fixed target, lepton colliders, B factories).

e.g. Dark Photons A' search at LHCb via rare heavy quark decay modes producing a A' which could either produce MET or a visible decay,  $A' \rightarrow \mu^+\mu^-$  with or without a displaced vertex.



Experiment	Lab	Production	Detection	Vertex	$M_{\rm ass}(M_{\rm e}V)$	$M_{\rm ass}R_{\rm es.}(M_{\rm eV})$	Beam	Ebeam~(GeV)	Ibeam or Lumi	Machine	1st Run	Next Run
APEX	JLab	e-brem	$\ell^+\ell^-$	no	65 - 600	0.5%	<i>e</i> <sup>-</sup>	1.1-4.5	150 µA	CEBAF(A)	2010	2018
A1	Mainz	e-brem	$e^+e^-$	no	40 - 300	?	<i>e</i> <sup>-</sup>	0.2-0.9	140 µA	MAMI	2011	-
HPS	JLab	e-brem	$e^+e^-$	yes	20 - 200	1–2	e <sup>-</sup>	1–6	50–500 nA	CEBAF(B)	2015	2018
DarkLight	JLab	e-brem	$e^+e^-$	no	< 80	?	$e^{-}$	0.1	10 mA	LERF	2016	2018
MAGIX	Mainz	e-brem	$e^+e^-$	no	10 - 60	?	<u>e</u> -	0.155	1 mA	MESA	<b>2020</b>	-
NA64	CERN	e-brem	$e^+e^-$	no	1 <b>- 5</b> 0	?	<i>e</i> <sup>-</sup>	100	$2 \times 10^{11} { m ~EOT/yr}$	SPS	2017	2022
Super-HPS	SLAC	e-brem	vis	yes	< 500	?	e <sup>-</sup>	4 - 8	$1 \ \mu A$	DASEL	?	?
(TBD)	Cornell	e-brem	$e^+e^-$	?	< 100	?	<i>e</i> <sup>-</sup>	0.1-0.3	100 mA	CBETA	?	?
VEPP3	Budker	annih	invis	no	5 - 22	1	<i>e</i> +	0.500	$10^{33}{ m cm}^{-2}{ m s}^{-1}$	VEPP3	2019	?
PADME	Frascati	annih	invis	no	1 - 24	2 - 5	<i>e</i> <sup>+</sup>	0.550	$\leq 10^{14}  e^+ \mathrm{OT/y}$	Linac	2018	?
MMAPS	Cornell	annih	invis	no	20 - 78	1 - 6	$e^+$	6.0	$10^{34}{\rm cm}^{-2}{\rm s}^{-1}$	Synchr	?	?
KLOE 2	Frascati	several	vis/invis	no	< 1.1  GeV	1.5	$e^+e^-$	0.51	$2\times 10^{32}{\rm cm}^{-2}{\rm s}^{-1}$	$DA\phi NE$	2014	-
Belle II	KEK	several	vis/invis	no	$\lesssim 10{\rm GeV}$	1 - 5	$e^+e^-$	$4 \times 7$	$1 \sim 10 \text{ ab}^{-1}/\text{y}$	Super-KEKB	2018	-
SeaQuest	FNAL	several	$\mu^+\mu^-$	yes	$\lesssim 10{ m GeV}$	3 – <mark>6</mark> %	р	120	10 <sup>18</sup> POT/y	MI	2017	2020
SHIP	CERN	several	vis	yes	$\lesssim 10{ m GeV}$	1 - 2	р	400	$2 \times 10^{20}$ POT/5y	SPS	2026	-
LHCb	CERN	several	$\ell^+\ell^-$	yes	$\lesssim 40{ m GeV}$	~ 4	pp	6500	$\sim 10{\rm fb}^{-1}/{ m y}$	LHC	2010	2015

TABLE I: Summary of dark photon experiments.